

# Exploring dark sectors with low-energy experiments

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Caltech

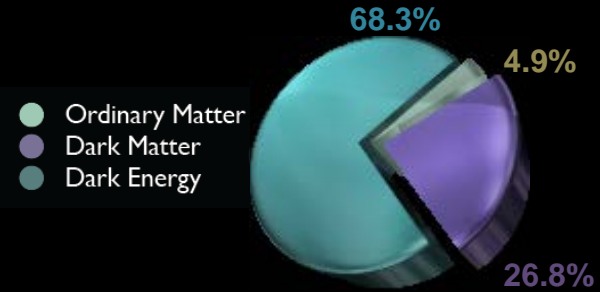
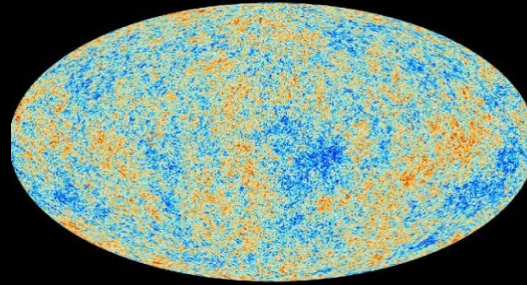
UCI Joint Particle Seminar  
UCI - February 2015

# Dark matter

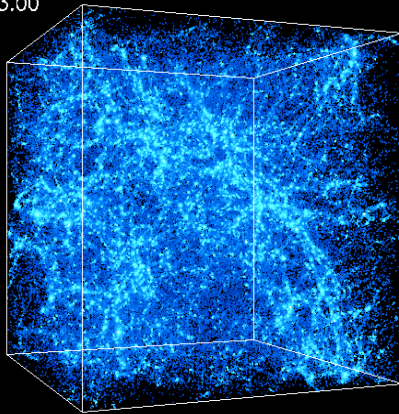
Gravitational lensing



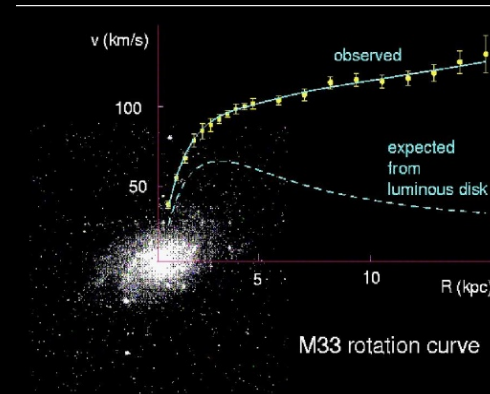
CMB



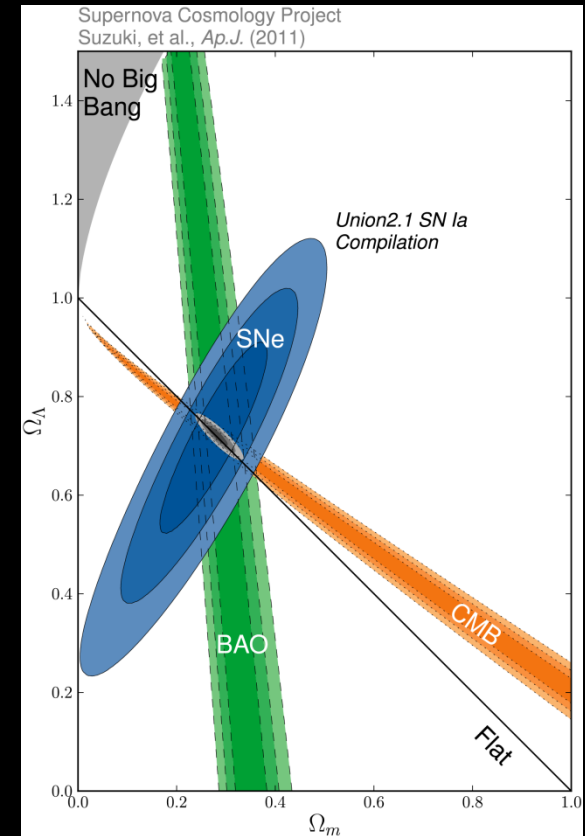
$Z = 3.00$



Structure

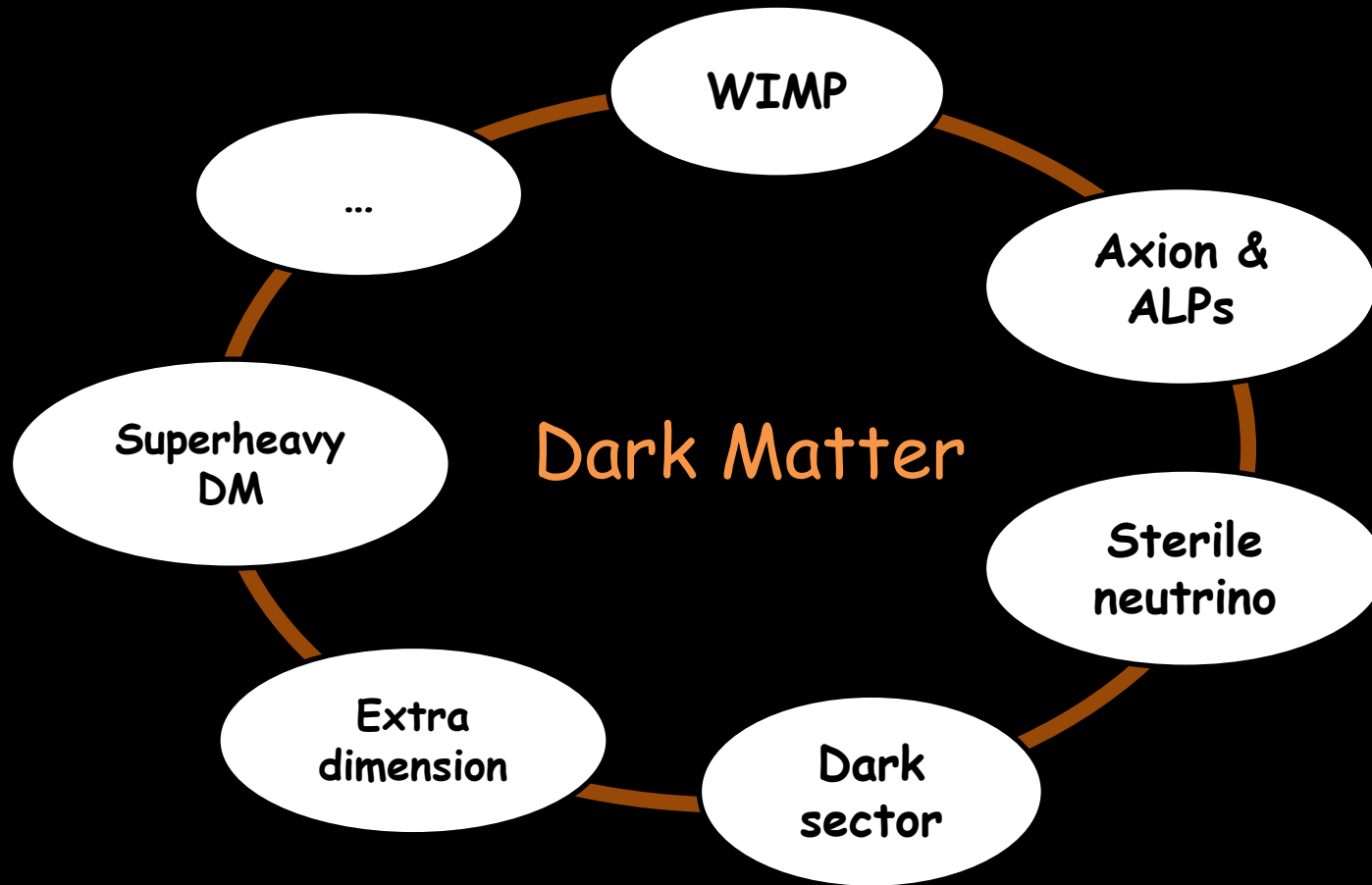


Rotation curve



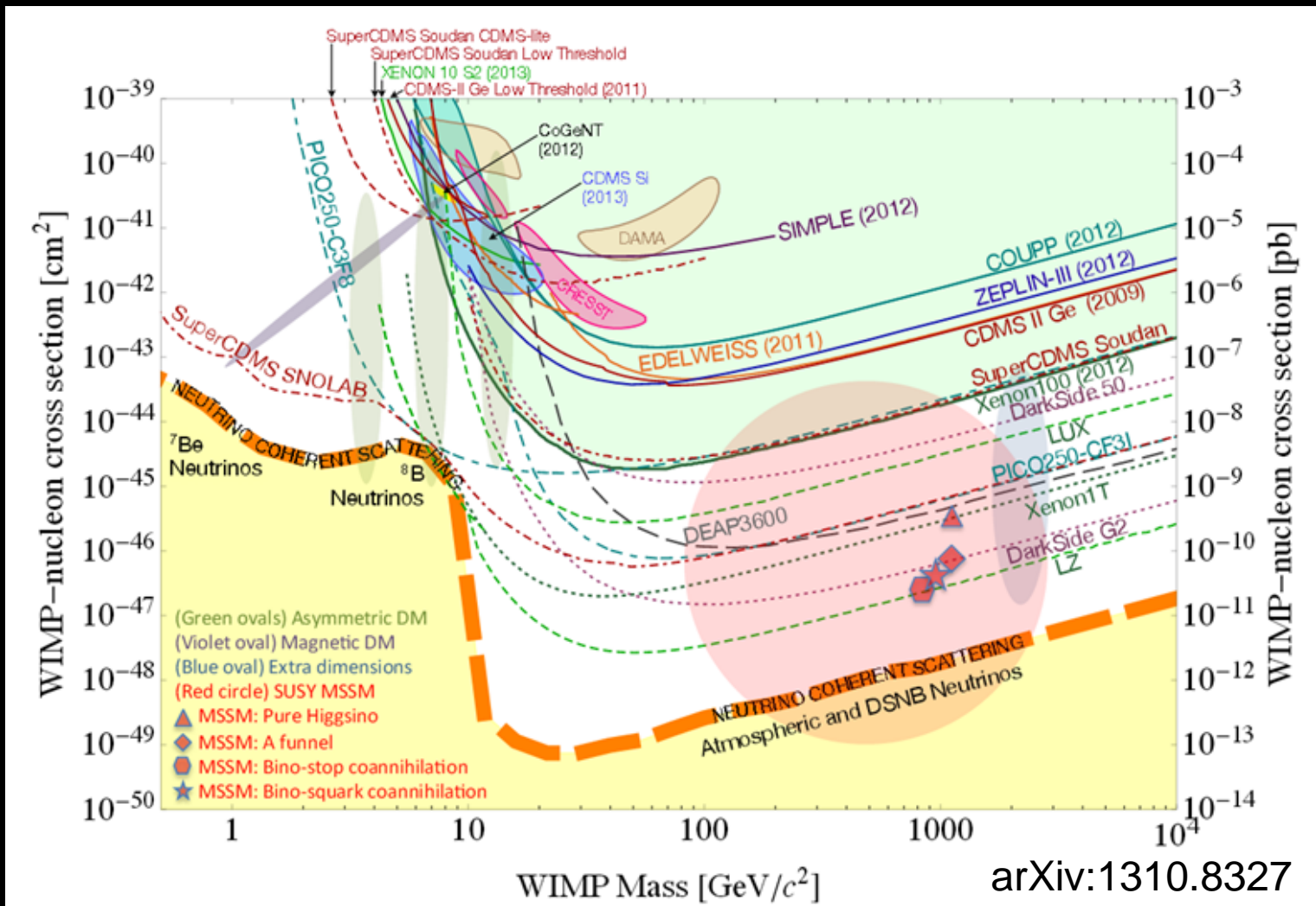
Strong evidence for dark matter

# A dizzying list of candidates...



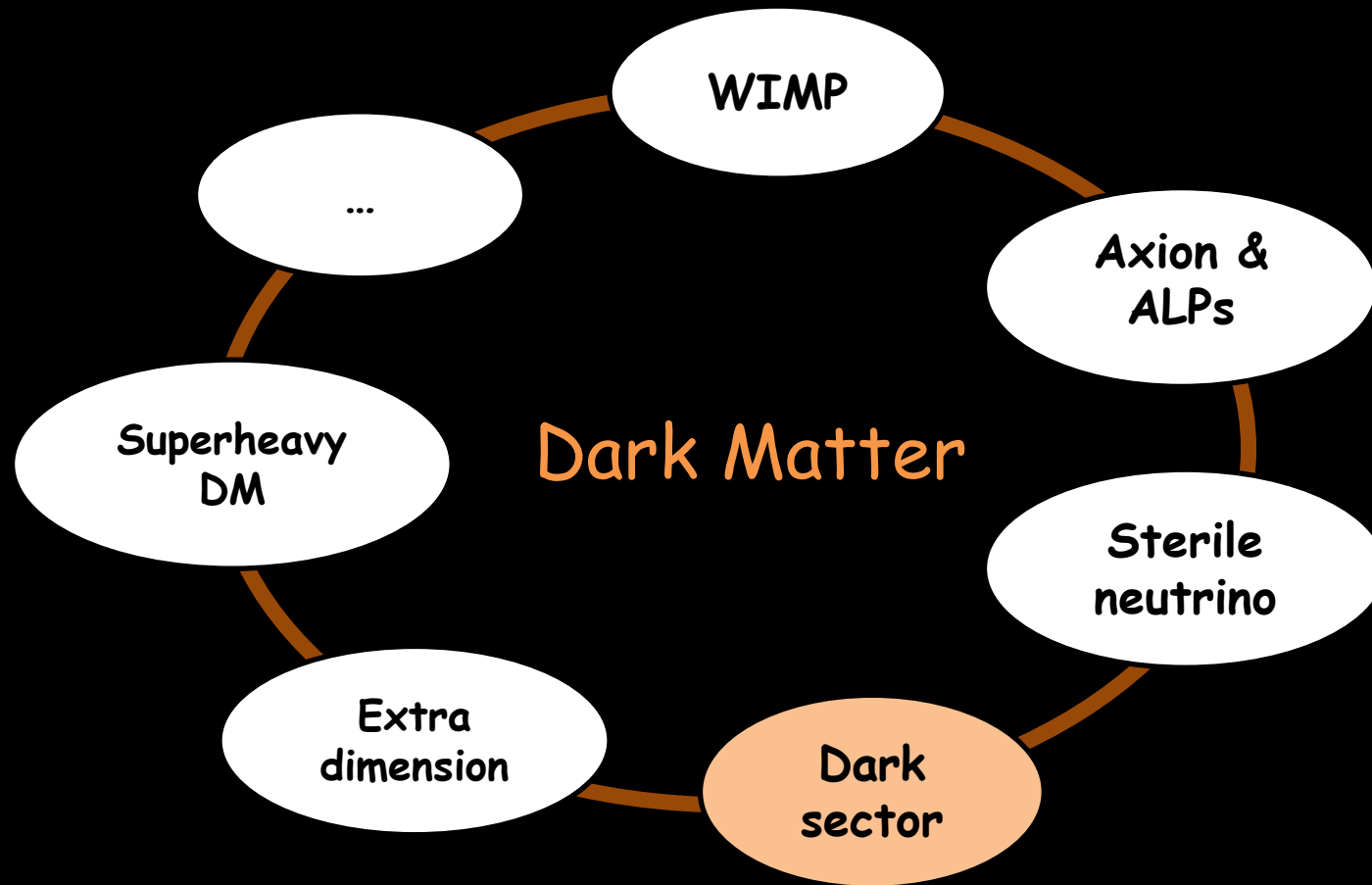
The WIMP (Weakly Interacting Massive Particle) paradigm is often considered as the most appealing scenario.

# Direct dark matter detection program



So far, no sign of WIMP and New Physics at the LHC!

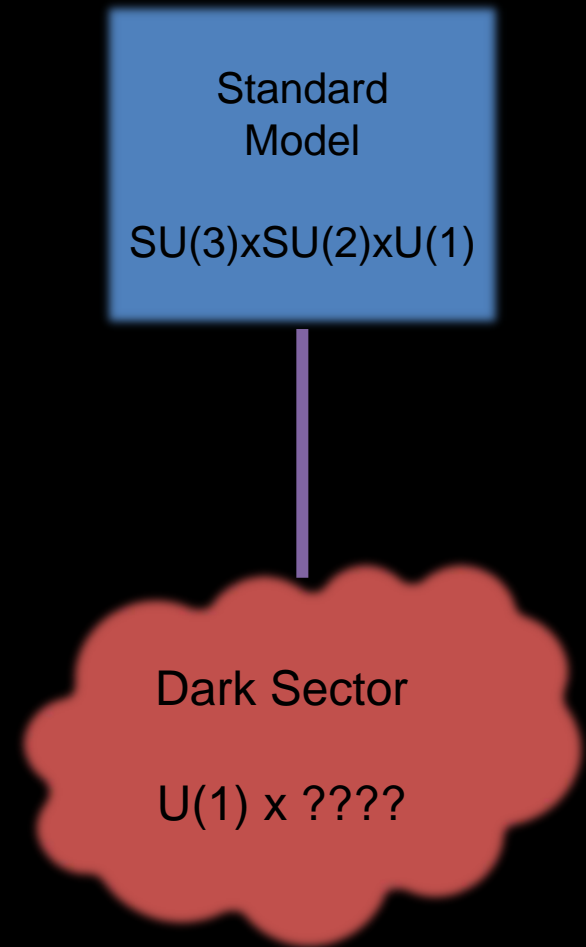
# A dizzying list of candidates...



Recent results from the LHC and direct detection experiments "challenge" the traditional WIMP paradigm and motivate the exploration of new ideas.

## A new possibility - dark sectors

- Recent anomalies observed by satellite and terrestrial experiments have motivated dark matter models introducing a **new sector with a 'dark' force**.
- **Dark sector** = new particles that do not couple directly to the SM content, but...
- There are **"portals"** between the dark sector and the SM.
- Implications for astrophysics, cosmology and particle physics.
- In particular, **low-energy colliders and fixed target experiments** offer an ideal environment to probe these new ideas.





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## Notation confusion

dark sector  
=  
hidden sector  
=  
secluded sector

Tip:  
Do not try to google "dark sector" anymore, use hidden sector instead!



# Dark sectors

## There might be dark sectors

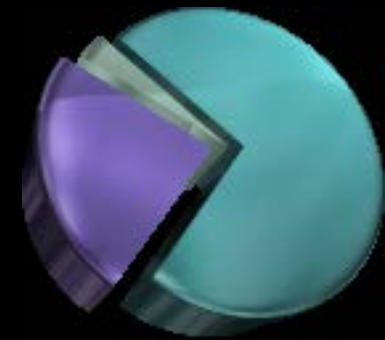
- New sectors that don't couple directly to the Standard Model.
- Theoretically motivated: string theory and many BSM scenarios include dark sectors with extra  $U(1)$ .
- Holdum's question ('86) : are there additional  $U(1)$ ? (PLB 166 (1986) 196)
- Dark photons ( $A'$ ) are the corresponding  $U(1)$  gauge bosons, mediating this dark force.

## Dark matter could be part of a dark sector

- Dark matter and other new particles may reside in dark sectors.
- Could have a very rich structure.

$SU(3)_C \times SU(2)_L \times U(1)_Y$

$U(1)_X \times ???$   
 $U(1)_Y \times ???$   
 $???$



## How could we detect them?

- Interaction between dark sector and SM occurs through high-dimension operator, referred to as "portals". At low-energy, the "vector portal" is dominant.

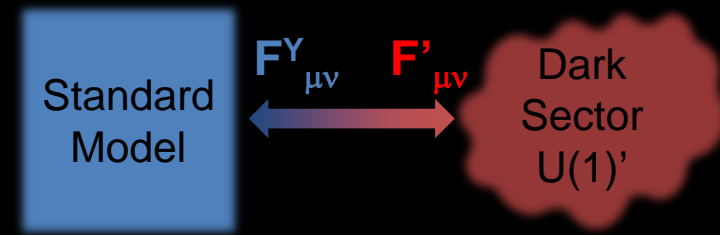


# Dark sector and vector portal

- Dark sector with a new gauge group  $U(1)'$  (similar to QED)
- One can add an effective interaction of the following form to the SM

$$\Delta\mathcal{L} = \frac{\varepsilon_Y}{2} F^{Y,\mu\nu} F'_{\mu\nu}$$

between the SM hypercharge and  $U(1)'$  fields, called **kinetic mixing**, with a mixing strength  $\varepsilon_Y$



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# Dark sector and vector portal

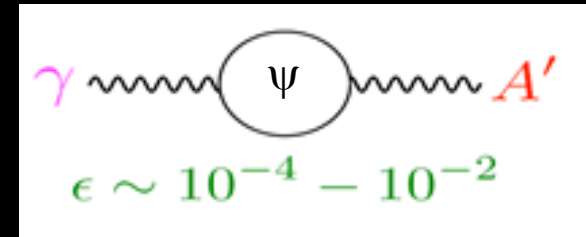
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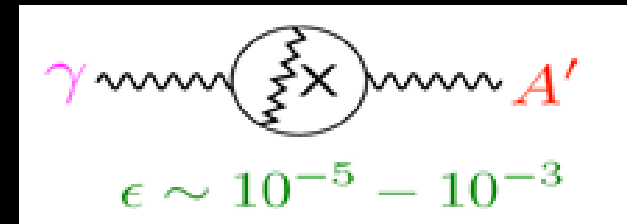
between the SM hypercharge and  $U(1)'$  fields, called **kinetic mixing**, with a mixing strength  $\varepsilon_Y$

- Could be realized by new heavy particles charged under both gauge groups.

heavy particle  $\psi$  with both dark and EM charges.



GUT (2 loops)



( $\rightarrow 10^{-7}$  if both  $U(1)$ 's are in unified groups)

typically  $\varepsilon_Y \sim 10^{-5} - 10^{-2}$

e.g. Arkani-Hamed & Weiner;  
Cheung, Ruderman, Wang, Yavin;  
Morrissey, Poland, Zurek;  
Essig, Schuster, Toro;

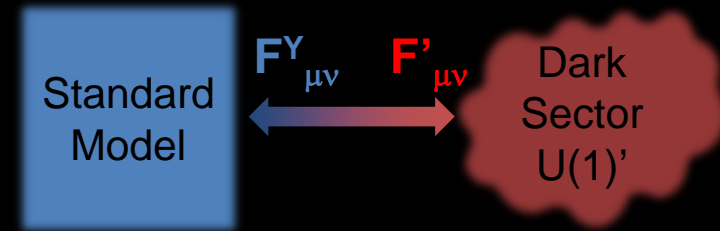
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between the SM hypercharge and  $U(1)'$  fields, called **kinetic mixing**, with a mixing strength  $\varepsilon_Y$

- Could be realized by new heavy particles charged under both gauge groups.
- After EWSB, there is a coupling between the dark photon and the photon (also the  $Z$ ), i.e. **a dark photon - SM fermion coupling**.



$$\Delta\mathcal{L} = \frac{\varepsilon_Y}{2} F^{Y,\mu\nu} F'_{\mu\nu}$$

↓ EWSB

$$\Delta\mathcal{L} = \frac{\varepsilon}{2} F^{EM,\mu\nu} F'_{\mu\nu} (+Z)$$

$$\varepsilon = \varepsilon_Y \cos \theta_w$$

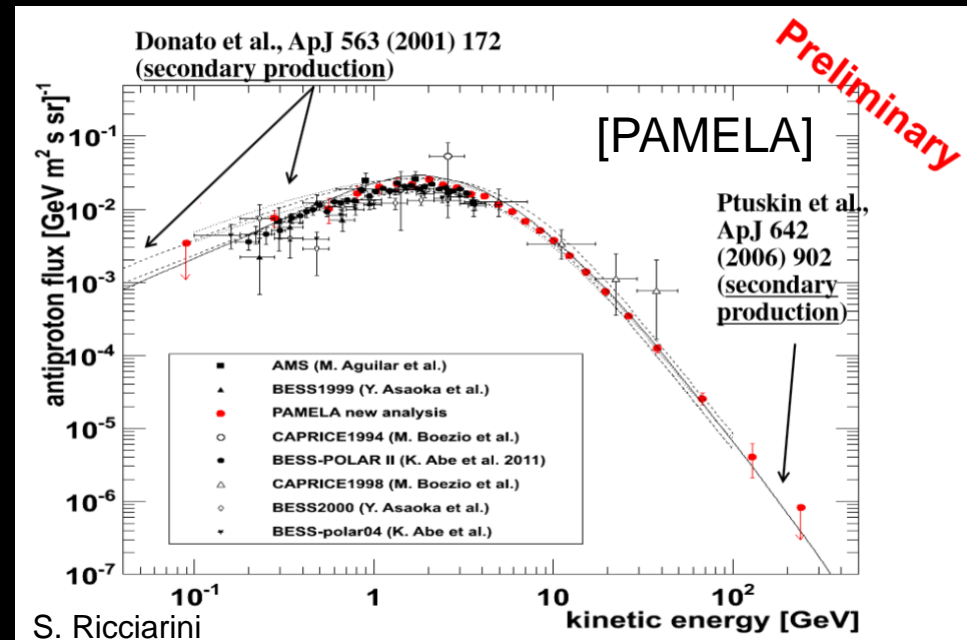
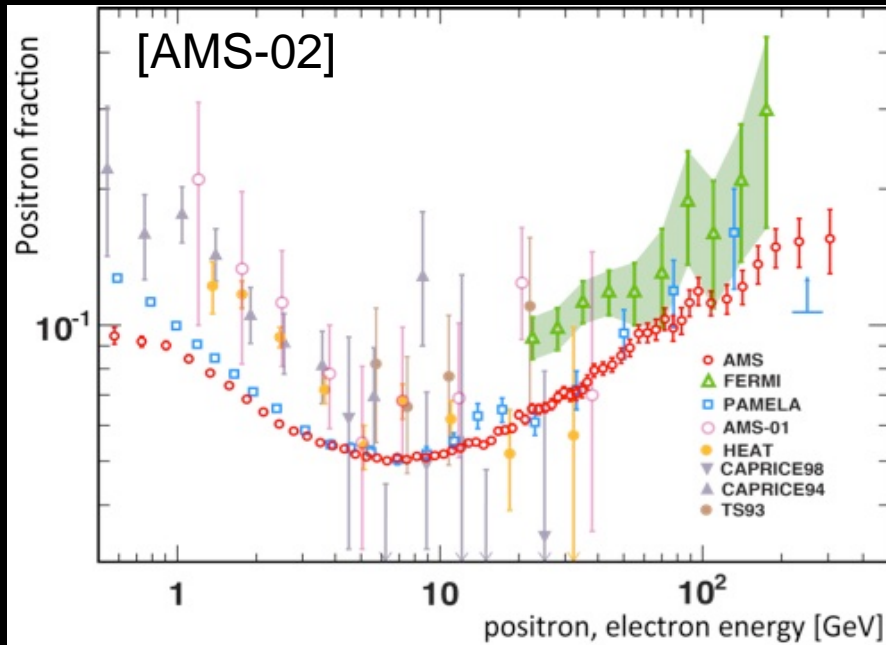
dark photon - SM fermion coupling with strength  $\alpha' = \varepsilon^2 \alpha$

Connection to dark matter?

# A few years ago... new astrophysical signals

Excess of electrons/positrons in the cosmic rays, first seen by Pamela, confirmed by Fermi & AMS-02.

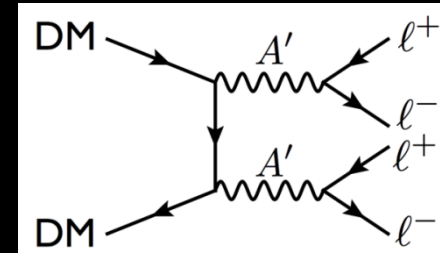
No comparable enhancement of antiprotons!



Could be explained by a simple dark sector model

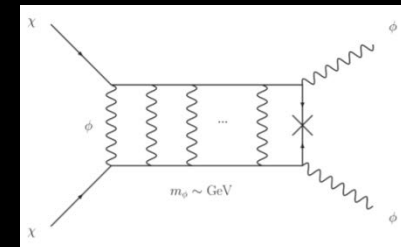
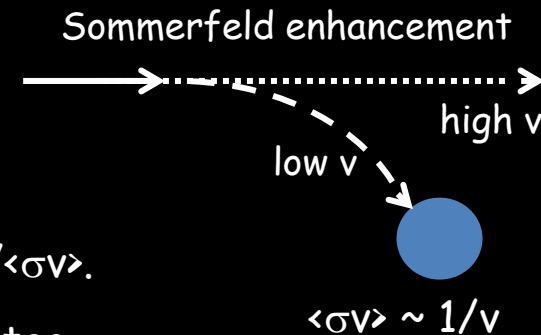
# The original idea: a light dark sector model

Wimp-like TeV-scale dark matter particles annihilate into light dark photons (10 MeV - few GeV range), which subsequently decay to electrons/positrons (Arkani-Hamed et al., Pospelov & Ritz):



- Large branching fraction to leptons
- Protons kinematically suppressed
- Hard energy spectrum
- Correct relic abundance with Sommerfeld enhancement

- Relic abundance depends on annihilation rate  $\Omega_{\text{DM}} \sim 1/\langle\sigma v\rangle$ .
- Annihilation rate derived from cosmic flux gives  $\Omega_{\text{DM}}$  too low by a factor 100-1000 ("boost" factors invoked to solve this problem for many models).
- Cross-section is enhanced at low velocities for light  $A'$ , boosting  $\Omega_{\text{DM}}$  to observed values.

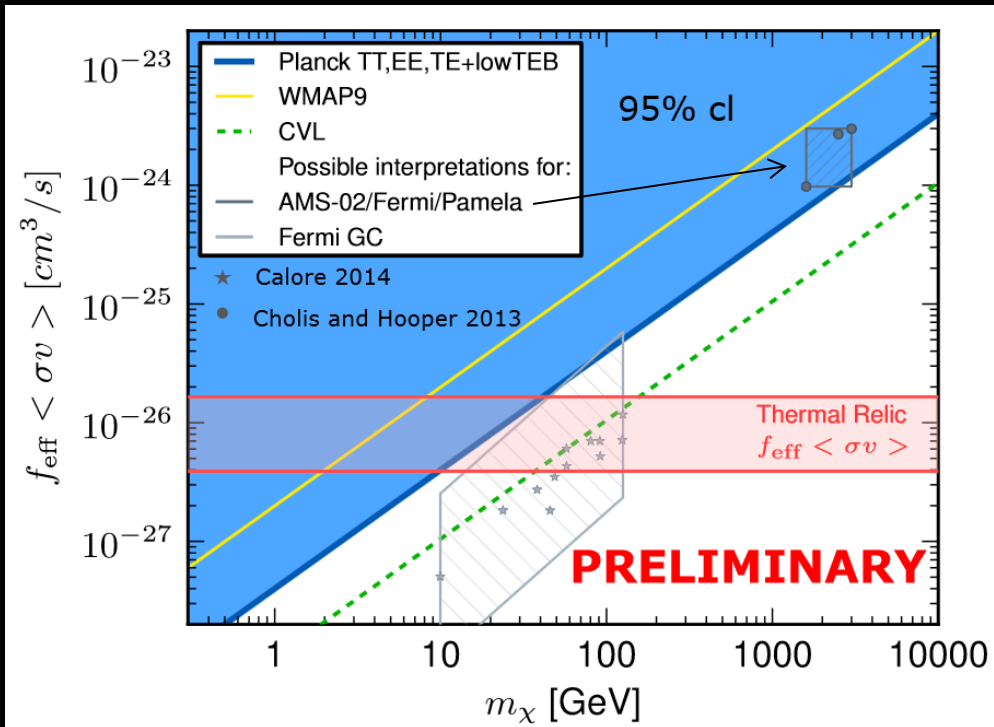
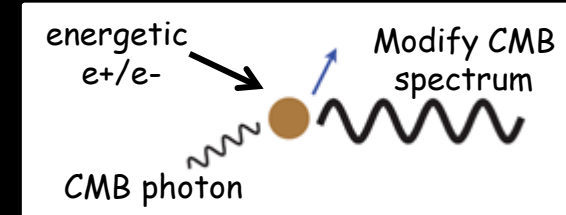


Sparked a wide interest in this class of models, there is just a tiny issue....

# Cosmological constraints

If DM annihilation into light dark photons is the source of  $e^-/e^+$  excess, other astrophysical phenomena should be observable (e.g. diffuse gamma ray emission, CMB).

In particular, primordial DM annihilation injects energy in the CMB  $\rightarrow$  distorts spectrum



Severe constraints from recent Planck measurement

There are still a few uncertainties, but plausible dark matter models of the Pamela excess that could explain all the current constraints are a very specific subset

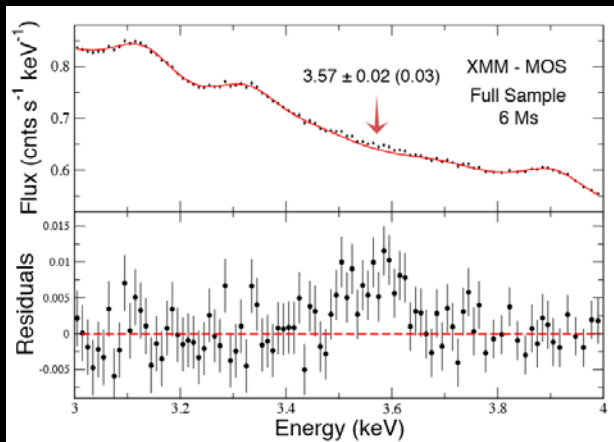
Any other anomalies?

Planck collaboration

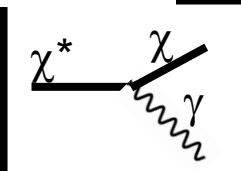


# Other anomalies

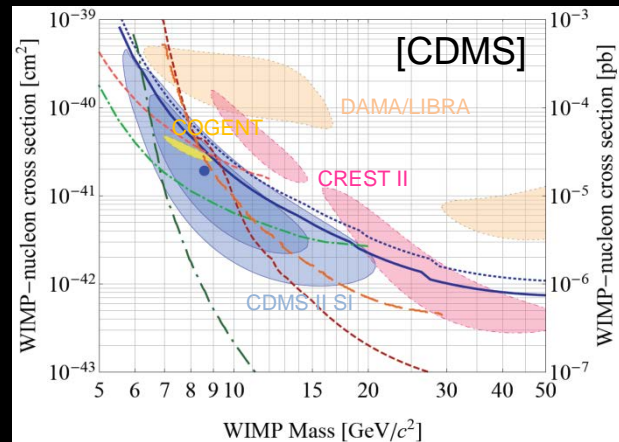
## Line at 3.55 keV



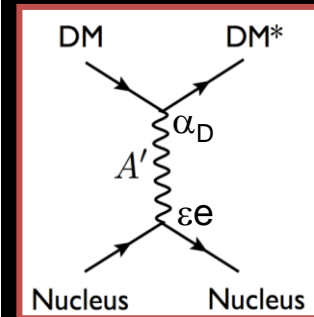
[Bulbul *et al*],  
[Boyarisky *et al*]



## Direct detection anomalies

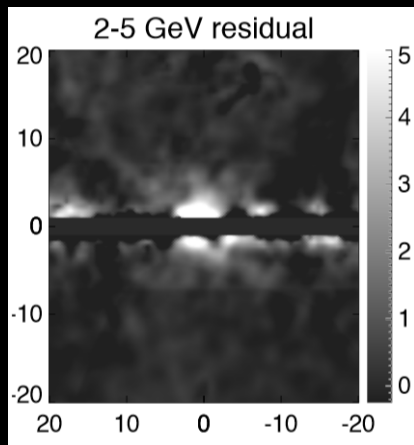


CDMS, arXiv:1304.4279



[Slatyer,  
Schuster&Toro,...]

## Galactic center



[Daylan *et al*],

And many others...

Could have other explanation: pulsars,  
instrumental effects, other new particles,...

When you have a dark hammer, you tend  
to see everything as a dark nail!

## At this point...

### New theory of dark matter based on dark sector(s)

- Light new mediator (dark photon  $A'$ ) with a MeV - GeV mass
- Mixing between dark sector - SM with  $\varepsilon \sim 10^{-5} - 10^{-2}$  (could be smaller)
- Could have a rich structure

### Anomalies from astrophysical data, direct detection and precision measurements

- Could be explained by dark sector
- Could have another origin, be statistical fluctuations or instrumental effects
- Dark matter could be composite with a dark sector sub-component
- ...

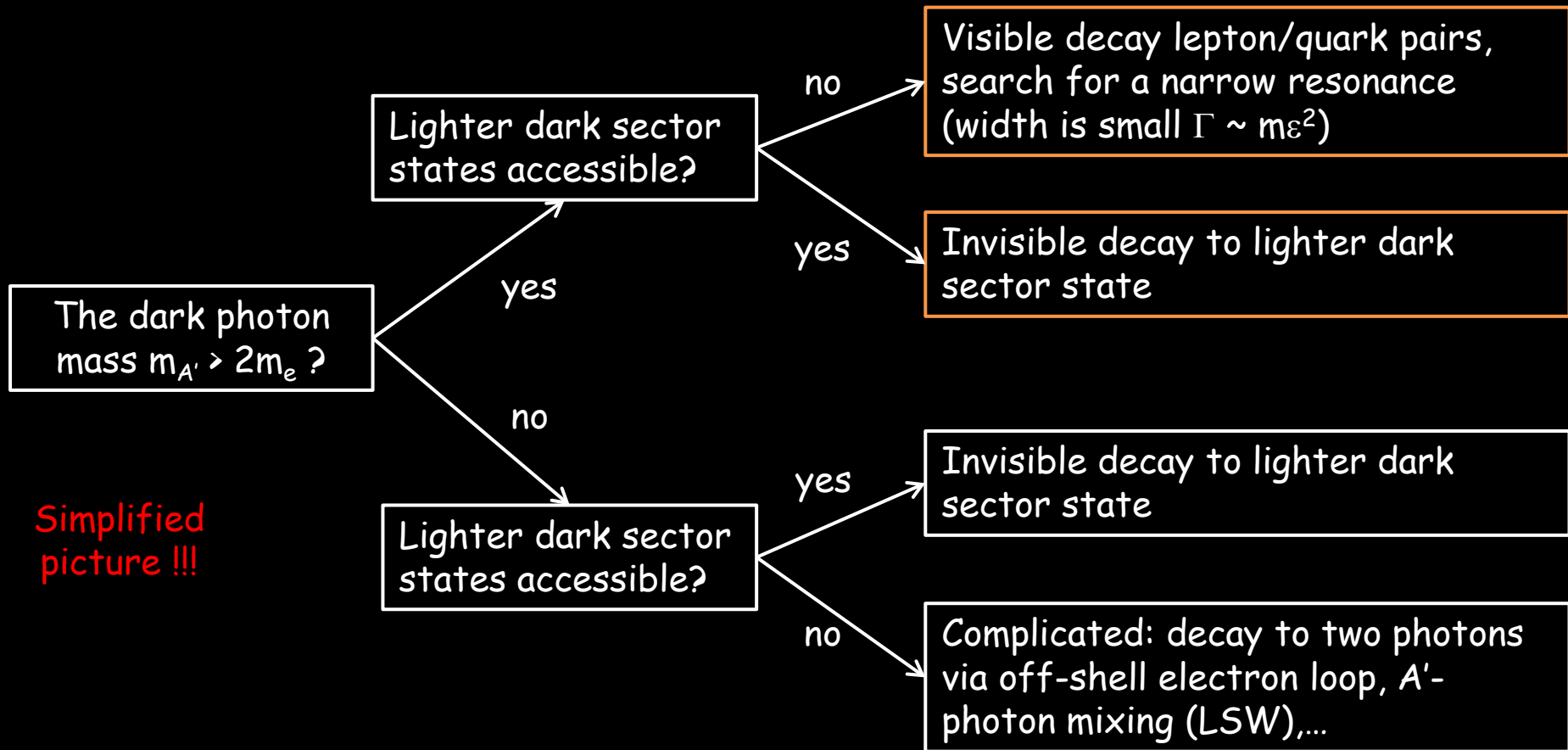
But it made us realize the amazing possibilities at the GeV-scale in a more general context, and the possibilities to probe them in laboratory at low energies !



Probing dark sectors  
at low-energy (and high-energy) colliders

# Particle physics implications

Particle physics experiments can produce dark photons. In fact, photons in any process can be replaced by dark photons (with an extra factor of  $\varepsilon^2$ ).

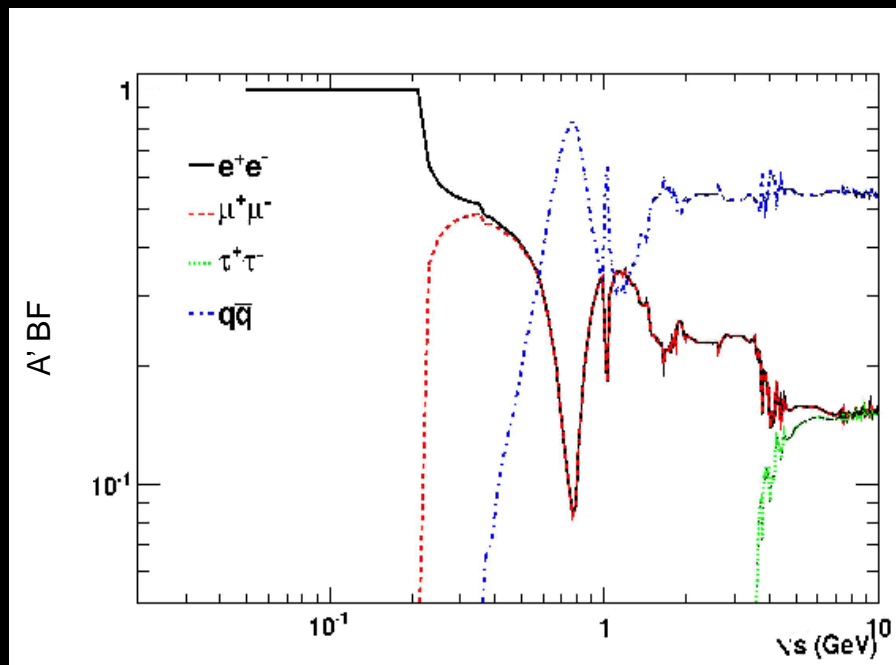


Search strategies depends on the mass hierarchy.

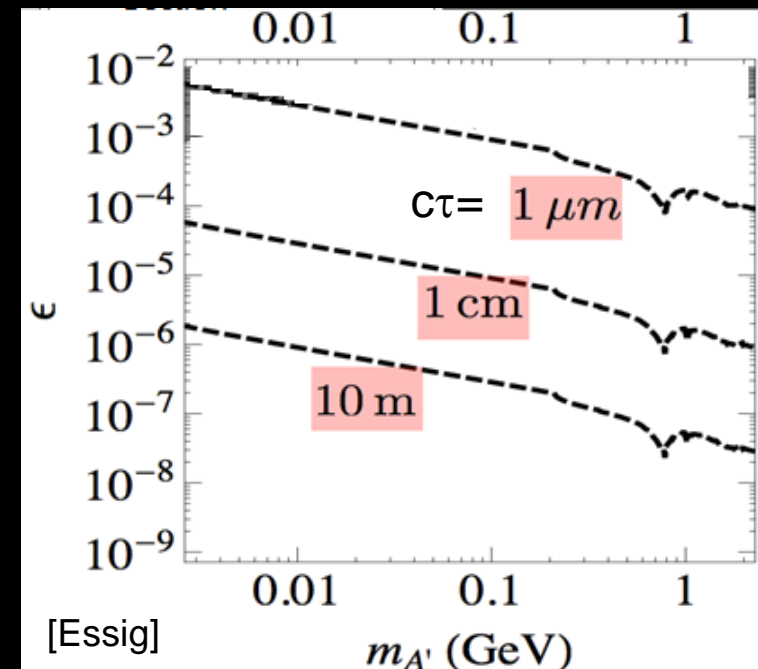
# Particle physics implications

Dark photon branching fraction into leptons depends only on the fermion electric charge.

Dark photon is small ( $\sim m\varepsilon^2$ ) and could be short or long-lived, depending on the parameters of the theory. Dark photon decays can either be prompt or displaced (visible case)



Lepton contribution dominates at low masses, still  $\sim 30\%$  at high masses!

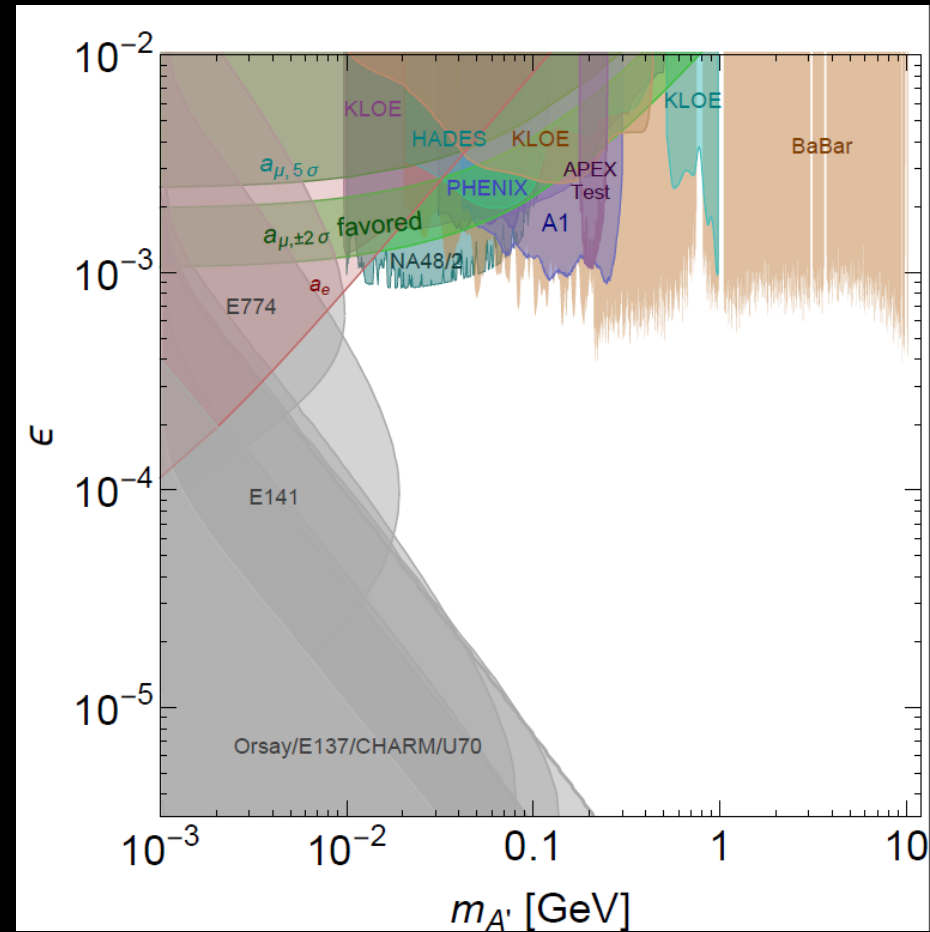


# Particle physics implications

Current constraints on the mixing parameter  $\varepsilon$  vs. the dark photon mass  $m_{A'}$  for visible  $A'$  decays

- electron/muon  $g-2$ ,
- beam dump experiments
- fixed target experiments
- neutral meson decays
- $e^+e^-$  colliders

Constraints from many type of experiments probing different regions of parameter space.



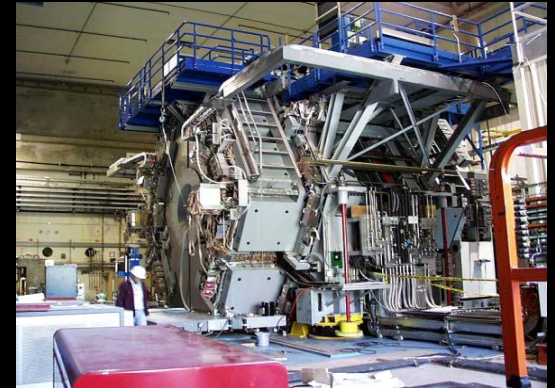
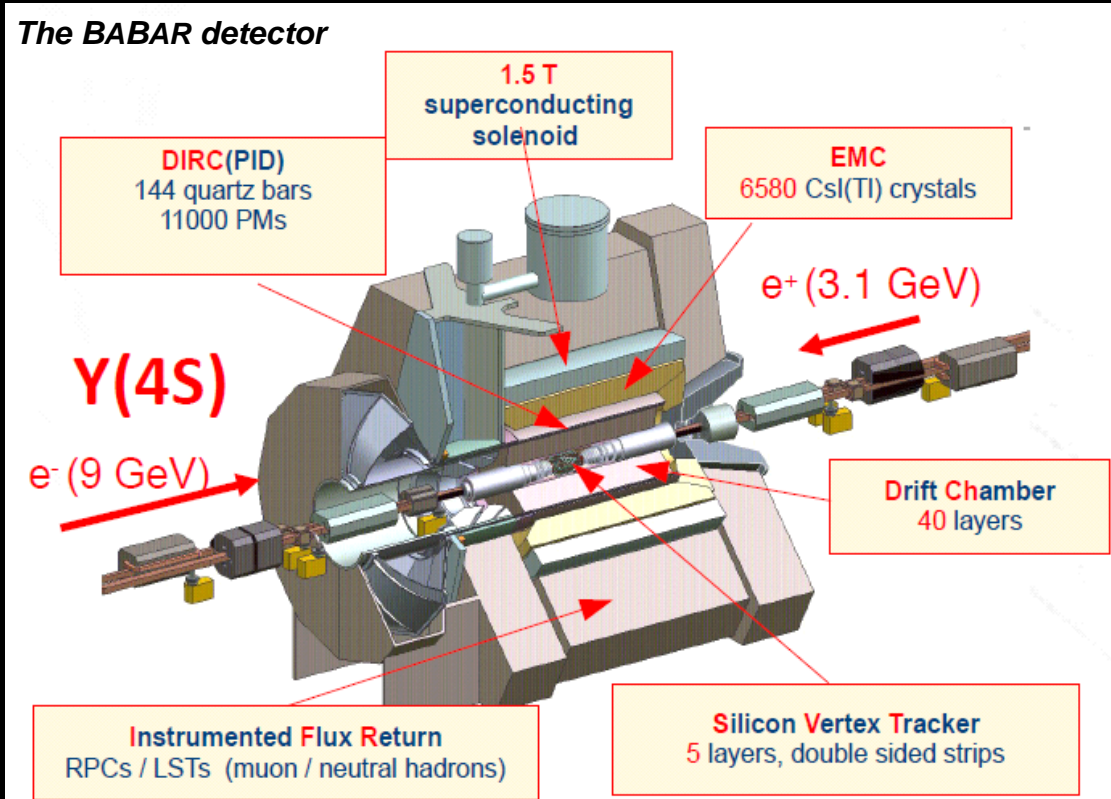
See arXiv:1311.0029

Low-energy high-luminosity  $e^+e^-$  colliders offer a low-background environment to search for MeV/GeV-scale dark sector (in particular high masses) and probe their structure



# The BABAR experiment

BABAR collected around  $500 \text{ fb}^{-1}$  of data around the  $\Upsilon(4S)$  resonance



B-factories offer an ideal environment to search for dark sector particles

## Search for dark photon

$$e^+e^- \rightarrow \gamma A', A' \rightarrow e^+e^-, \mu^+\mu^-, \pi^+\pi^-$$

$$e^+e^- \rightarrow \gamma A', A' \rightarrow \text{invisible}$$

$$\pi^0 \rightarrow \gamma l^+l^-, \eta \rightarrow \gamma l^+l^-, \phi \rightarrow \eta l^+l^-, \dots$$

## Search for dark Higgs boson

$$e^+e^- \rightarrow h' A', h' \rightarrow A' A'$$

$$e^+e^- \rightarrow h' A', h' \rightarrow \text{invisible}$$

## Search for dark boson(s)

$$e^+e^- \rightarrow \gamma A' \rightarrow W' W'$$

## Search for dark hadrons

$$e^+e^- \rightarrow \pi_D + X, \quad \pi_D \rightarrow e^+e^-, \mu^+\mu^-$$

## Search for dark scalar (s) and dark pseudoscalar (a)

$$B \rightarrow K^{(*)} s \rightarrow K^{(*)} l^+l^-$$

$$B \rightarrow K^{(*)} a \rightarrow K^{(*)} l^+l^-$$

$$B \rightarrow ss \rightarrow 2(l^+l^-)$$

$$B \rightarrow K 2(l^+l^-)$$

$$B \rightarrow 4(l^+l^-)$$

## Search for “muonic/tauonic dark force”

$$e^+e^- \rightarrow \mu^+\mu^- Z', Z' \rightarrow \mu^+\mu^-, \tau^+\tau^-, \text{inv.}$$

$$e^+e^- \rightarrow \tau^+\tau^- Z', Z' \rightarrow \mu^+\mu^-, \tau^+\tau^-, \text{inv.}$$

## Search for leptophilic dark scalar

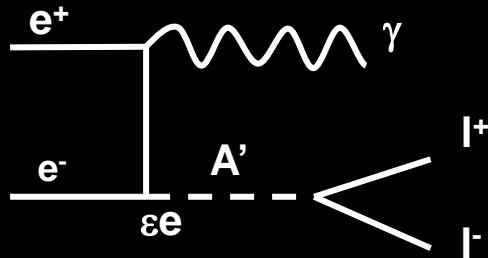
$$e^+e^- \rightarrow \tau^+\tau^- h', h' \rightarrow \mu^+\mu^- \text{ (4 leptons + MET)}$$

Large set of channels (few experiments can explore all these at once), can study the properties of a dark sector in detail

# Direct dark photon production

A dark photon can be produced in

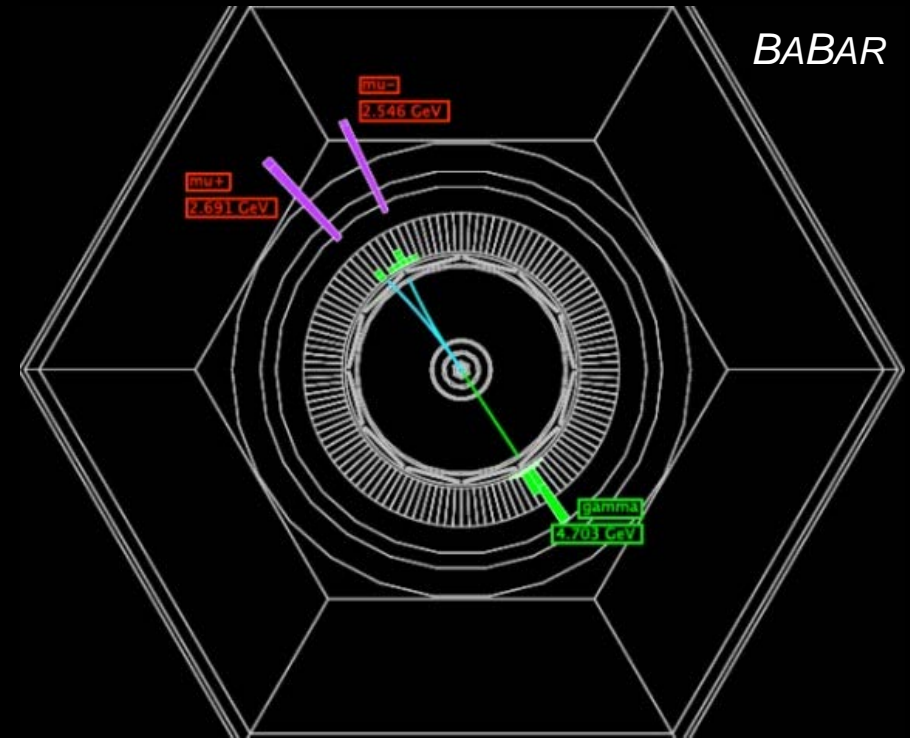
$$e^+e^- \rightarrow \gamma A', A' \rightarrow e^+e^-, \mu^+\mu^-$$



## Event selection

- 2 tracks + 1 photon
- Constrained fit to the beam energy and beam spot
- Particle identification (e/mu)
- Kinematic cuts to improve purity
- Quality cuts on tracks and photons

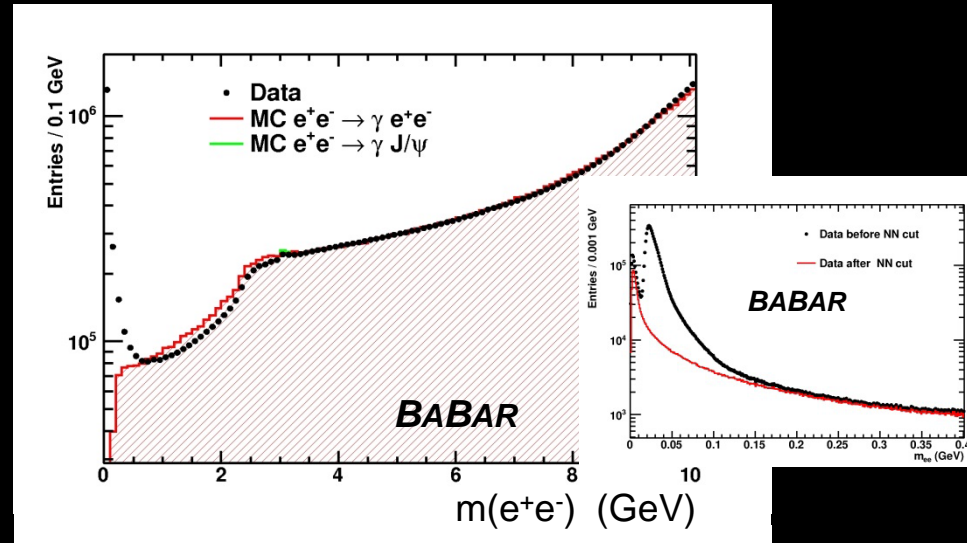
Typical event



- Tracks
- Photon
- Signal in muon/hadron detector

## Di-electron mass spectrum

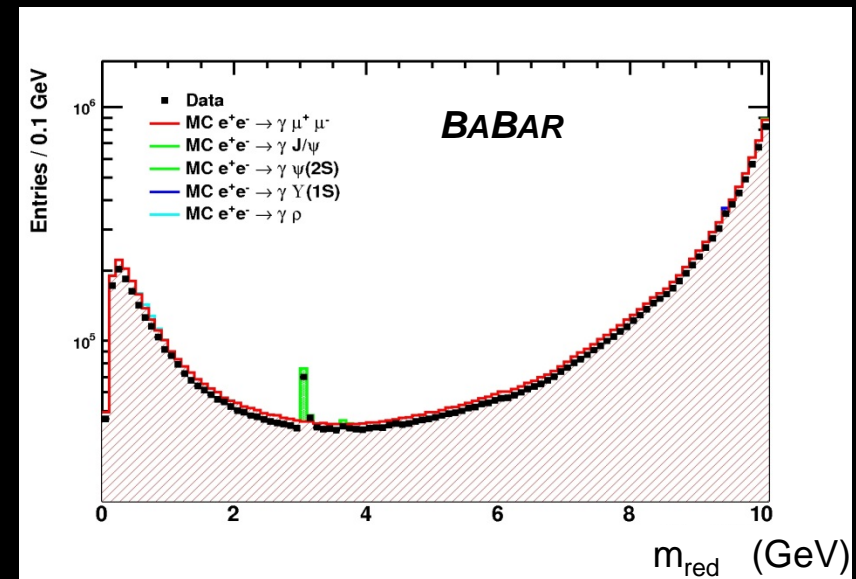
- Globally well reproduced by BHWIDE above 1 GeV, cut-off in the MC (co-linear tracks) affects low mass region. Madgraph reproduces well the low mass region.
- Background from photon conversions suppressed by neural network



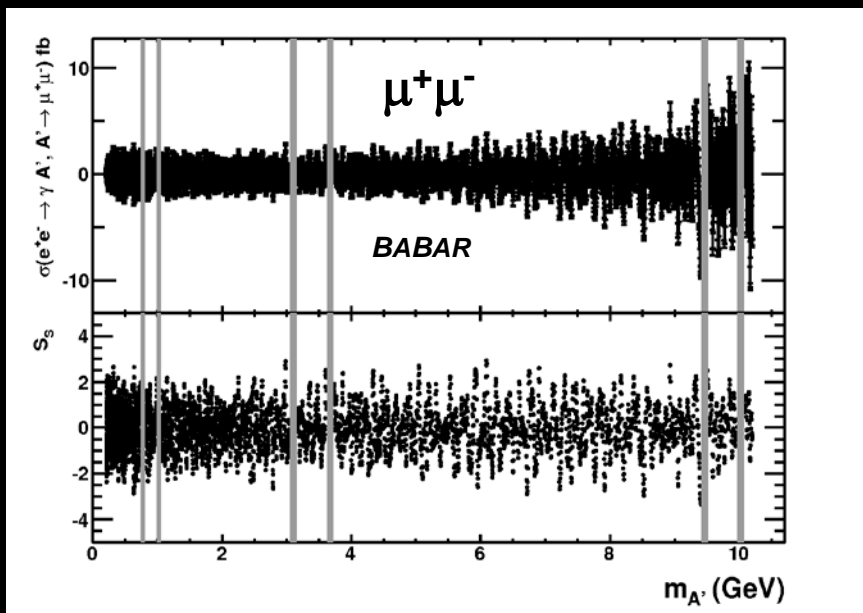
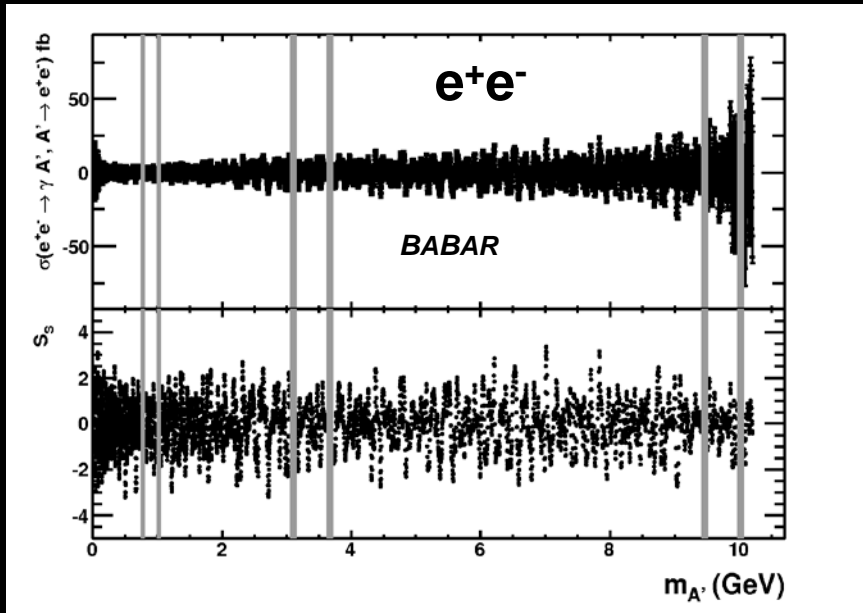
## Di-muon mass spectrum

- Plot the reduced mass (smoother near threshold):  $m_{\text{red}} = (m_{\mu\mu}^2 - 4 m_{\mu}^2)^{1/2}$
- Globally well reproduced by KK2F, correct for differences in efficiencies

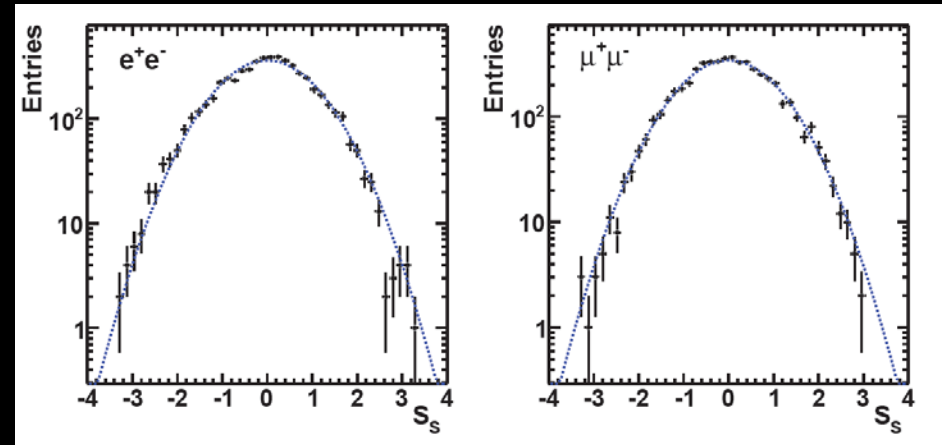
Good data-MC agreement at the  $J/\psi$ ,  $\Psi(2S)$ ,  $\Upsilon(1S)$  resonances







## Distribution of statistical significances

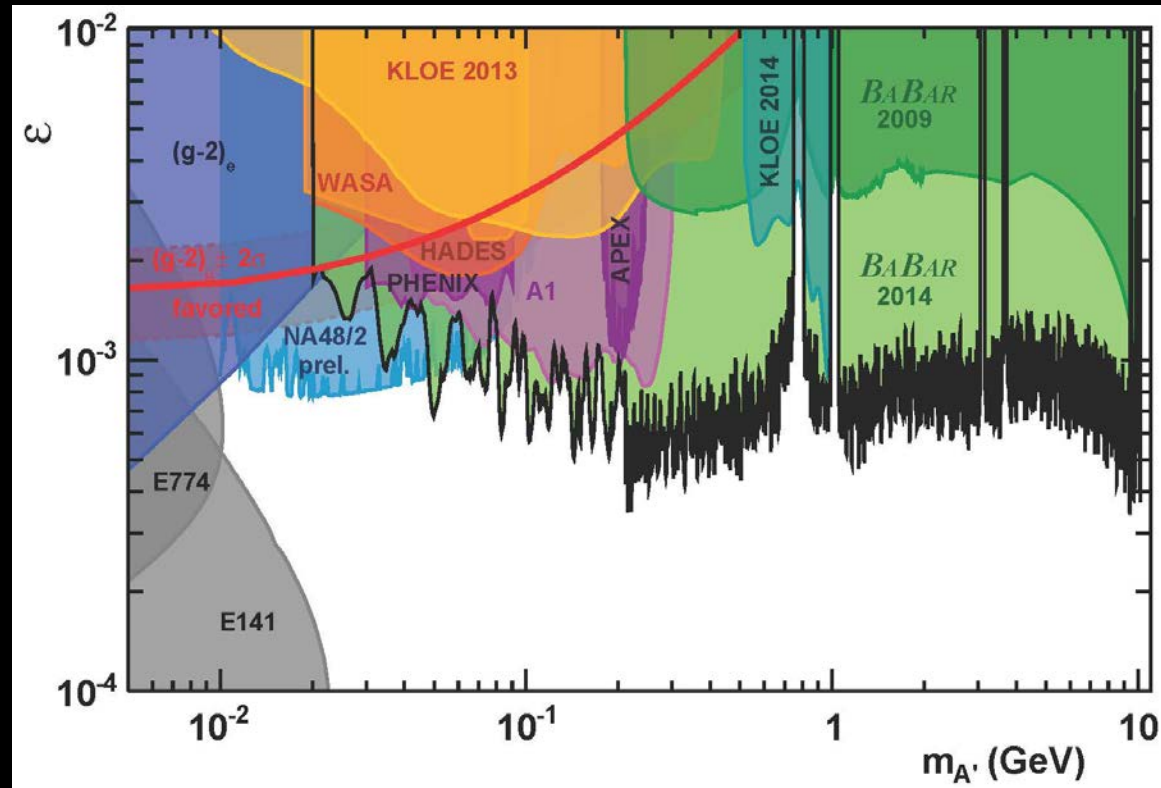


## Largest significances:

- $3.4\sigma$  for electrons @  $7.02 \text{ GeV}$   
→  $0.6\sigma$  with trial factors
- $2.9\sigma$  for muons @  $6.09 \text{ GeV}$   
→  $0.1\sigma$  with trial factors

**Consistent with null hypothesis**

- Low mass has from large backgrounds and sub-optimal trigger efficiency, but still competitive
- The  $e^+e^- \rightarrow \gamma A', A' \rightarrow \pi^+\pi^-$  final state can further probe the region near the  $\rho$  meson, currently difficult to access by other experiments.



Together with PHENIX and NA48/2, the full “g-2 band” is excluded for purely visible decays

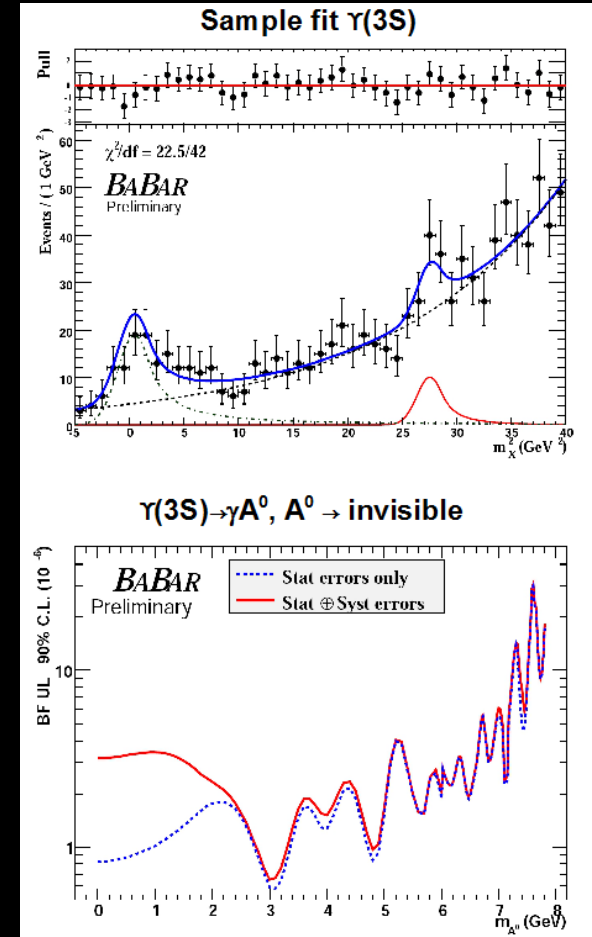
There is still plenty of interesting parameter space to explore!



## Invisible dark sector

- Several scenarios where dark photons decay to invisible final states, e.g lighter dark sector particles (sub-GeV),...
- At  $e^+e^-$  colliders, we can search for  $e^+e^- \rightarrow \gamma A'$ ,  $A' \rightarrow$  invisible by tagging the recoil photon in "single photon" events.
- Currently only a measurement of  $Y(2S,3S) \rightarrow \gamma A^0$ ,  $A^0 \rightarrow$  invisible at *BABAR* with  $A^0$  a light CP-odd Higgs

$Y(3S) \rightarrow \gamma A^0$ ,  $A^0 \rightarrow$  invisible,  
new analysis in progress +  
extension to  $A'$

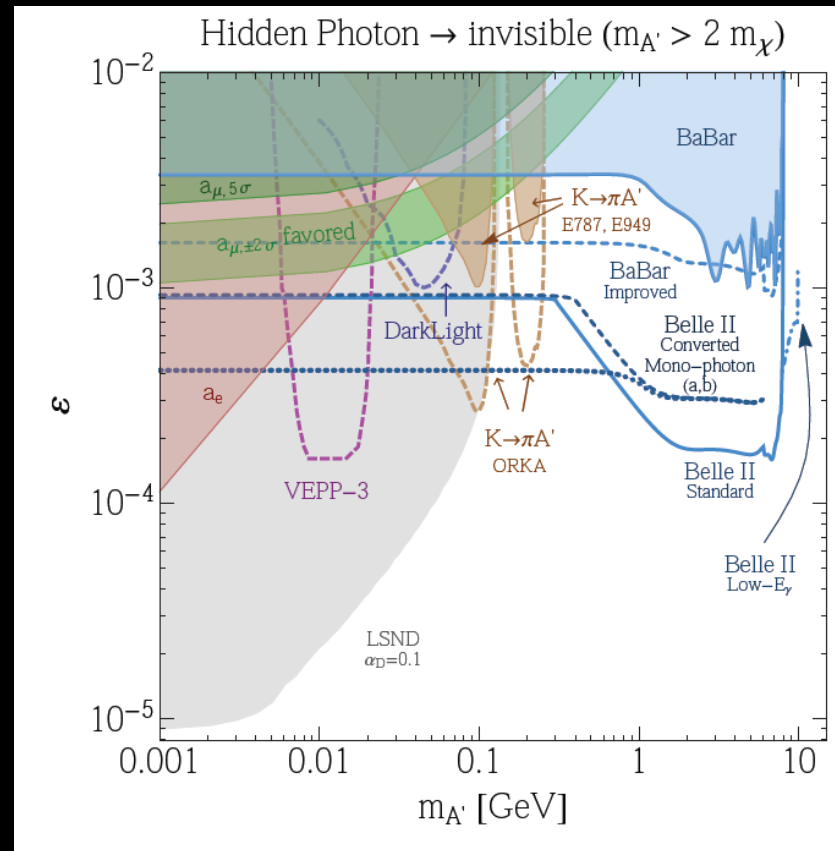


# Invisible dark photon decays

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- Analysis extended to full dataset and the dark photon case, expect **limits on  $\epsilon$  at the level of  $10^{-3}$** .
- Constraints from many other experiments!

Essig et al., arXiv:1309.5084

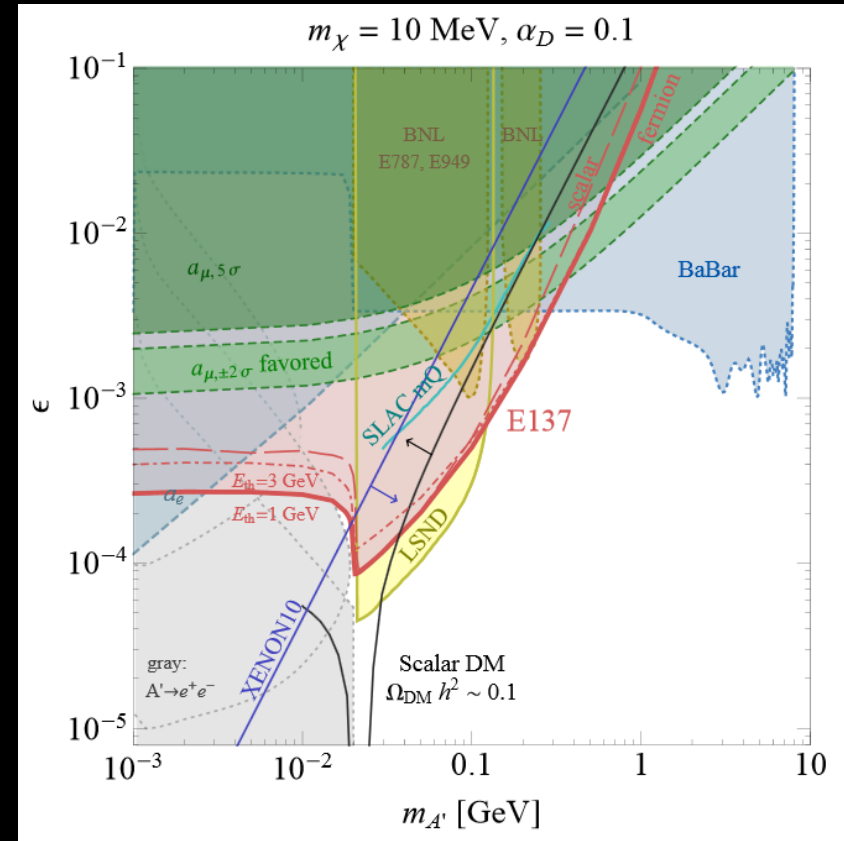


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Batell et al., arXiv:1406.2698



Some of the limits are model dependent!

# Dark Higgs boson

The dark photon mass is usually generated via the Higgs mechanism, **adding a dark Higgs boson ( $h'$ )** to the theory, which could be light.

A minimal scenario has a **single dark photon and a single dark Higgs boson.**

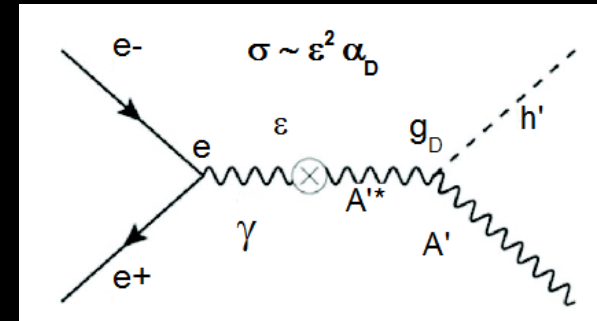
The  $h'$  could be produced in the Higgsstrahlung process  $e^+e^- \rightarrow A'^* \rightarrow h' A'$ , which is also sensitive to the dark sector coupling constant  $\alpha_D = g_D^2 / 4\pi$

Decay topology depends on the mass hierarchy:

- $m_{h'} > 2m_{A'}$  : prompt decays
- $m_{h'} < 2m_{A'}$  : displaced and invisible decays

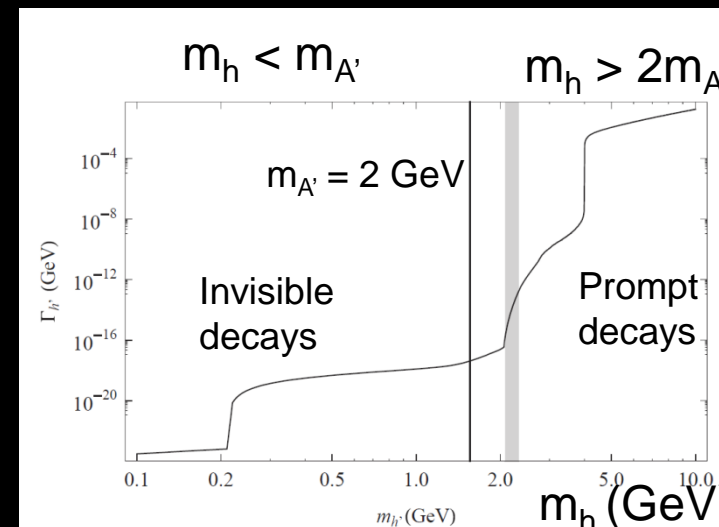
Searches for prompt  $h'$  decays at BABAR / Belle:

$$e^+e^- \rightarrow A'^* \rightarrow h' A', h' \rightarrow A' A'$$



Only suppressed by  $\epsilon^2$

Dark Higgs decay topology



# Dark Higgs boson

The dark photon mass is usually generated via the Higgs mechanism, adding a dark Higgs boson ( $h'$ ) to the theory, which could be light.

A minimal scenario has a single dark photon and a single dark Higgs boson.

The  $h'$  could be produced in the Higgsstrahlung process  $e^+e^- \rightarrow A'^* \rightarrow h' A'$ , which is also sensitive to the dark sector coupling constant  $\alpha_D = g_D^2 / 4\pi$

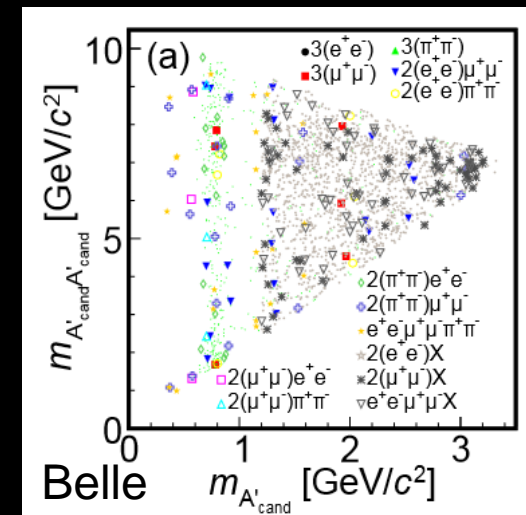
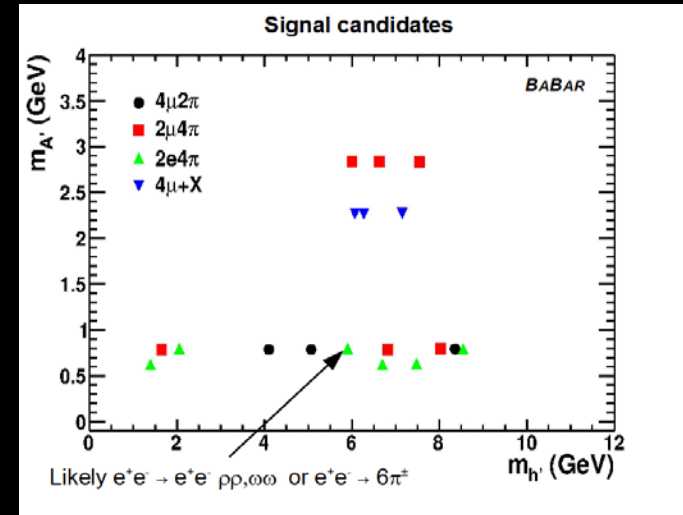
Decay topology depends on the mass hierarchy:

- $m_{h'} > 2m_{A'}$  : prompt decays
- $m_{h'} < 2m_{A'}$  : displaced and invisible decays

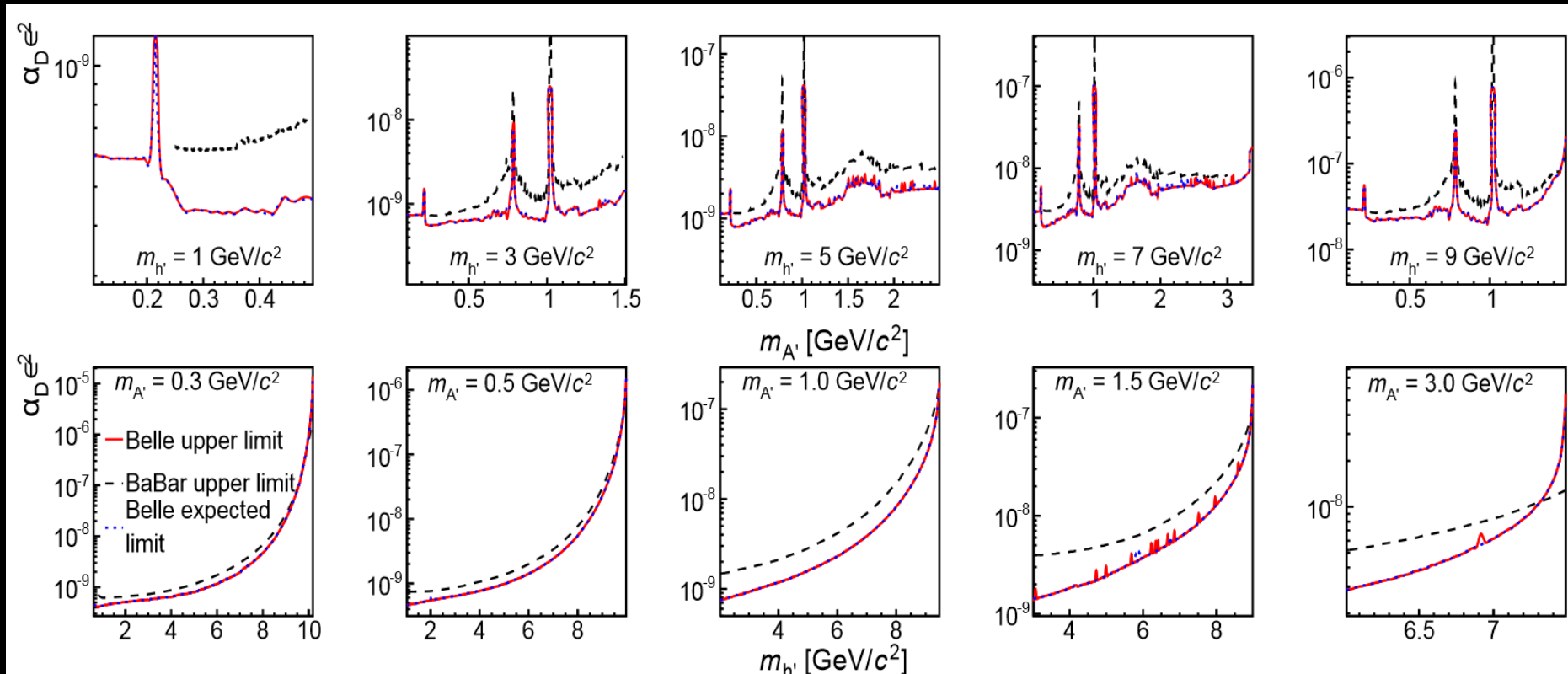
Searches for prompt  $h'$  decays at BABAR / Belle:

$$e^+e^- \rightarrow A'^* \rightarrow h' A', h' \rightarrow A' A'$$

Signal candidates (3 entries / event)



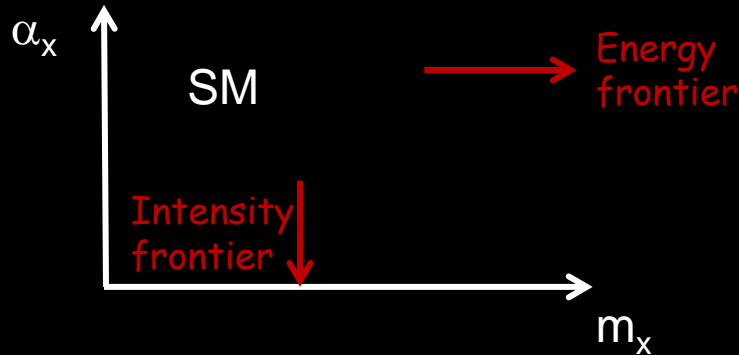
No significant signal observed, set limits on the product  $\alpha_D \varepsilon^2$



Colliders are well suited to explore these possibilities

# Dark sector searches at LHC

Direct production of dark photon suppressed at high energy



$$\text{Amplitude} \sim \frac{\alpha_x}{q^2 + m_x^2}$$

Difficult to probe  $\alpha_x < 10^{-6}$  and  $m_x \sim \text{GeV}$  at LHC (hard ISR emission also suppressed)

Instead, new particles (e.g SUSY) could decay into dark sector particles with a large BF.

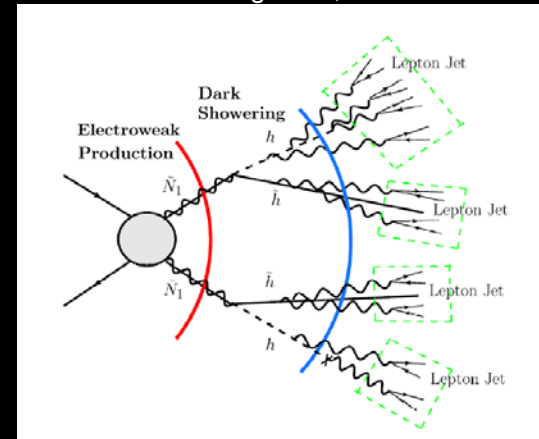
In case of SUSY, bottom of cascade no longer stable, decays into dark photons  $\rightarrow$  lepton jets.

Main characteristics:

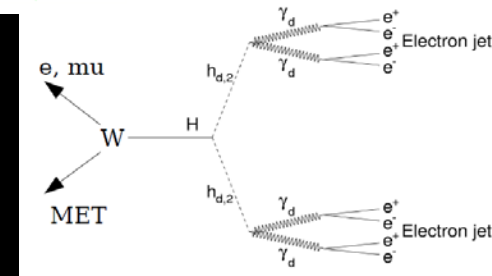
- Many leptons final state (e.g. lepton jets)
- Boosted dark sector particles  $\rightarrow$  displaced vertices

But New Physics needed in some models !!!

Cheung et al., arXiv:0909.0290



W+H  $\rightarrow$  e- jet [A. Haas]



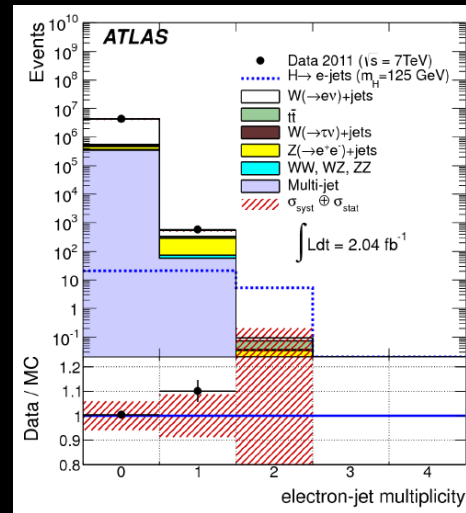


# Dark sector searches at LHC

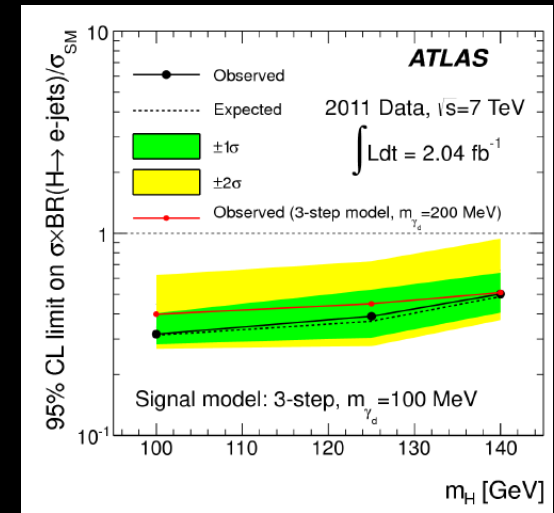
Search for

$W+H \rightarrow \text{electron-jets} + X$

No excess of events with two electron jets observed



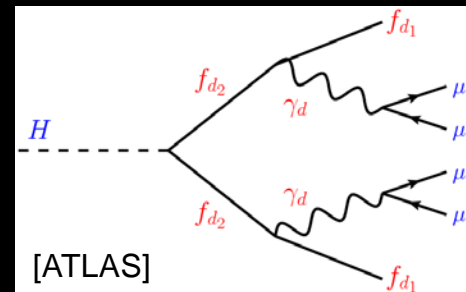
ATLAS Collab., New J.Phys. 15 (2013) 043009



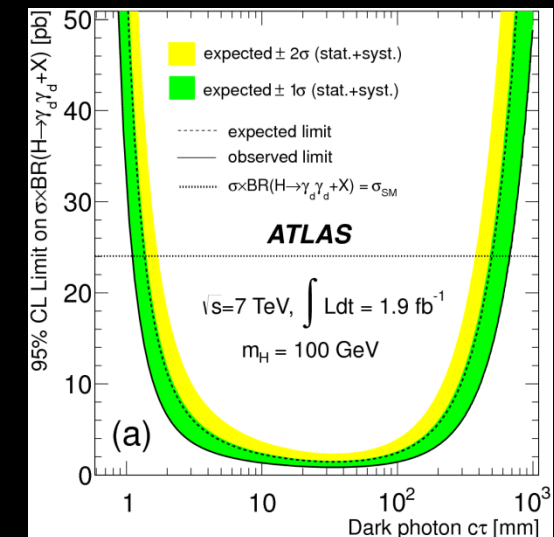
Search for

$H \rightarrow A' A' + X$

No signal observed



Atlas Collab., PLB 721 (2013) 32



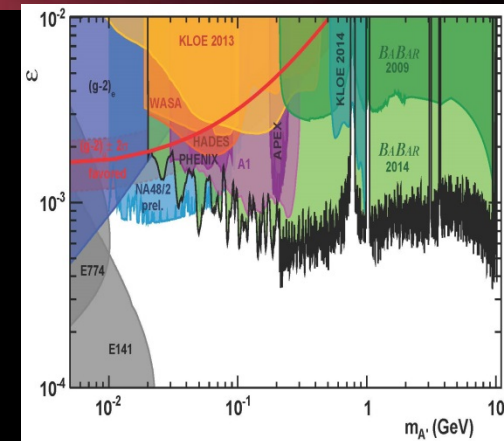
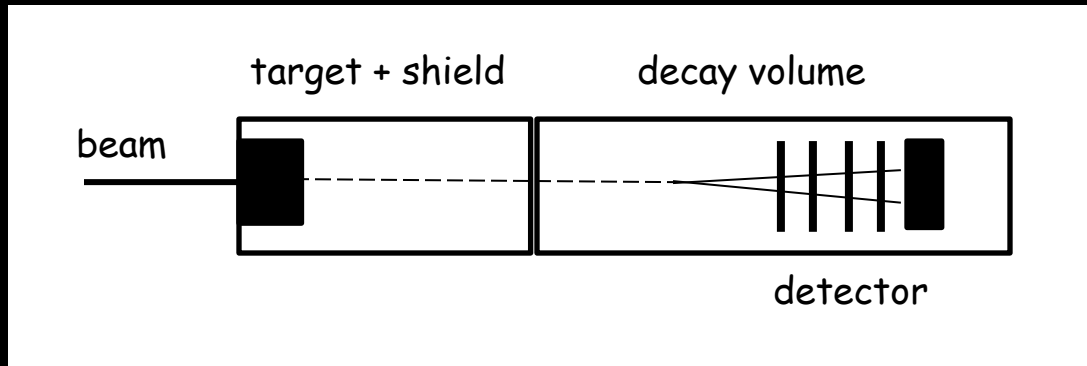
+ searches for SUSY lepton jets,  $H \rightarrow \text{muon jets}$  and possible searches for direct production, rare Z decays,...

Interesting program pursued at LHC

# Other constraints and future initiatives

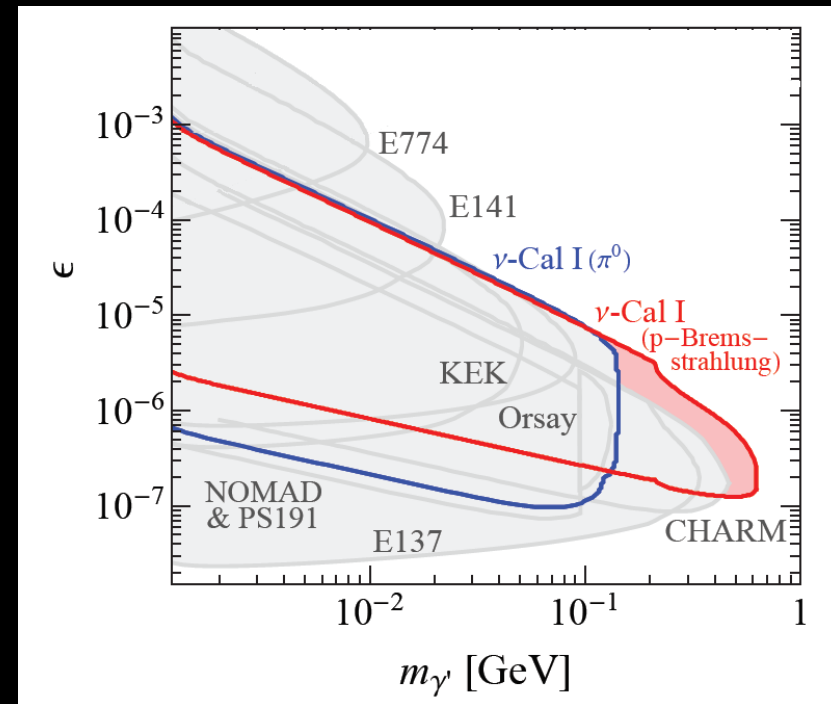


# Beam dump experiments



- Beam produces hadronic and/or EM shower
- Secondary particles emit  $A'$
- Dark photons can decay near the detector, and be reconstructed as narrow resonances
- Original experiments looking for  $\nu$ , axions, light Higgs, ... **have been reinterpreted as constraints on dark photon production**
- Sensitive to low mixing values at large masses, complementary to other approaches

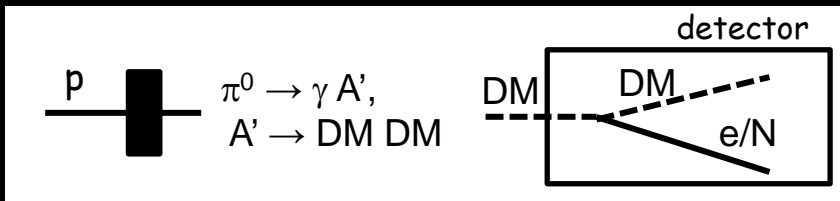
Blumlein & Brunner, arXiv: 1311.3870



# Beam dump and invisible $A'$ decays

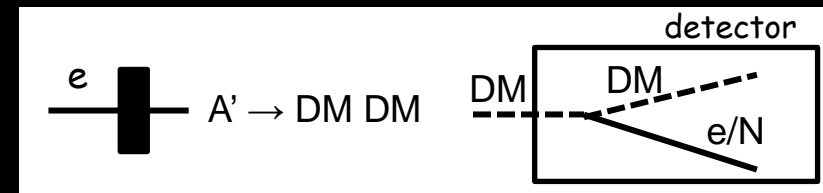
## Proton-beam

- Invisible DM produced in pion decay
- Neutrino factory ideal for probing this scenario (MicroBoone, Nova, LBNE,...)

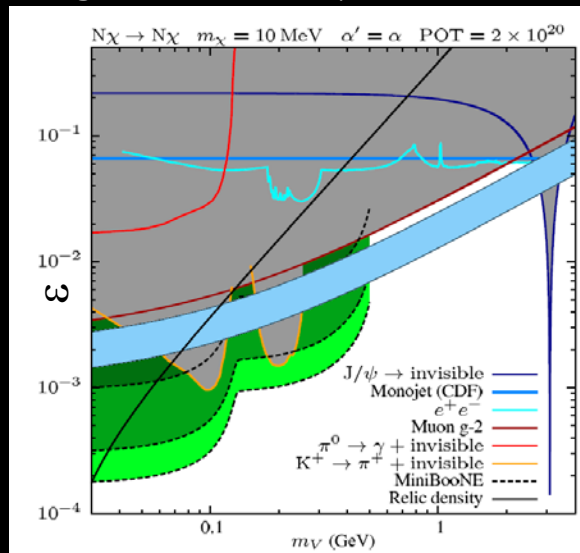


## Electron-beam

- Low background
- Small mass detector
- Favorable kinematics

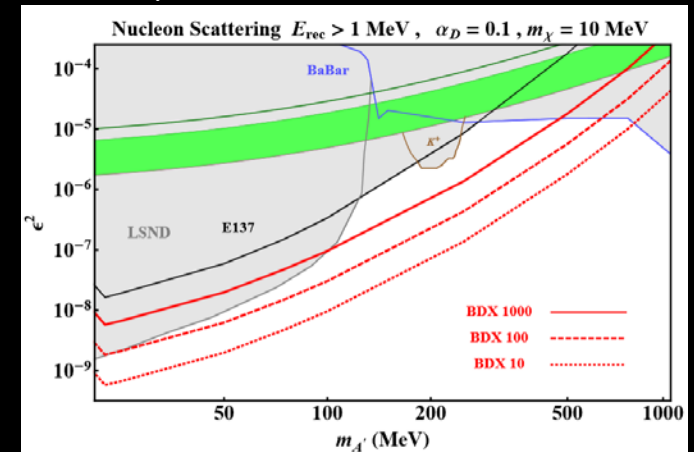


## e.g. MiniBoone expected reach



Aguilar-Arevalo *et al.*, arXiv:1211.2258

## BDX experiment (arXiv:1406.3028)

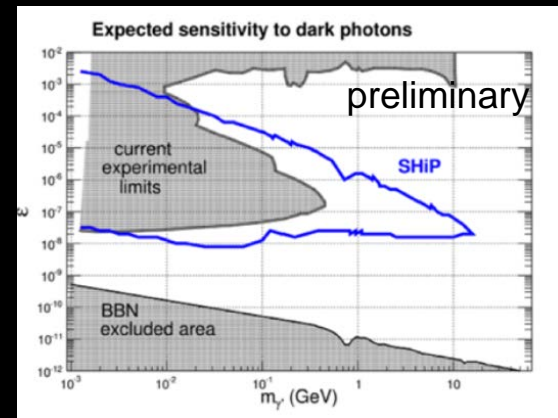
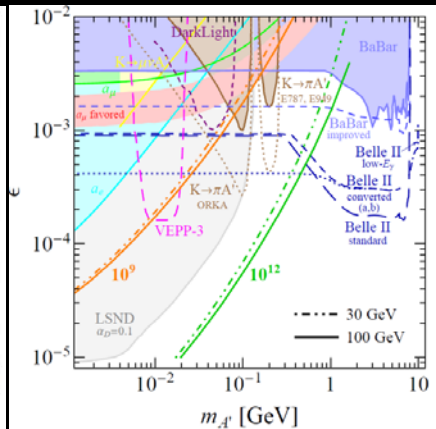
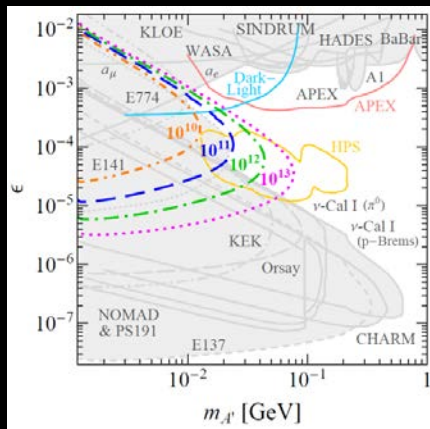
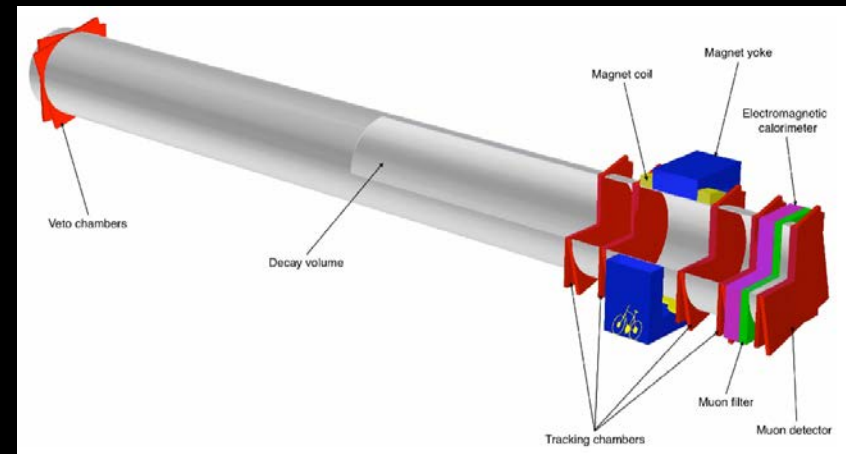
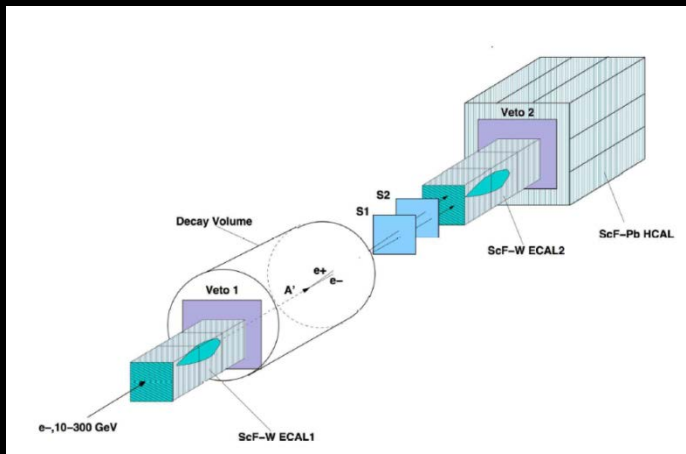


Even better @ arXiv:1411.1404

# Beam dump experiment proposal at CERN

Using the CERN SPS e- beam  
(arXiv:1312.3309)

The SHIP proposal at CERN  
(<http://ship.web.cern.ch/ship/>)

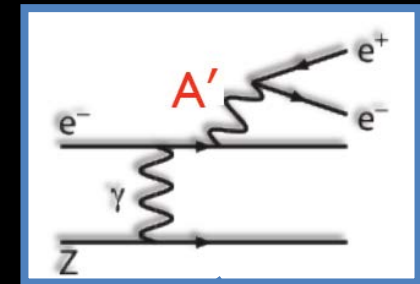


W. Bonivento

# Fixed target experiments

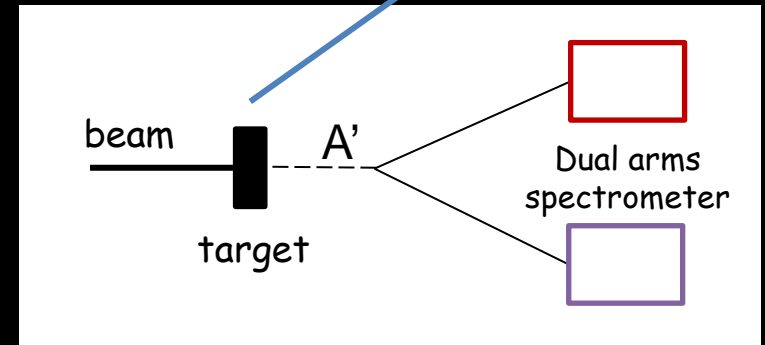
## Fixed target experiments

- Electron beam on fixed target radiates  $A'$
- Decay product detected by dual arm spectrometer



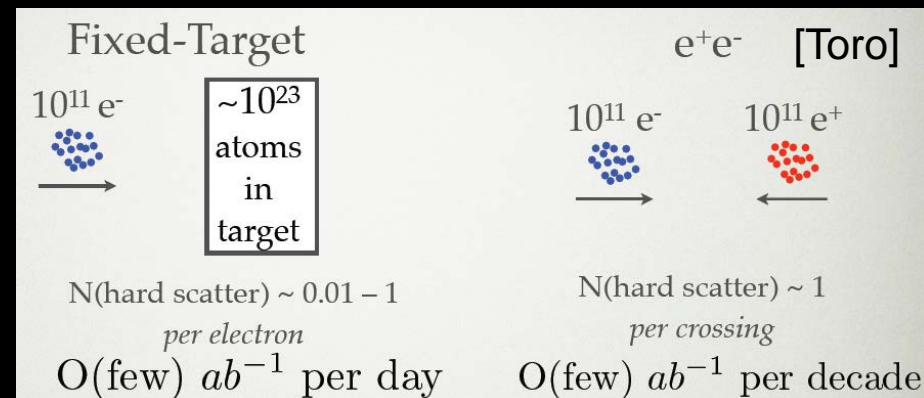
## Fixed target have huge luminosity

- Much denser target
- Cross-section  $\propto Z^2$  and  $1/m^2$



## But small signal and large background

- Small bump on top of background
- Displaced vertices boosts sensitivity





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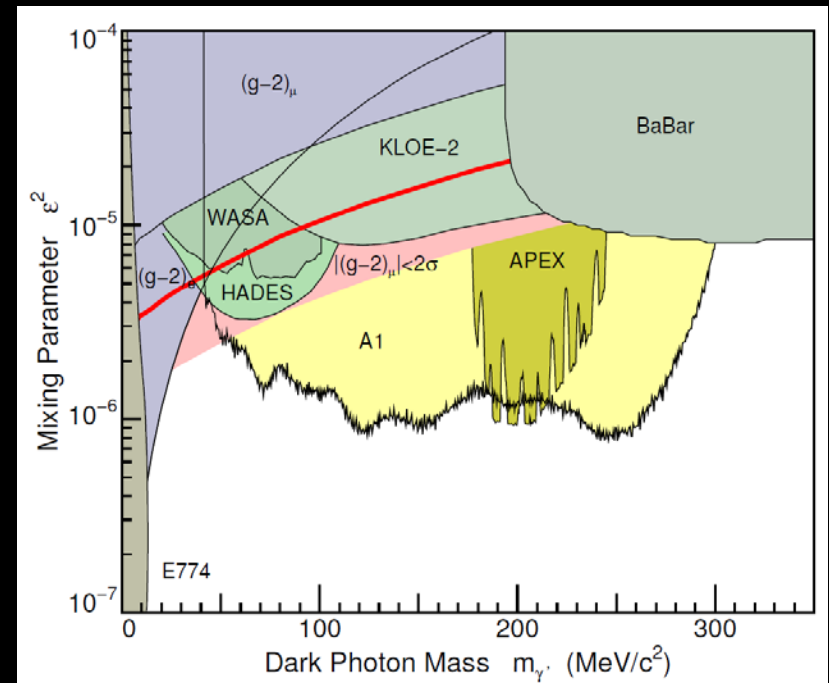
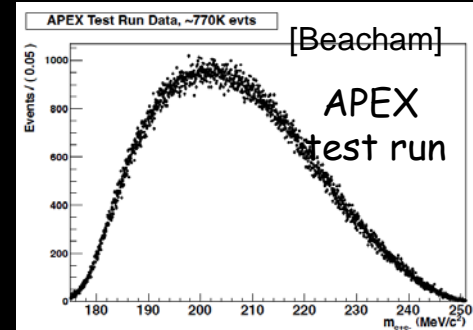
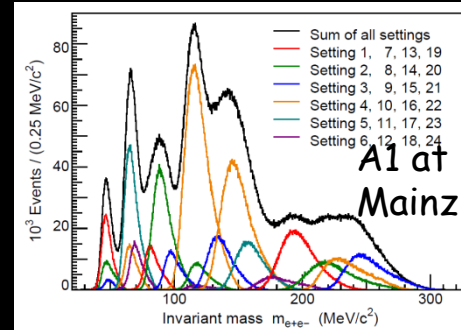
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## Recent results

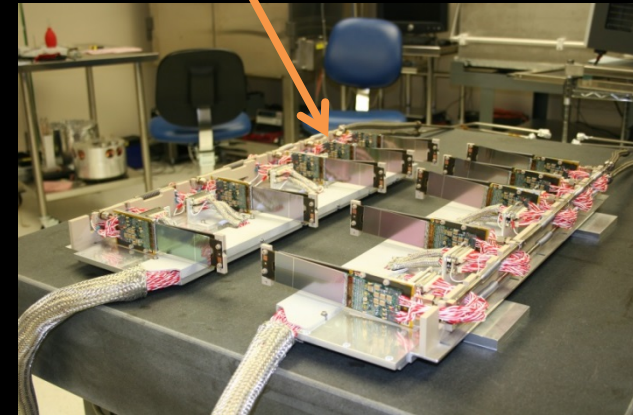
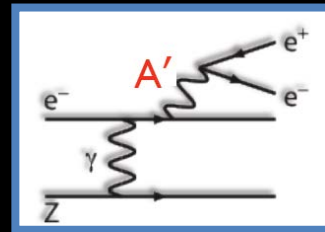
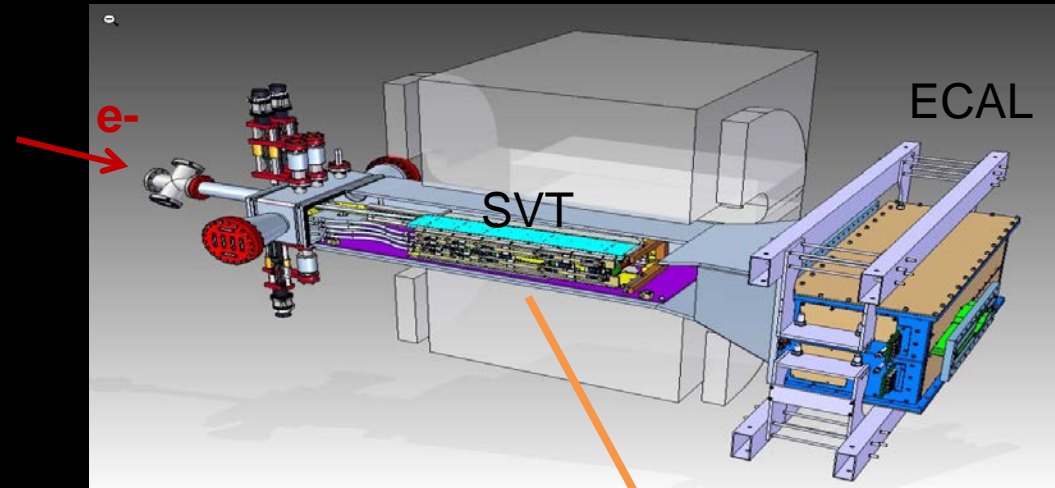
- A1 at Mainz: 850 MeV  $e^-$  beam
- APEX at Jlab: 6 GeV  $e^-$  beam



Expect to improve  
sensitivity in near future

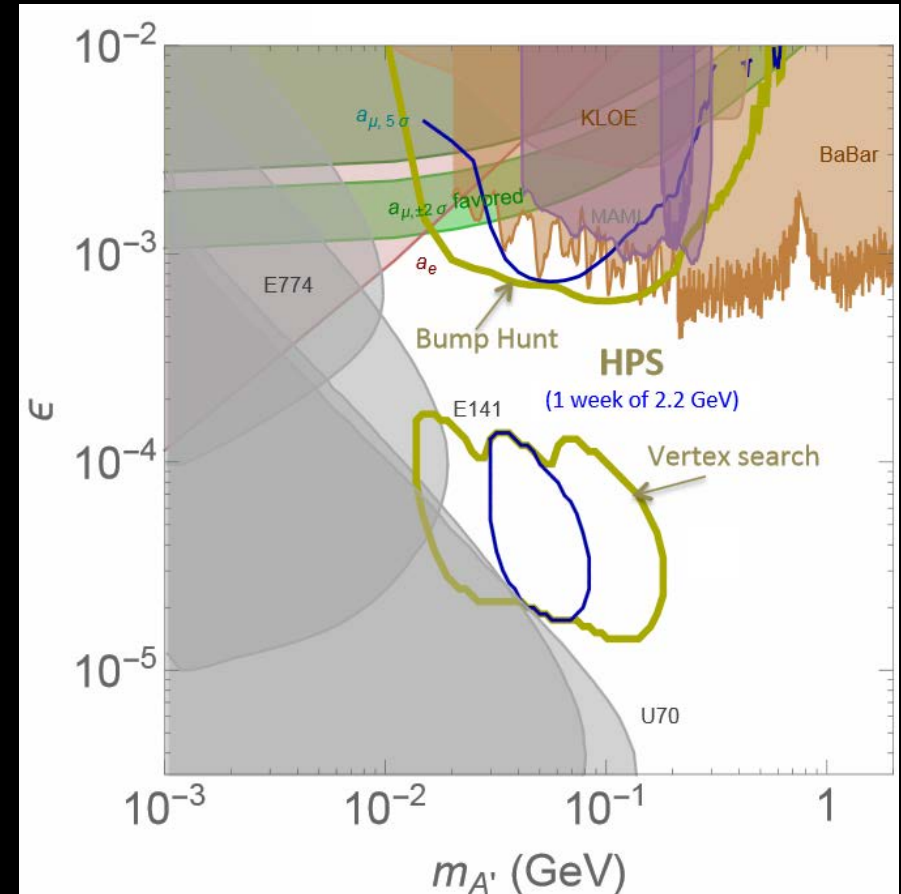
## Heavy Photon Search experiment at JLab

- Large forward-acceptance spectrometer
- Electron beam hits a target, radiates dark photons which converts into an  $e^+e^-$  pair.
- Silicon vertex tracker to measure  $e^+e^-$  mass and vertex position
- $\text{PbWO}_4$  crystal calorimeter to identify  $e^+e^-$  and trigger
- High rate trigger and DAQ
- Search for prompt (bump hunt) and displaced  $A'$  decays (vertex)
- Scheduled to be running spring 2015



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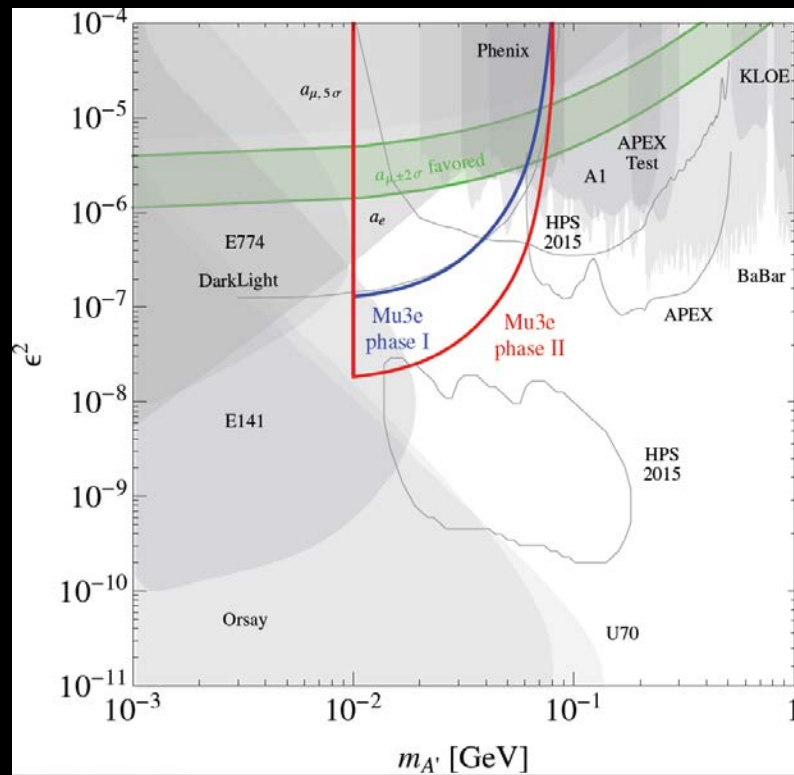
1 week 1.1 GeV  
 +1 weeks 2.2 GeV  
 +2 weeks 4.4 GeV

\*<https://confluence.slac.stanford.edu/display/hpsg/HPS+Proposals>

# The Mu3e experiment

## Mu3e experiment at PSI

- Search for Lepton Flavor Violation (LFV) in muon decay  $\mu \rightarrow eee$  with a sensitivity down to  $10^{-16}$ .
- Low mass silicon vertex detector, high granularity and precision
- The decay  $\mu \rightarrow e\nu\nu\gamma$  has a large branching fraction. This is ideal to...
- ... search for  $\mu \rightarrow e\nu\nu A'$ ,  $A' \rightarrow ee$ . Final state contains also three electrons and missing energy. Main background is  $\mu \rightarrow eee\nu\nu$  (BF =  $3.6 \times 10^{-5}$ ).
- Search for narrow peak over smooth background (prompt decays)
- Experiment should start in 2015 with lower beam intensity, 2018 for upgraded beam.

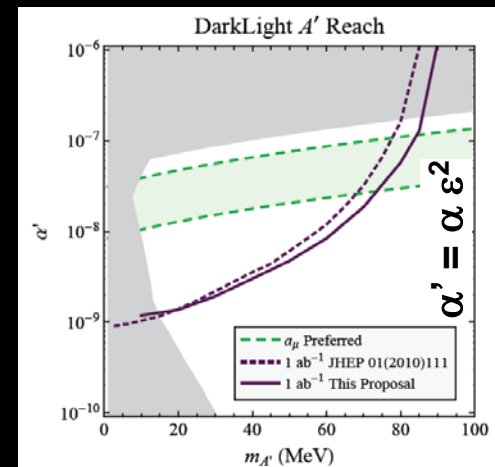
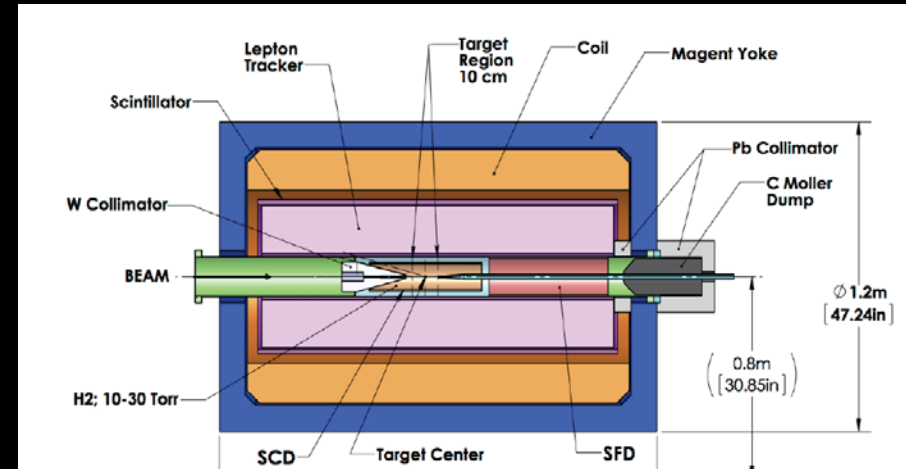


BE, R. Essig, Y. Zhong, arXiv: 1411.1770

Can probe low mass region

## DarkLight\* at Jlab

- Compact  $4\pi$  detector
- Electron beam (100 MeV) on gaseous hydrogen target
- Measure the full reaction  $e^- p \rightarrow e^- p A'$
- Measure visible and invisible  $A'$  decays for  $m_{A'} < 90$  MeV
- Test run at Jlab FEL to demonstrate concept
- Expect to run in 2016



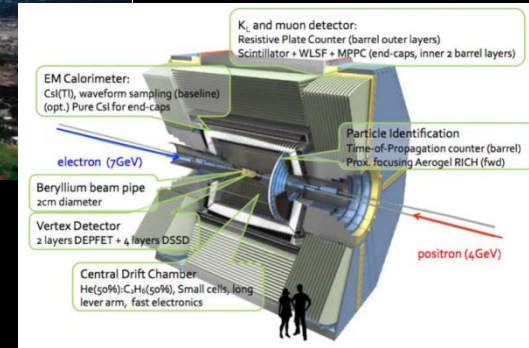
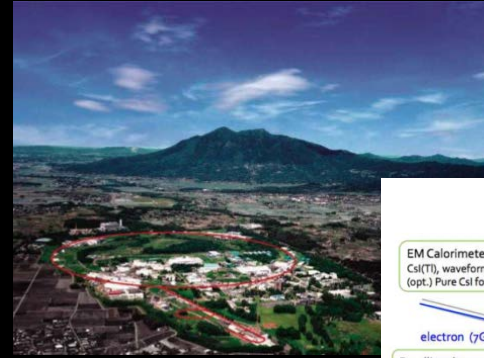
\*DarkLight = Detecting  $A$  Resonance Kinematically with  $e$ lectrons Incident on a Gaseous Hydrogen Target  
[http://dmtpc.mit.edu/DarkLight/DarkLightProposal\\_PAC39.pdf](http://dmtpc.mit.edu/DarkLight/DarkLightProposal_PAC39.pdf)



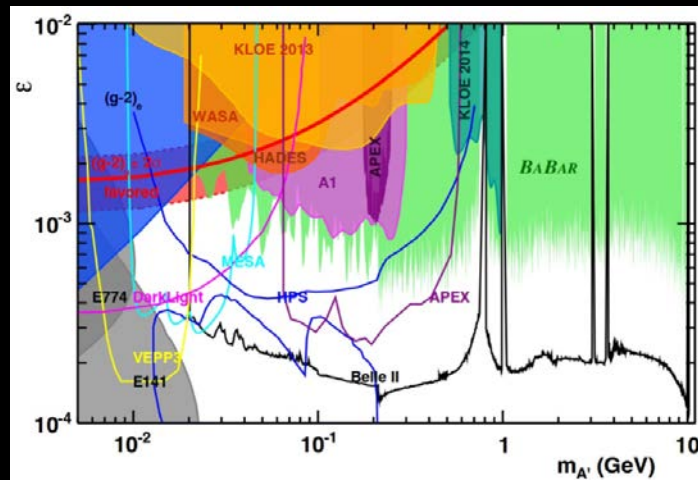
# The Belle II experiment

## Belle II experiment

- High luminosity  $e^+e^-$  collider at the Y(4S) center-of-mass energy at KEKB (Japan)
- Collect 100x more data than BABAR
- Will start taking physics data late 2016
- Expected to probe  $A' \rightarrow \text{visible}$  and  $A' \rightarrow \text{invisible}$  decays and considerably improve current constraints.

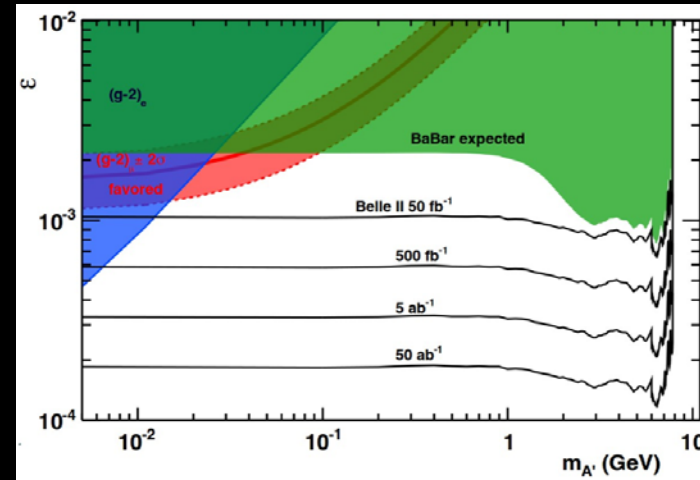


### $A' \rightarrow \text{visible}$



C. Hearty

### $A' \rightarrow \text{invisible}$

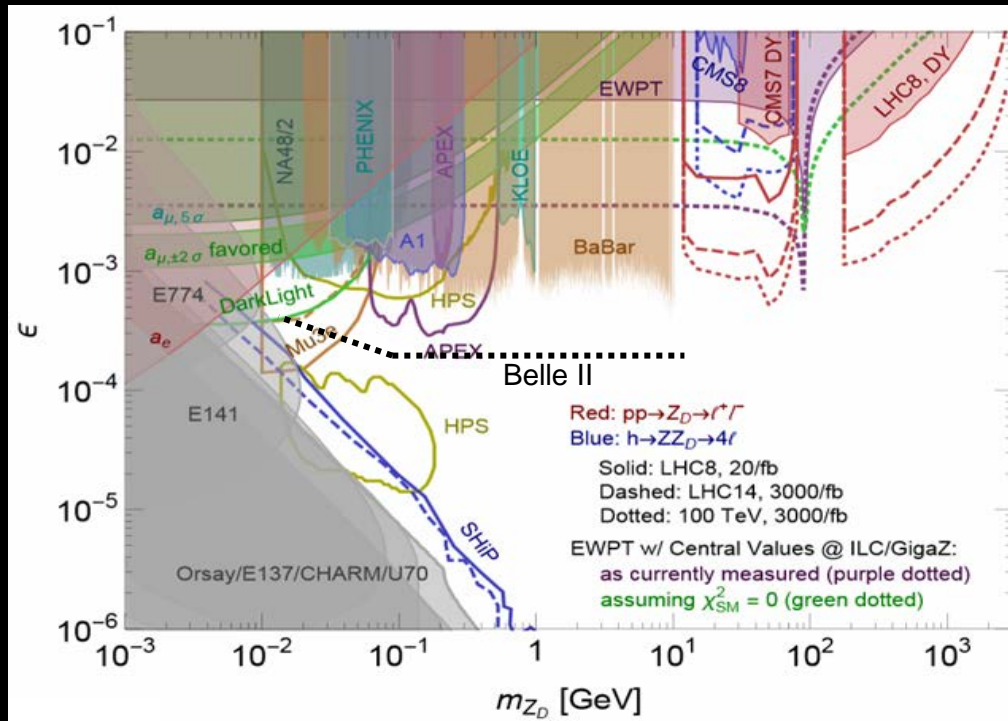


C. Hearty

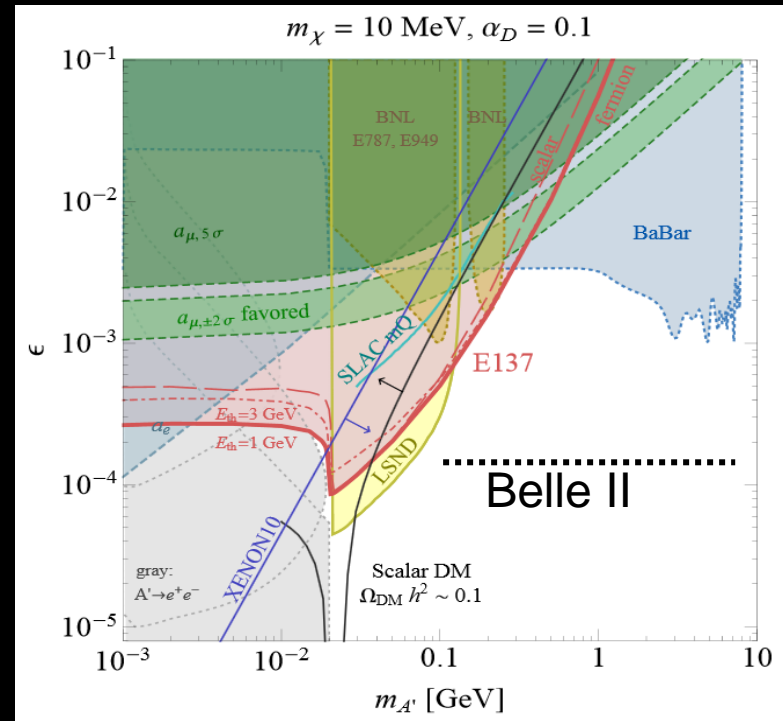


# Summary plots

$A' \rightarrow \text{visible}$



$A' \rightarrow \text{invisible}$



Future experiments will probe a large fraction of the low mass region, but there is still a lot of ground to cover!

People often ask me what are the practical implications of my research.

Living in LA, it seems natural to point them towards entertainment...



# Summary

- There are still many intriguing possibilities to explore at the GeV scale, and dark forces open a new window on physics far beyond the SM.
- A fraction of the dark photon parameter space has already been probed by current experiments:  $g-2$ , fixed target, beam dump,  $e^+e^-$  colliders,...
- But there is still a lot of uncharted territory!
- New experiments at existing facilities will further explore this parameter space, hopefully resulting in a game-changing discovery.

*And remember, nothing bad ever happens when you work with dark force!*