

UCI School of Physical Sciences

Imaging Planet Formation with Professor Steph Sallum

Welcome, we will begin shortly

*For questions, please utilize the Q&A feature at the bottom of
your screen*

Text PSBLS to 41444 to give!

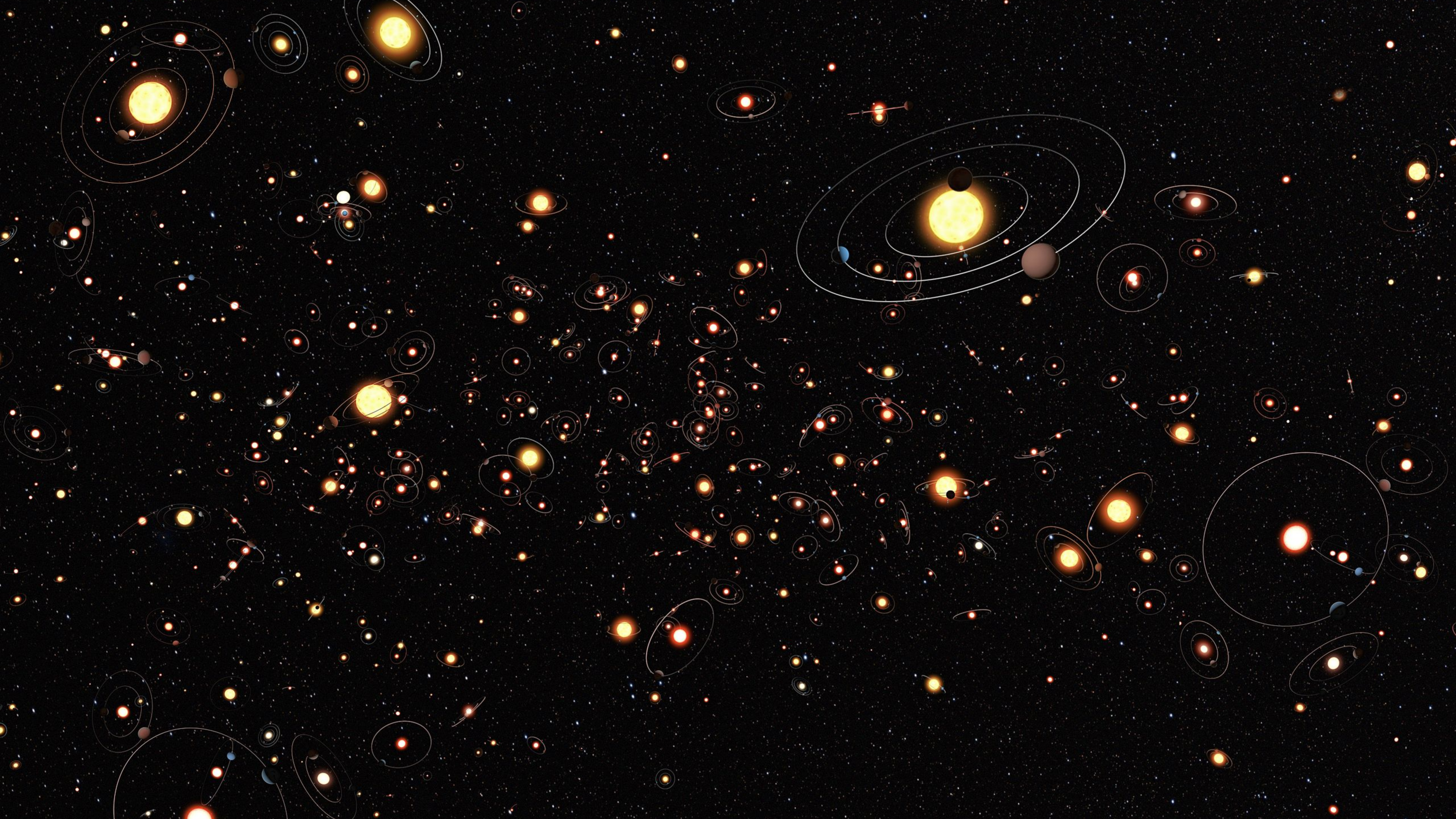
Imaging Planet Formation

UCI SoPS Virtual Lecture Series

Prof. Steph Sallum

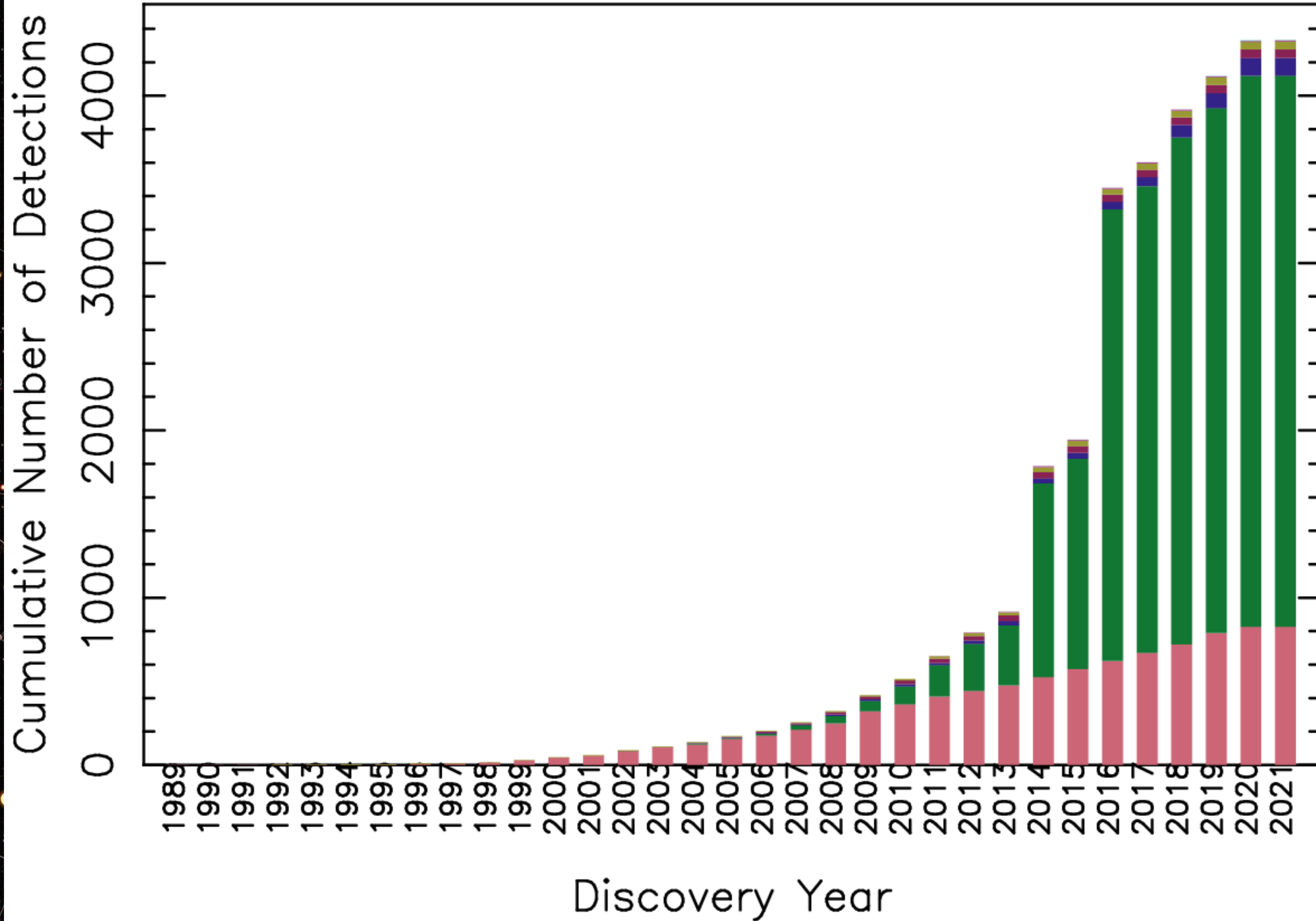
Why study planet formation?





Total Number of Discovered Exoplanets

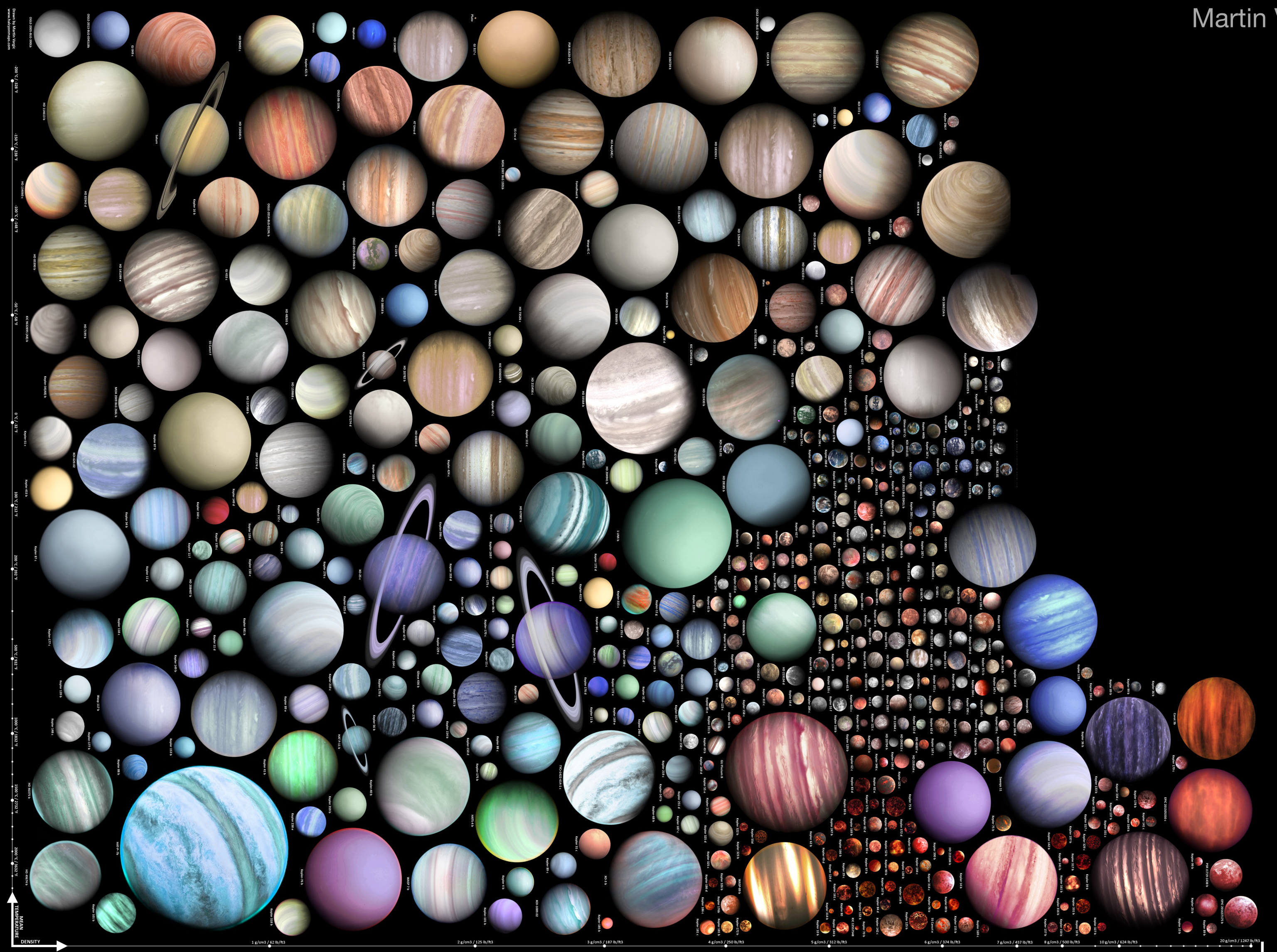
21 Jan 2021
exoplanetarchive.ipac.caltech.edu



Exoplanet Diversity



-330 F
-60 F
212 F
2700 F
3600 F



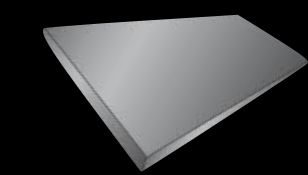
1 g/cm³



2.75 g/cm³



8 g/cm³

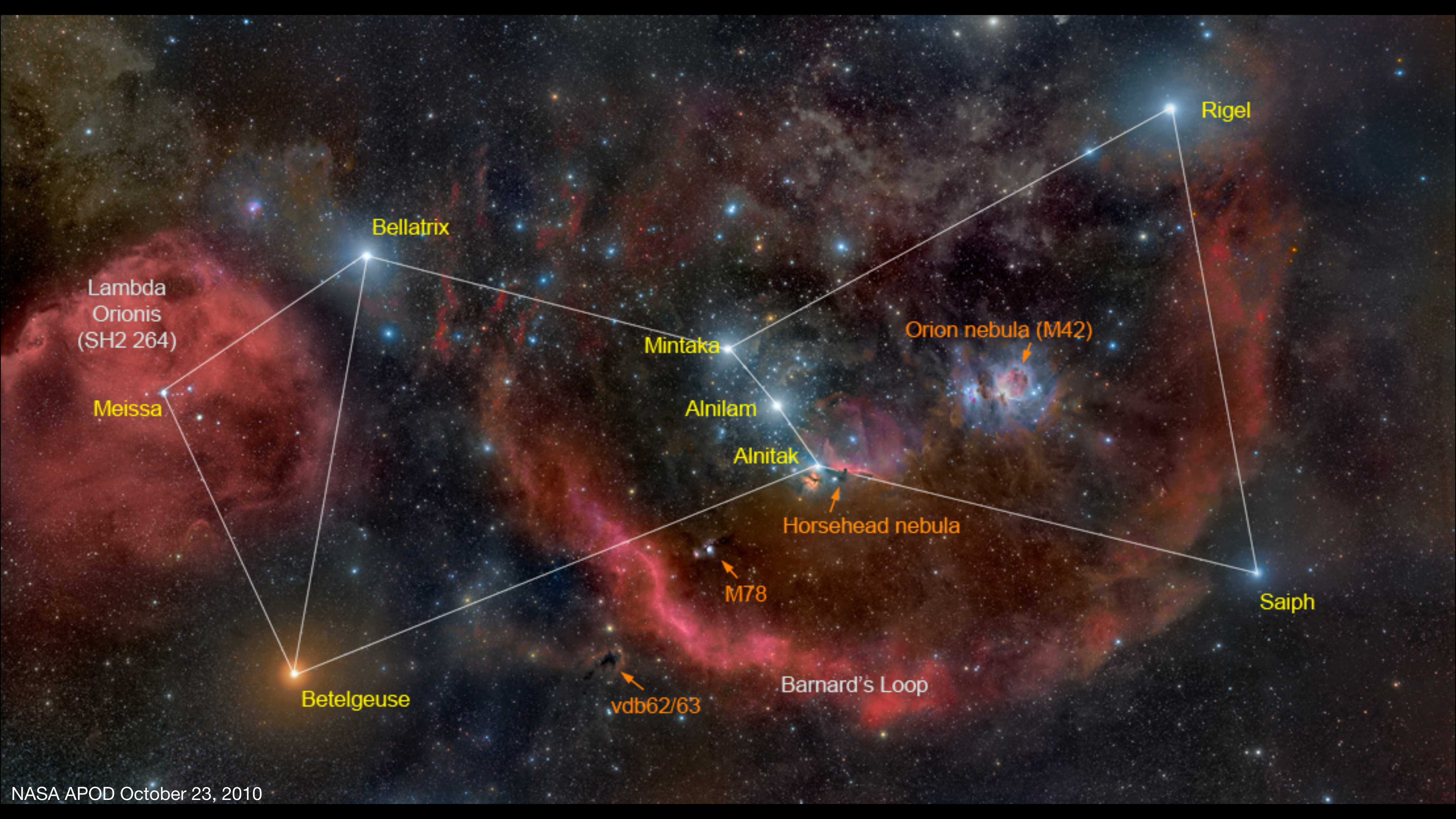


20 g/cm³



Planet Formation





Rigel

Bellatrix

Lambda
Orionis
(SH2 264)

Meissa

Mintaka

Orion nebula (M42)

Alnilam

Alnitak

Horsehead nebula

Saiph

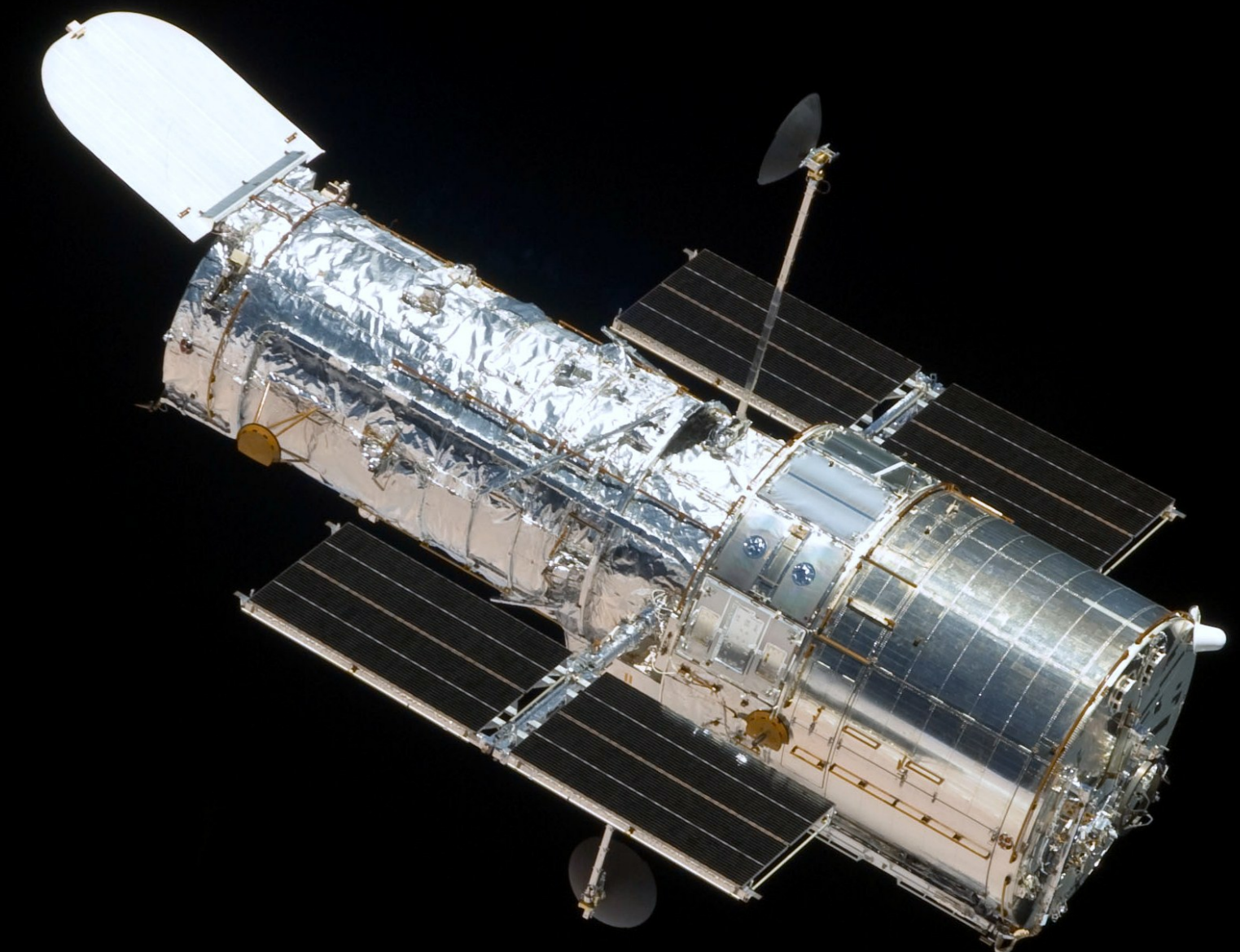
M78

Betelgeuse

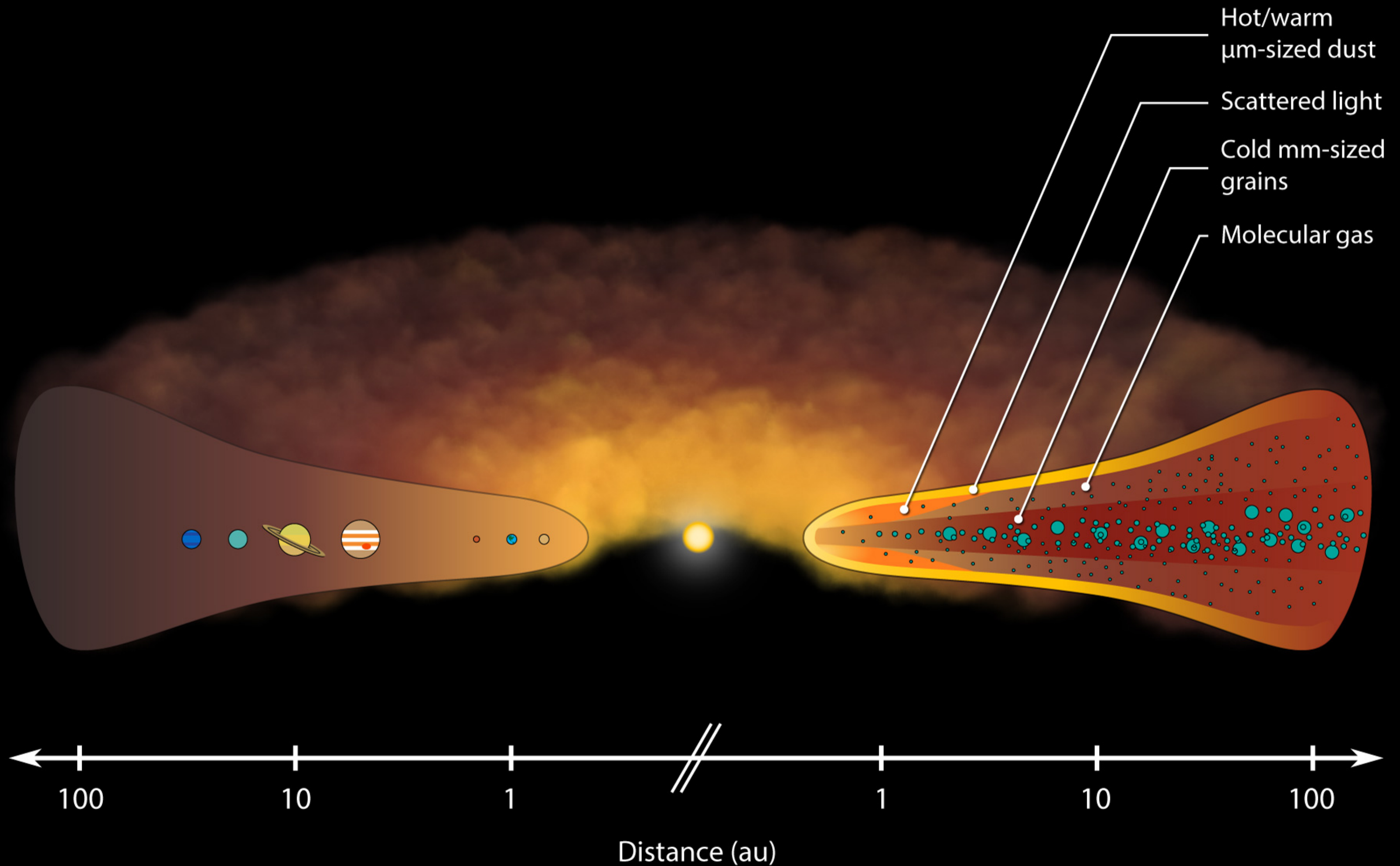
vdb62/63

Barnard's Loop

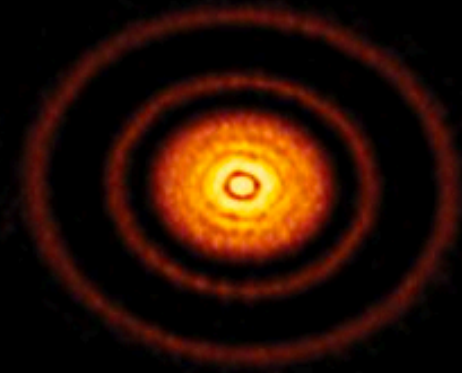
Protoplanetary Disks: The Sites of Planet Formation



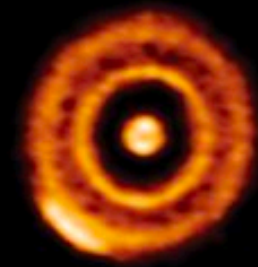
Protoplanetary Disks: The Sites of Planet Formation



Protoplanetary Disks: The Sites of Planet Formation



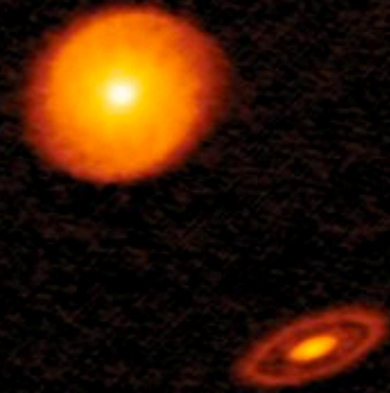
AS 209



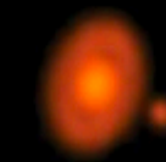
HD 143006



IM Lup



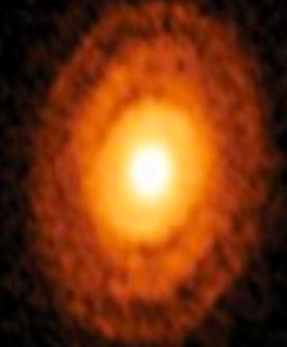
AS 205



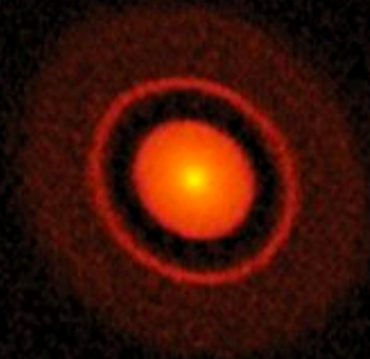
HT Lup



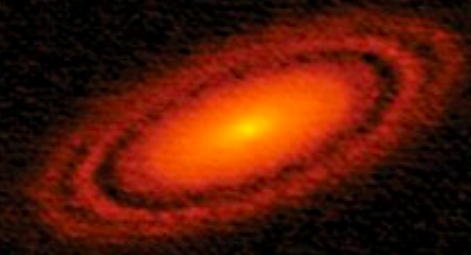
MY Lup



Wa Oph 6



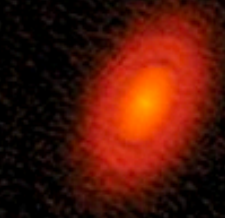
Elias 24



DoAr 25



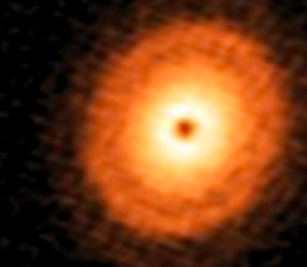
DoAr 33



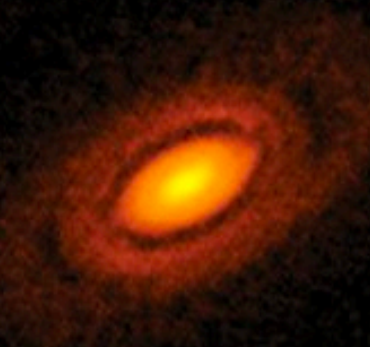
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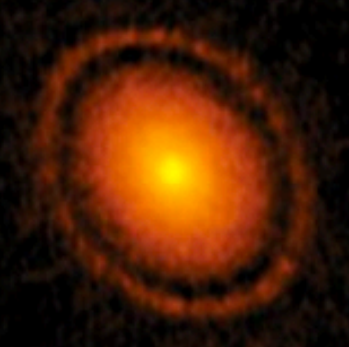
Sz 114



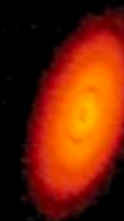
Sz 129



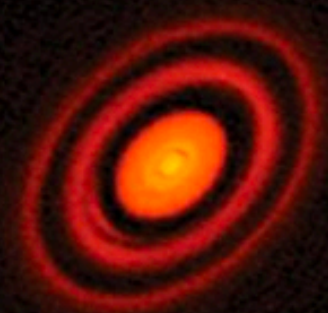
Elias 27



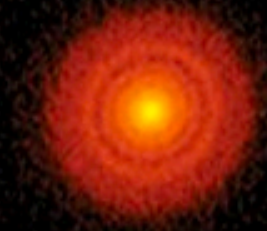
GW Lup



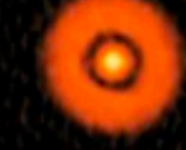
HD 142666



HD 163296



RU Lup

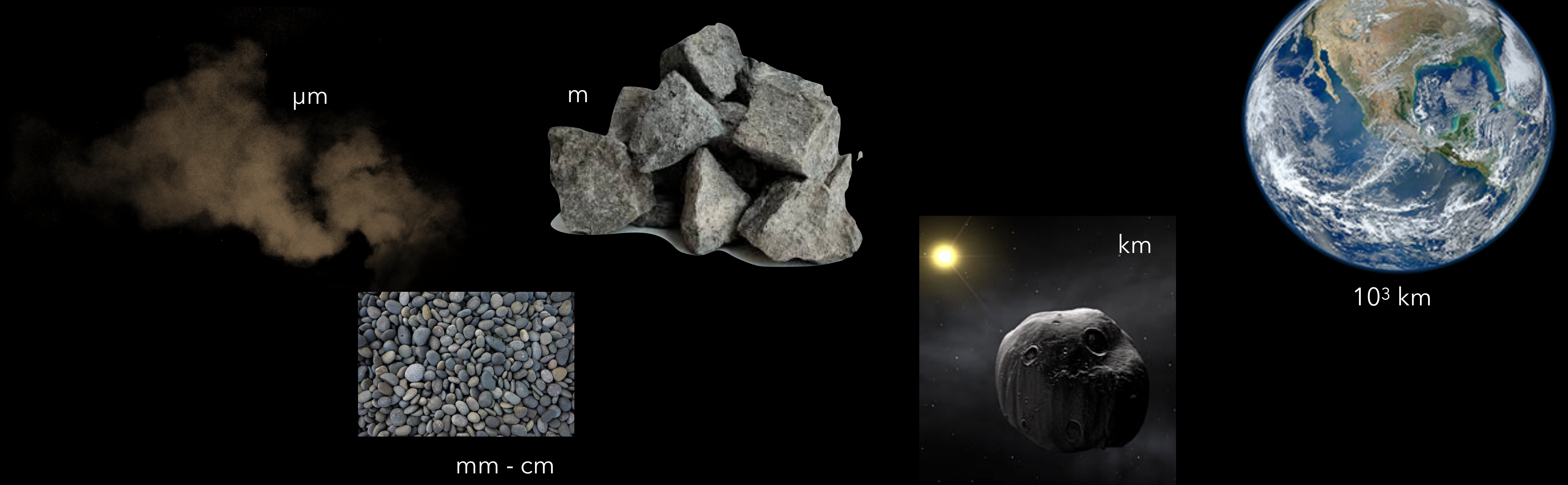


SR4

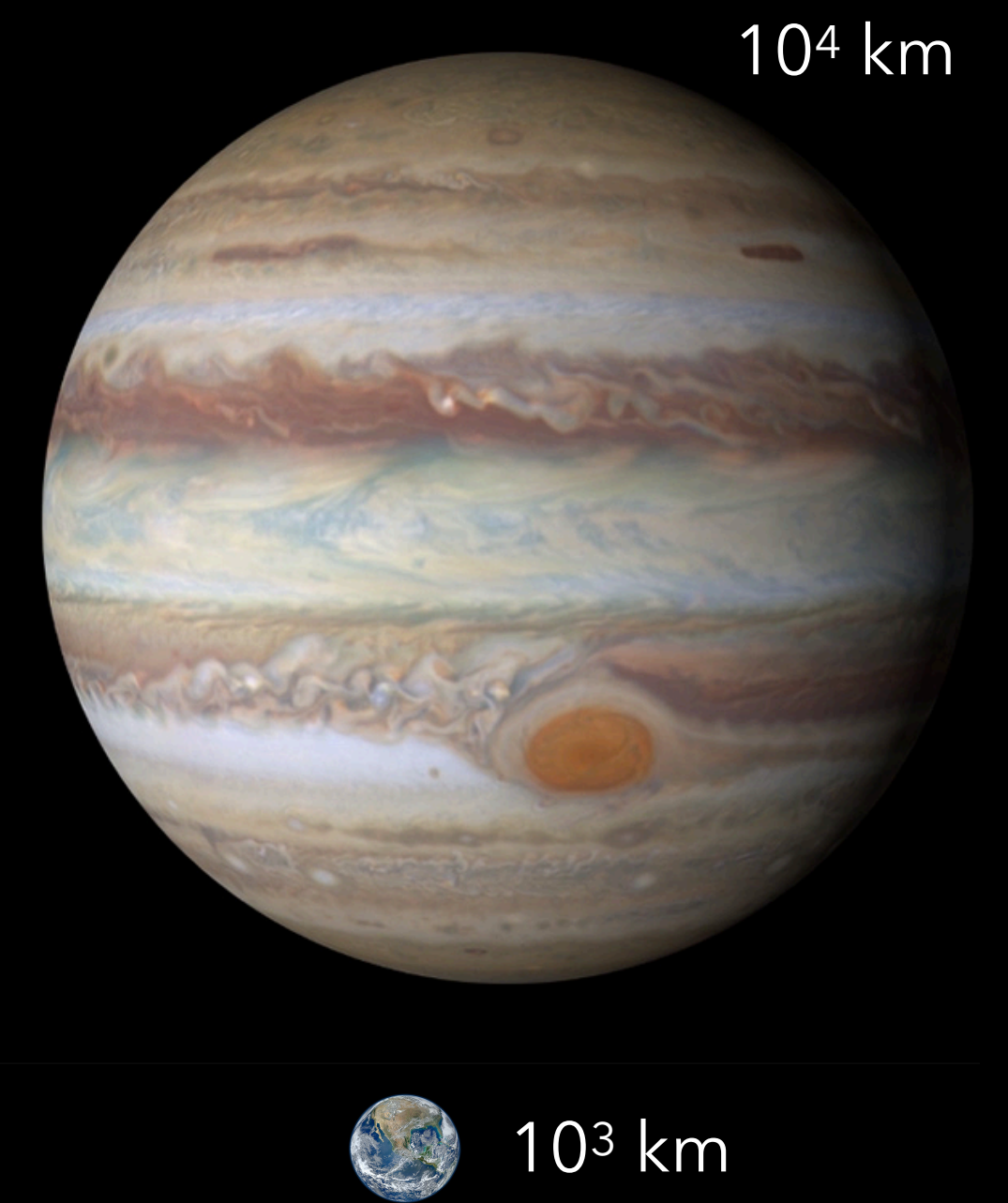
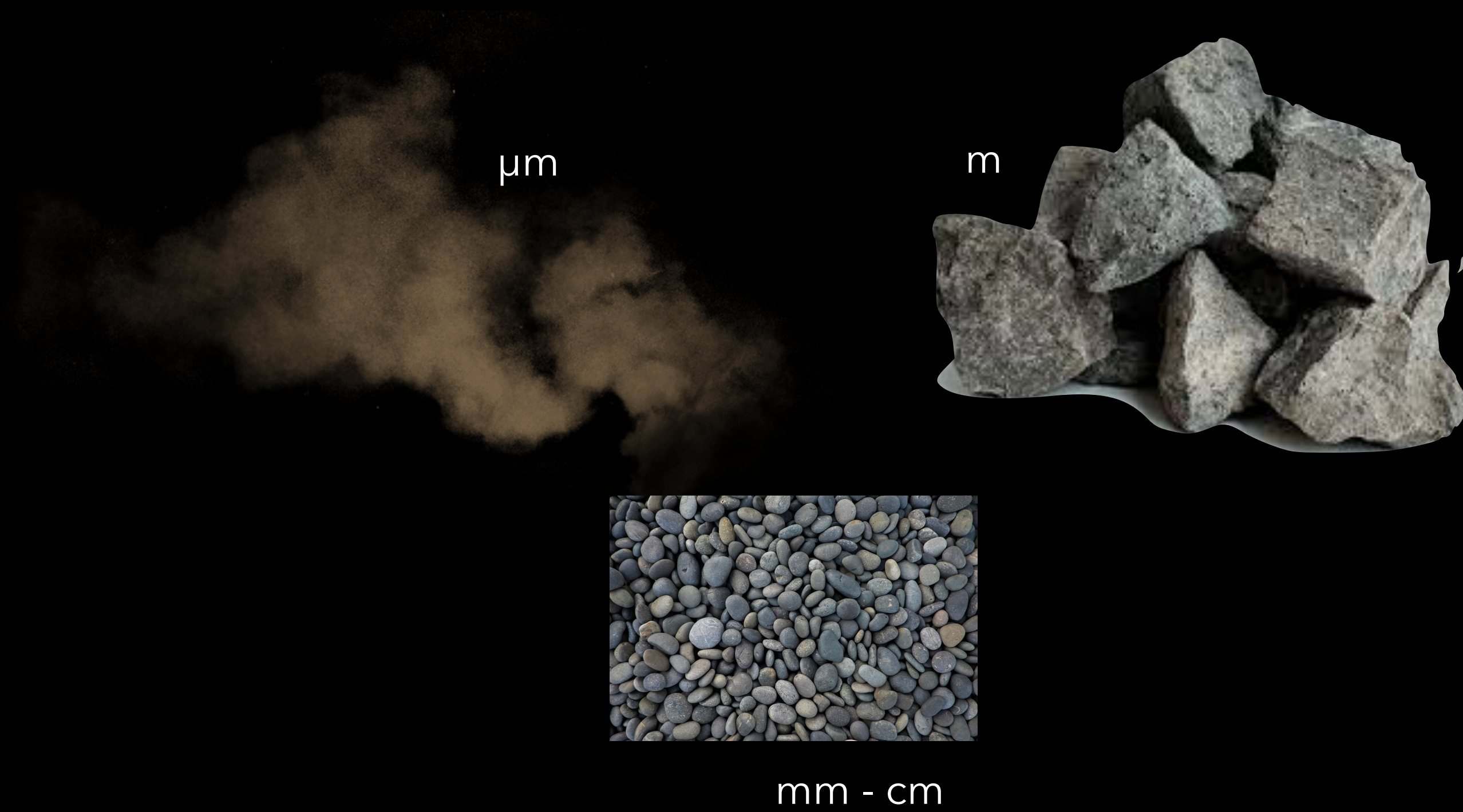


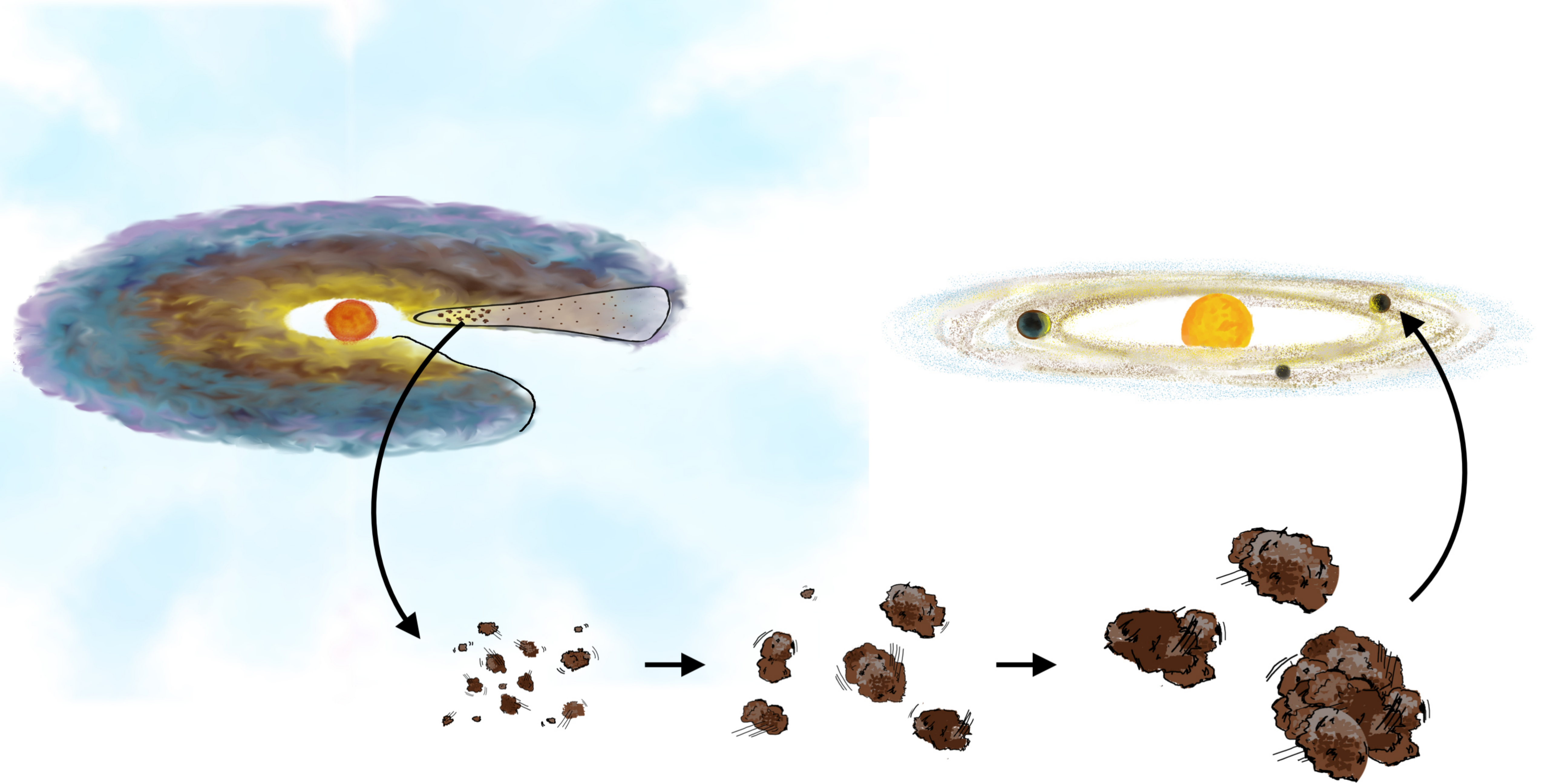
WSB 52

Planet Formation

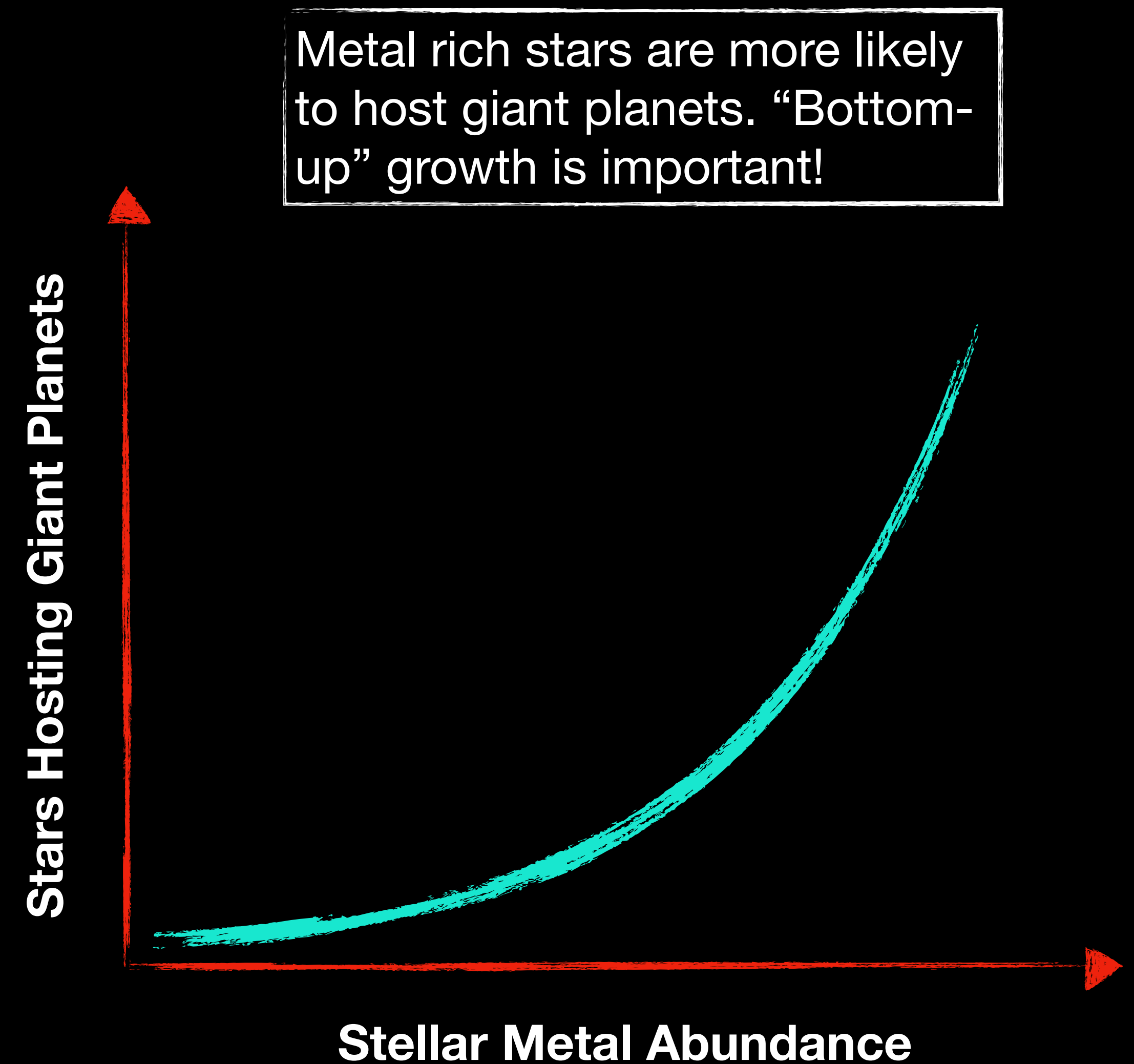
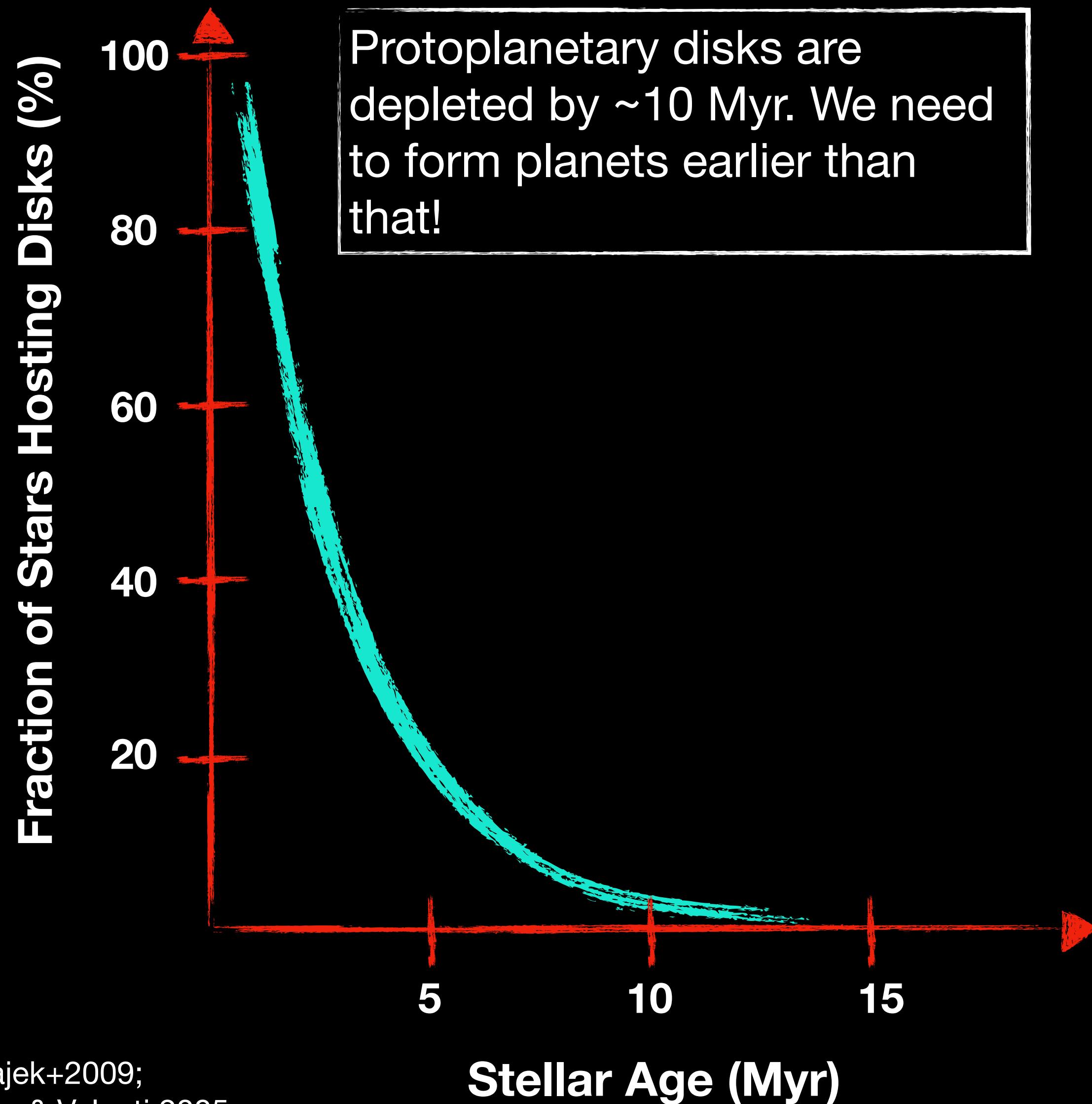


Planet Formation



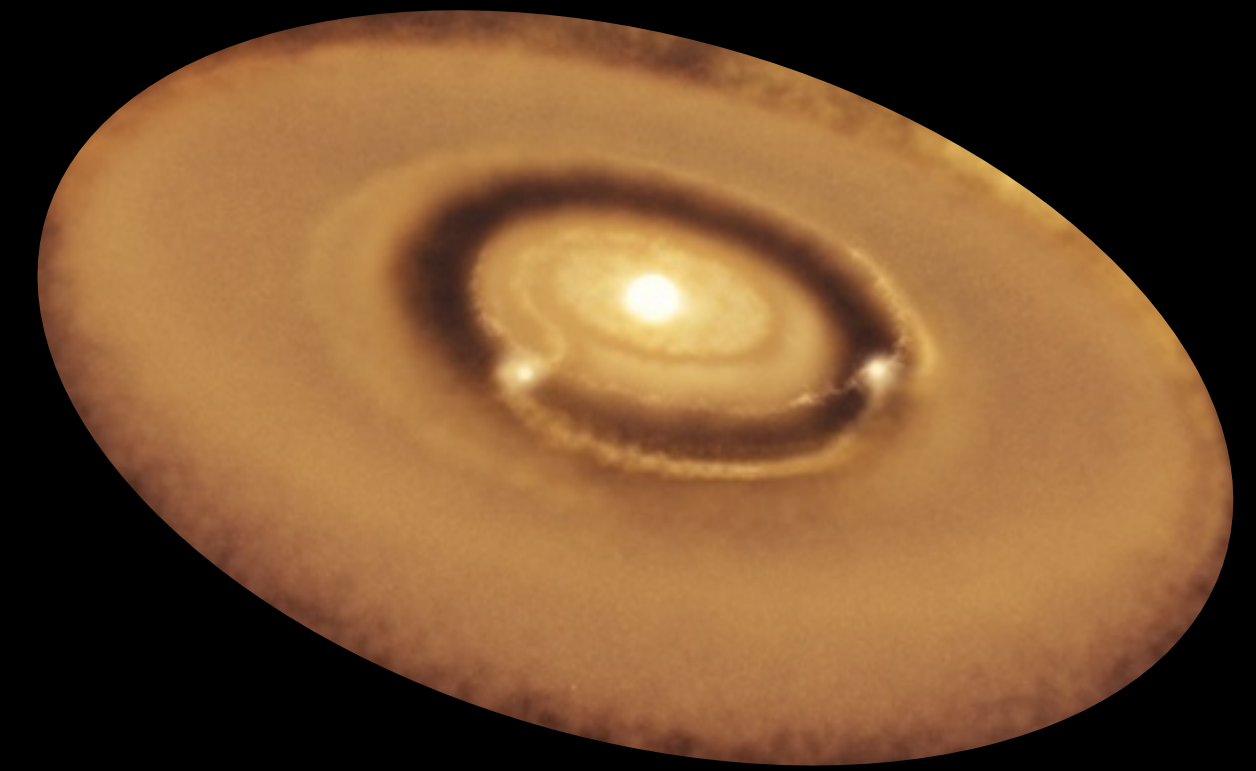


Planet Formation: Some Things We Know

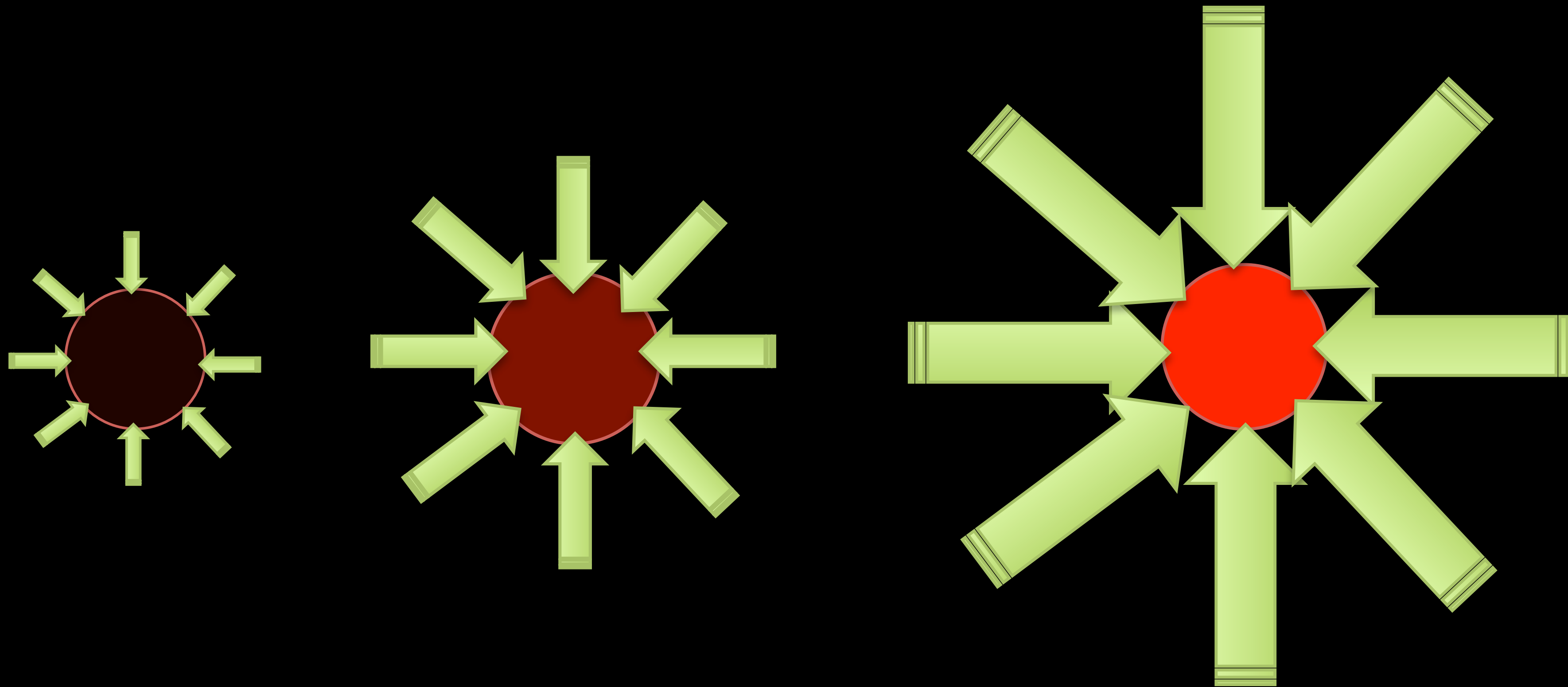


Planet Formation

Some Things We Don't Know



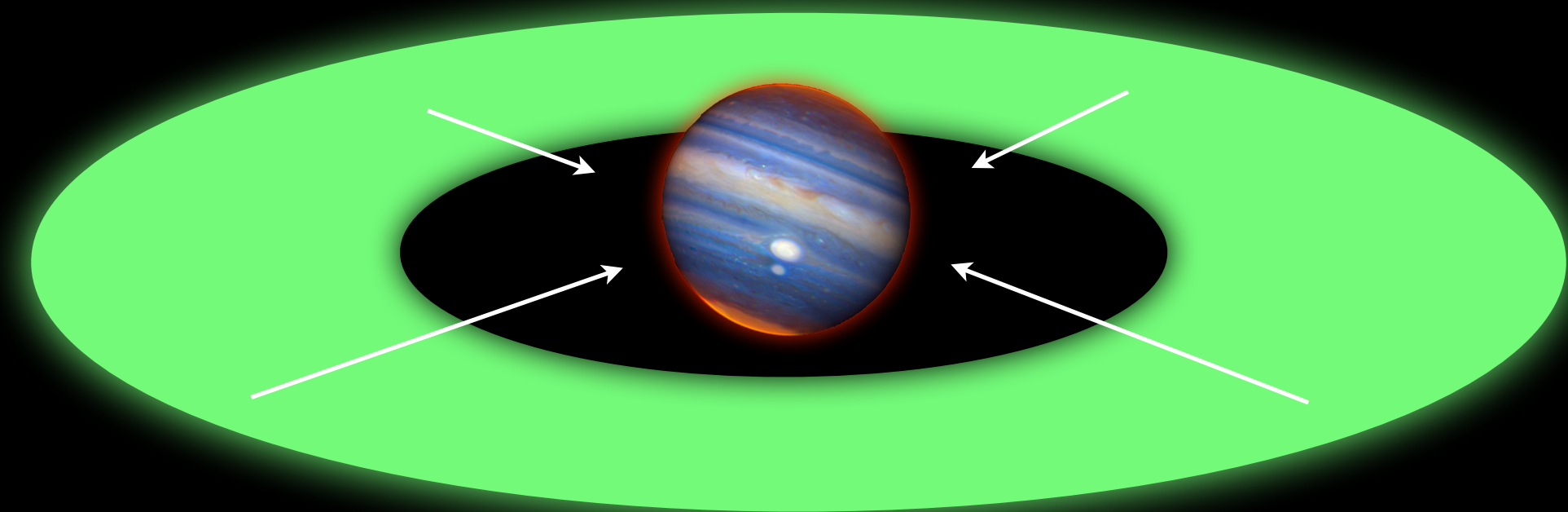
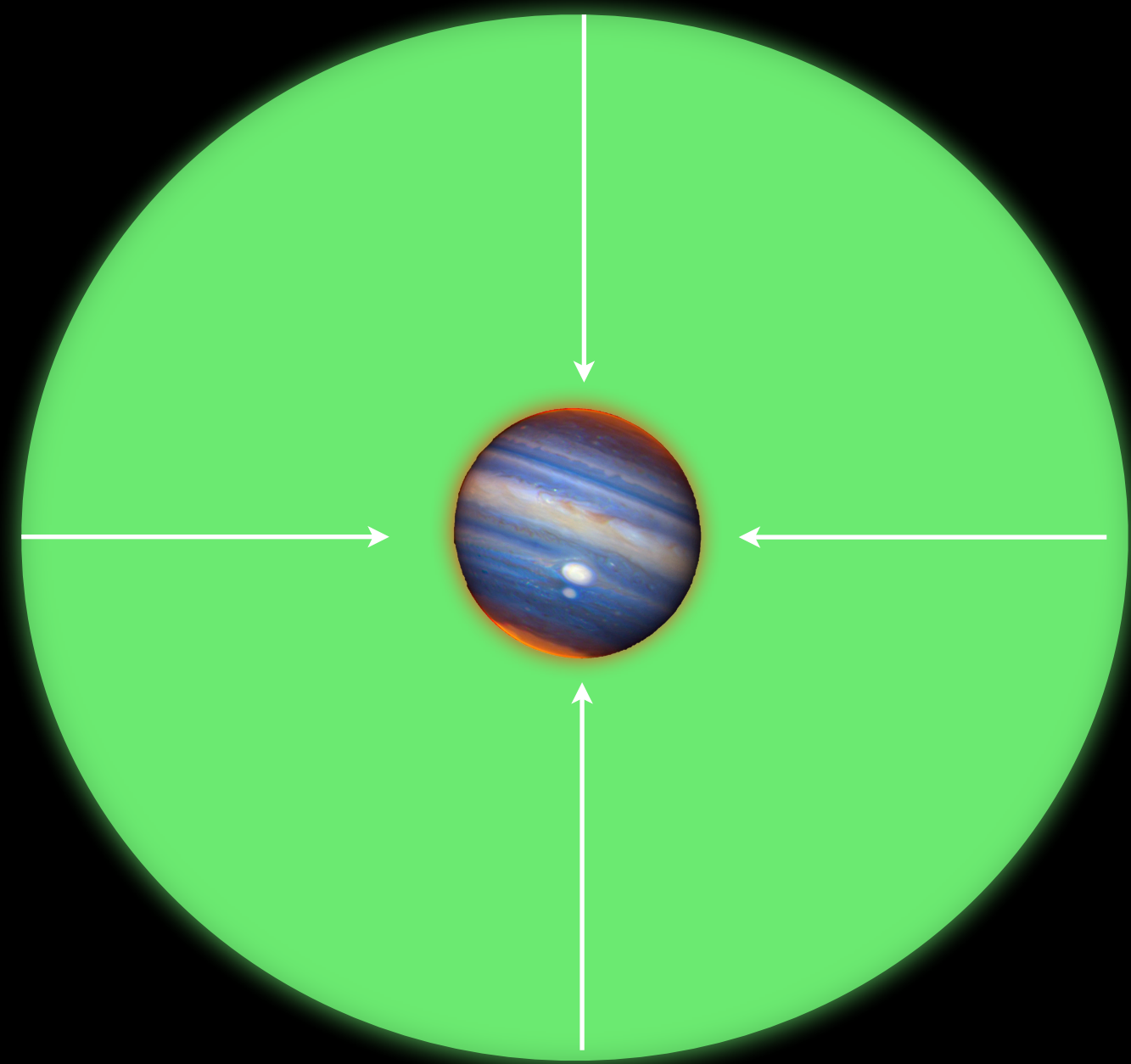
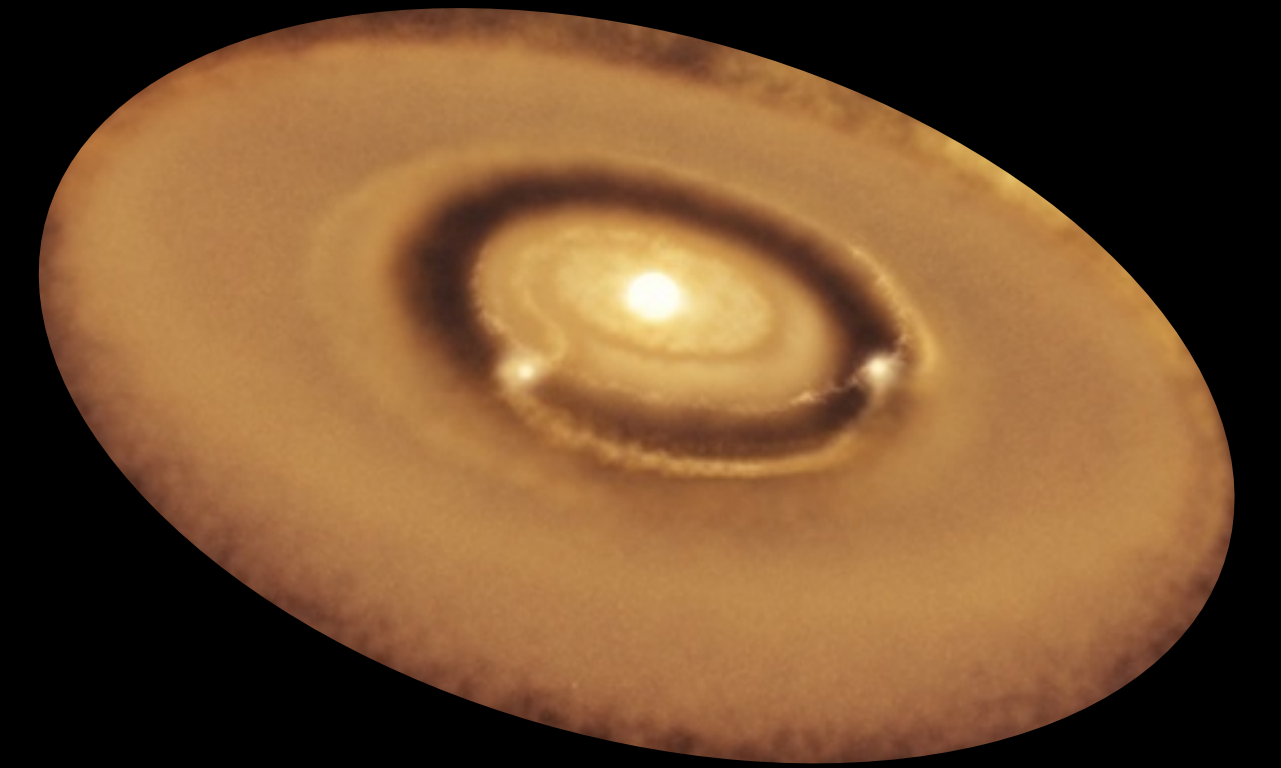
- How fast does material fall onto a forming planet?



Planet Formation

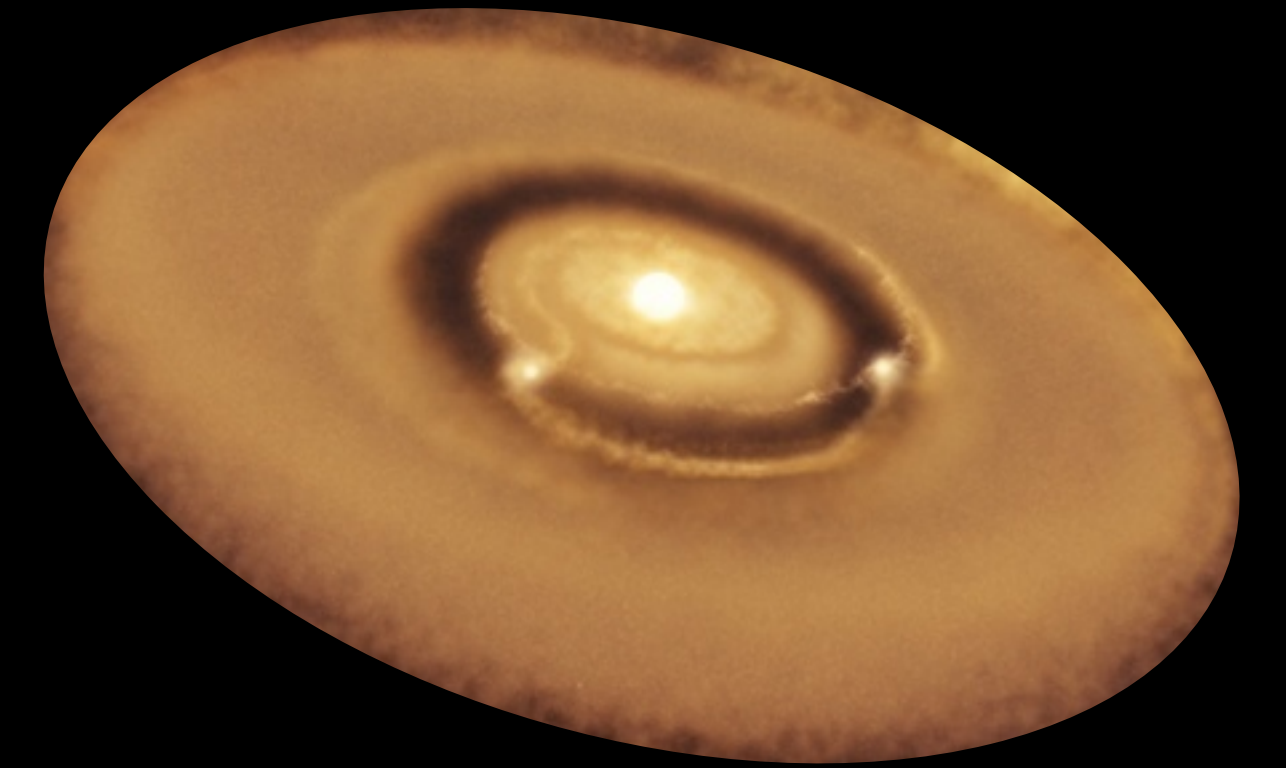
Some Things We Don't Know

- How fast does material fall onto a forming planet?
- What path does the infalling material take?

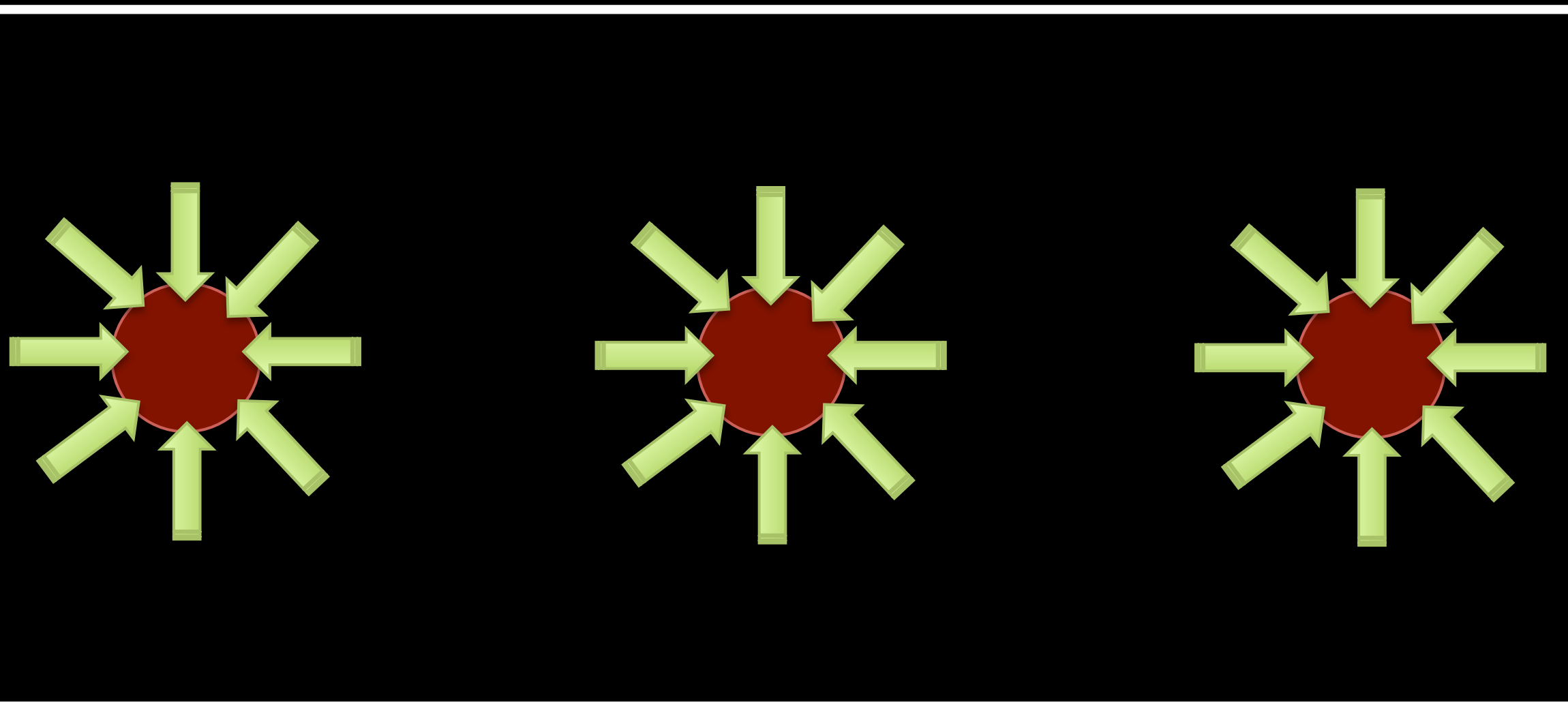


Planet Formation

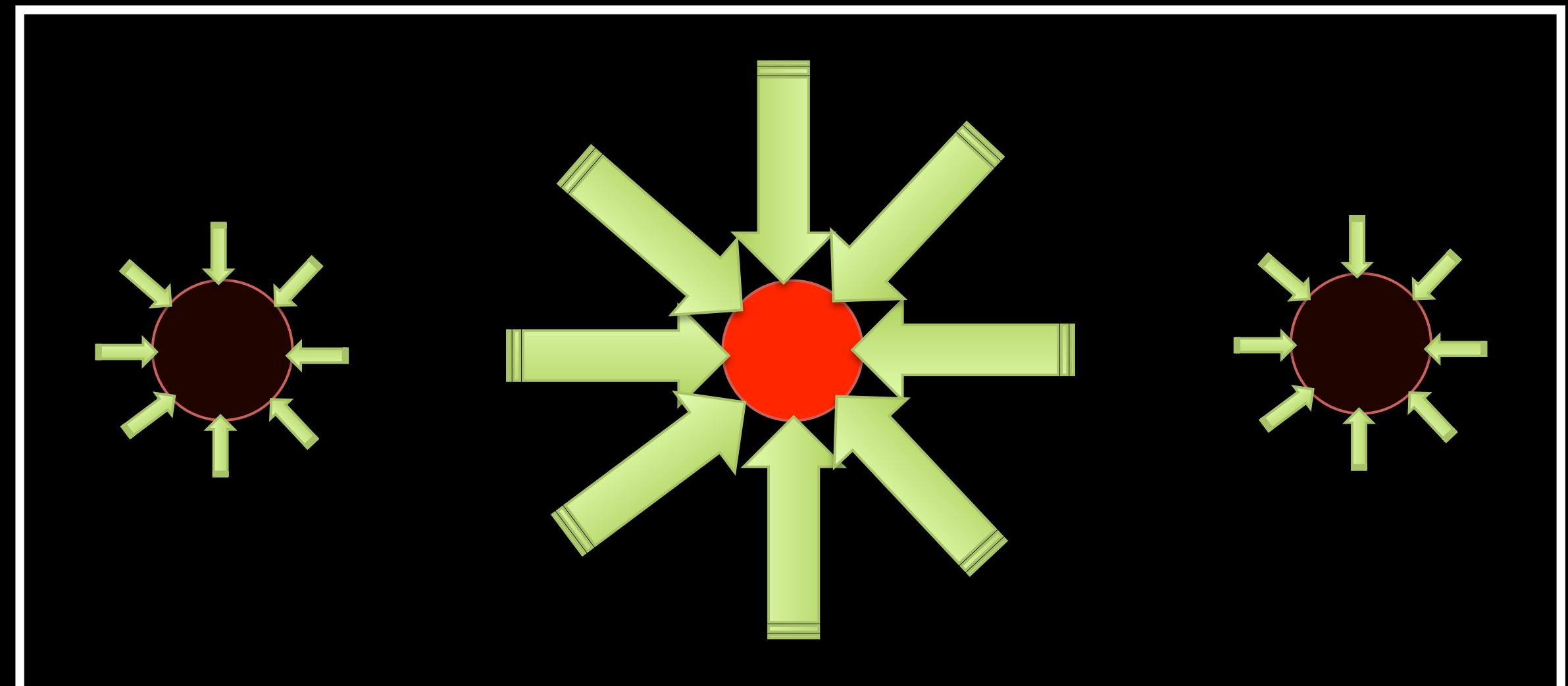
Some Things We Don't Know



- How fast does material fall onto a forming planet?
- What path does the infalling material take?
- How steadily do planets accumulate mass?

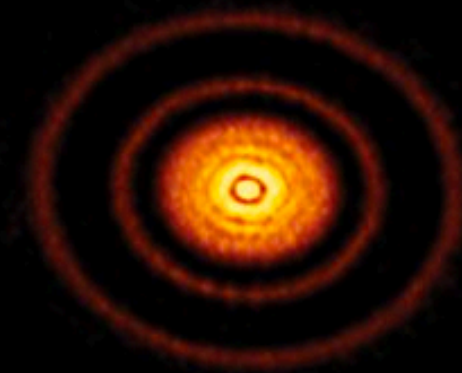


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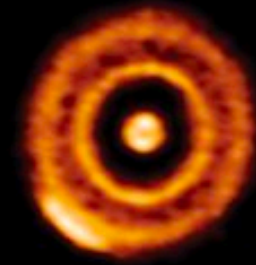


Imaging Planet Formation

Imaging protoplanets: where to search



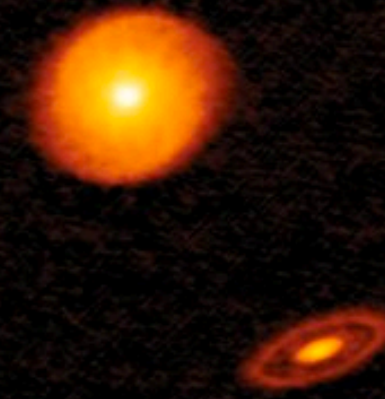
AS 209



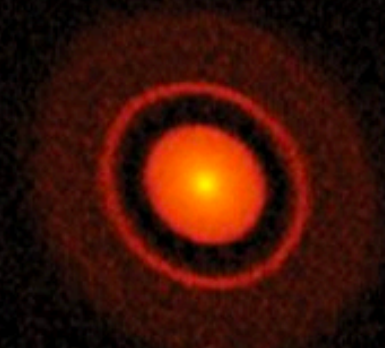
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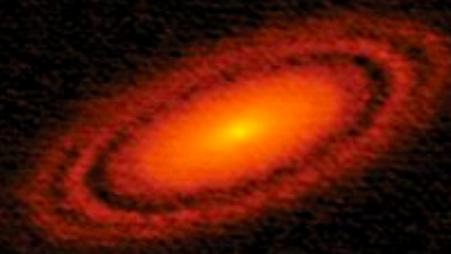
IM Lup



AS 205



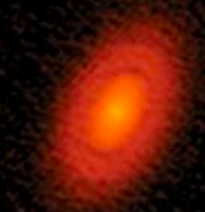
Elias 24



DoAr 25



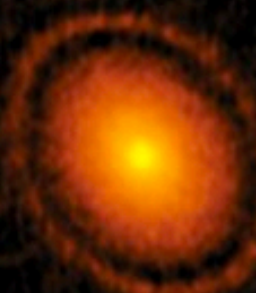
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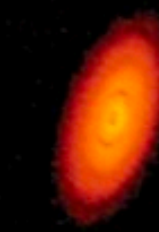
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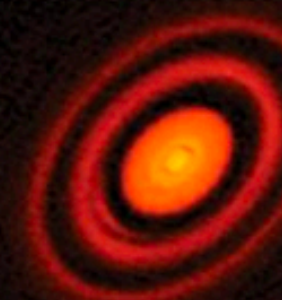
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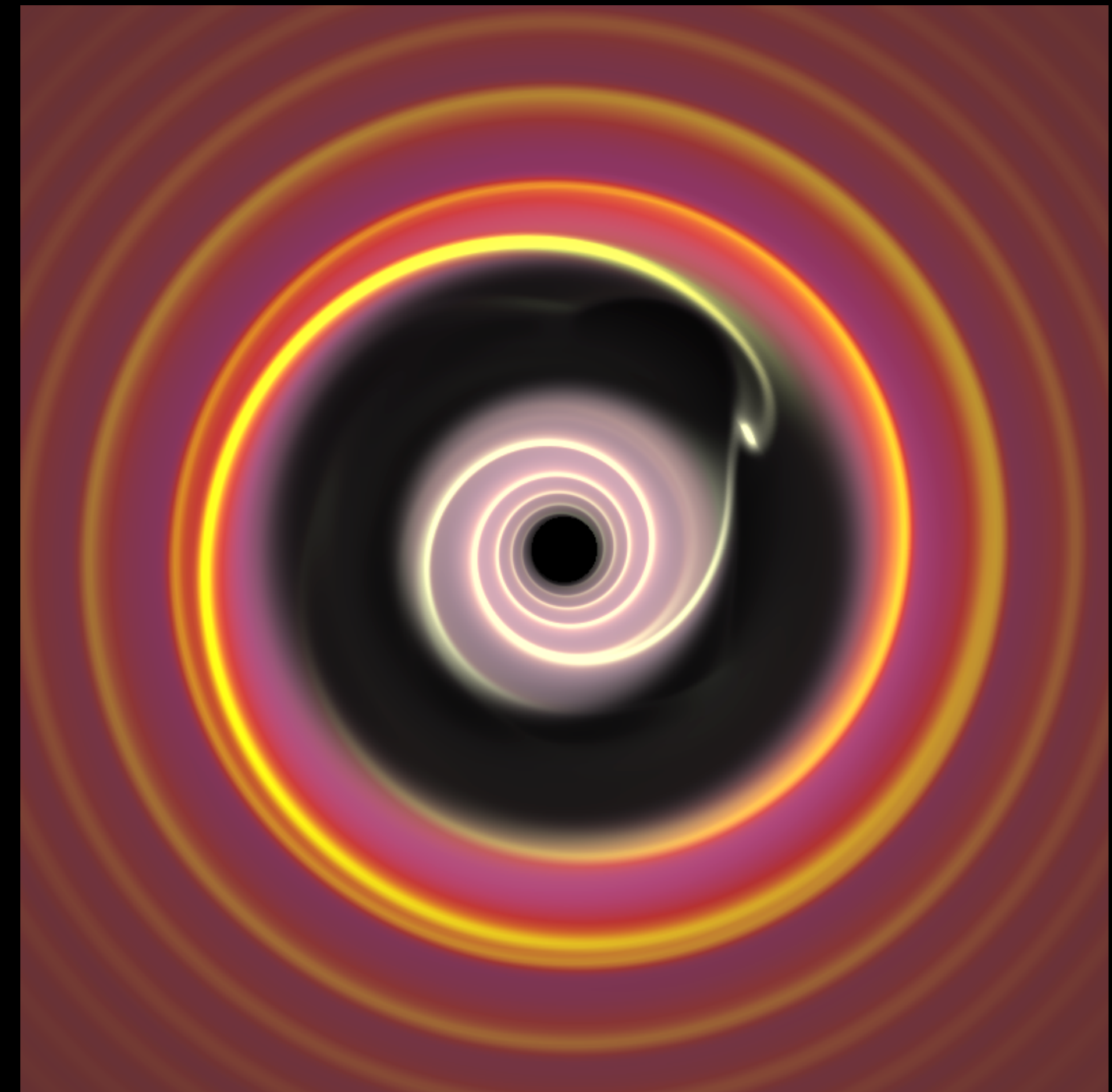
GW Lup



HD 142666

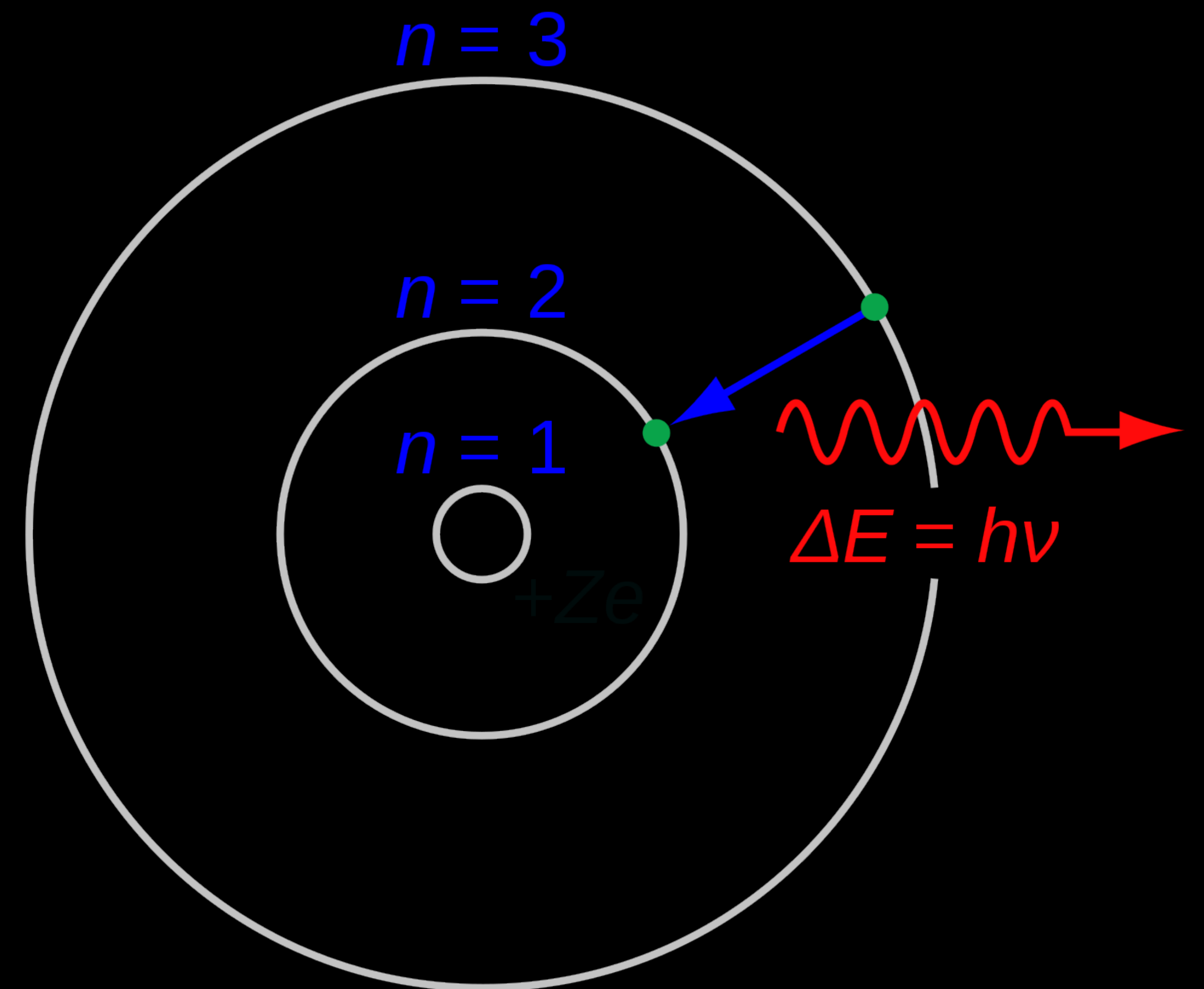
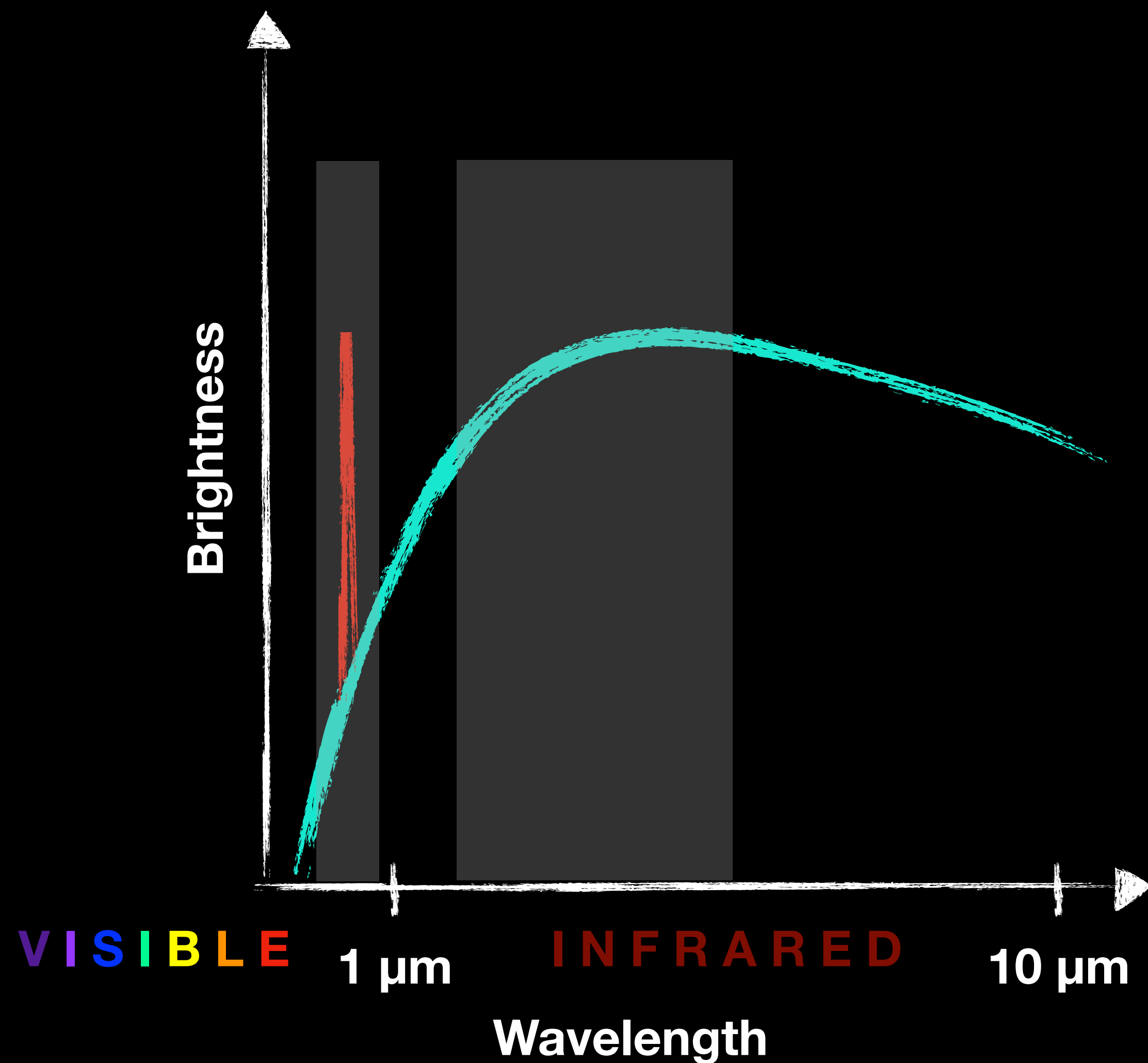


HD 163296



Phil Armitage

Imaging protoplanets: what wavelengths to use



Wikipedia: H-alpha

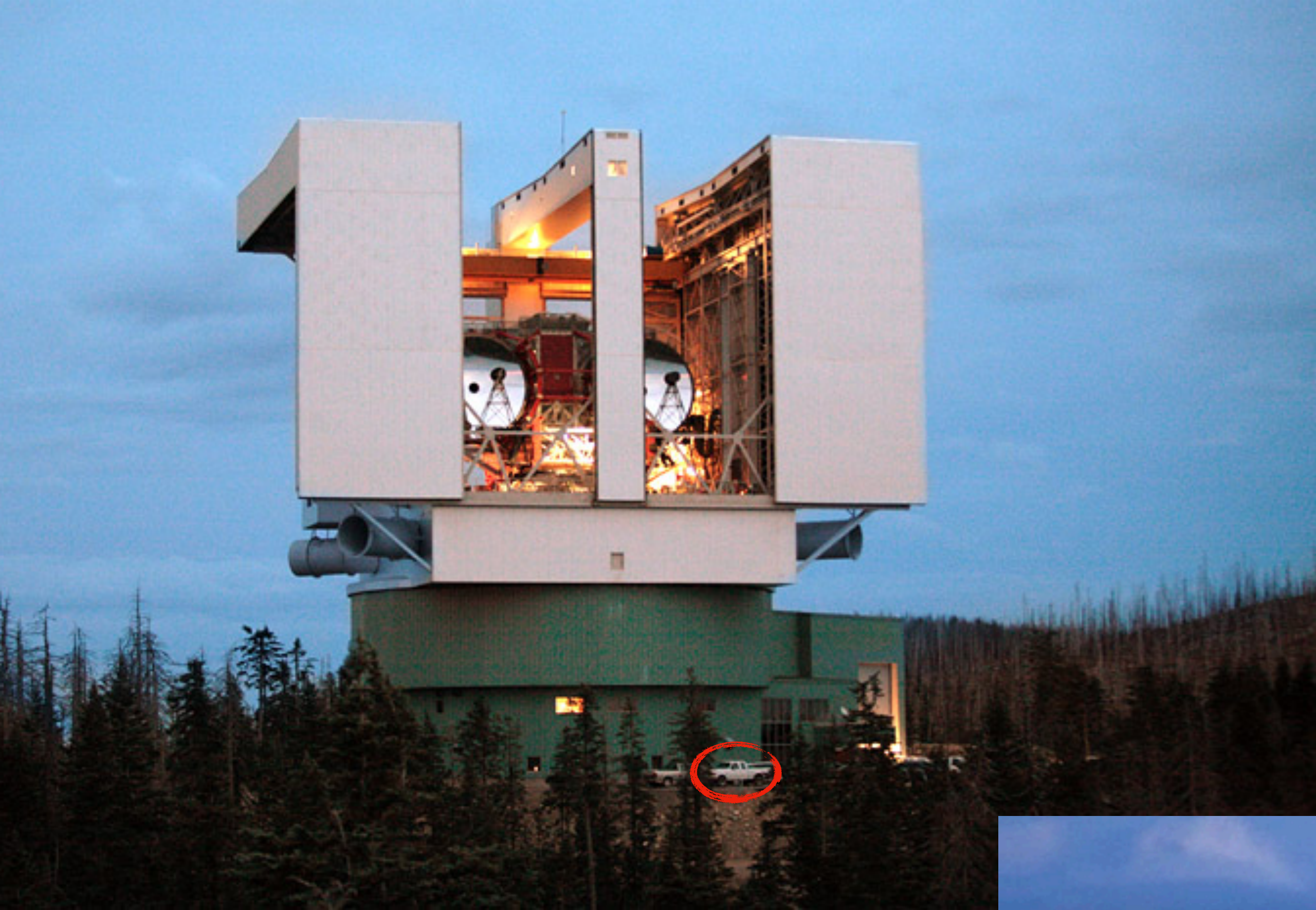
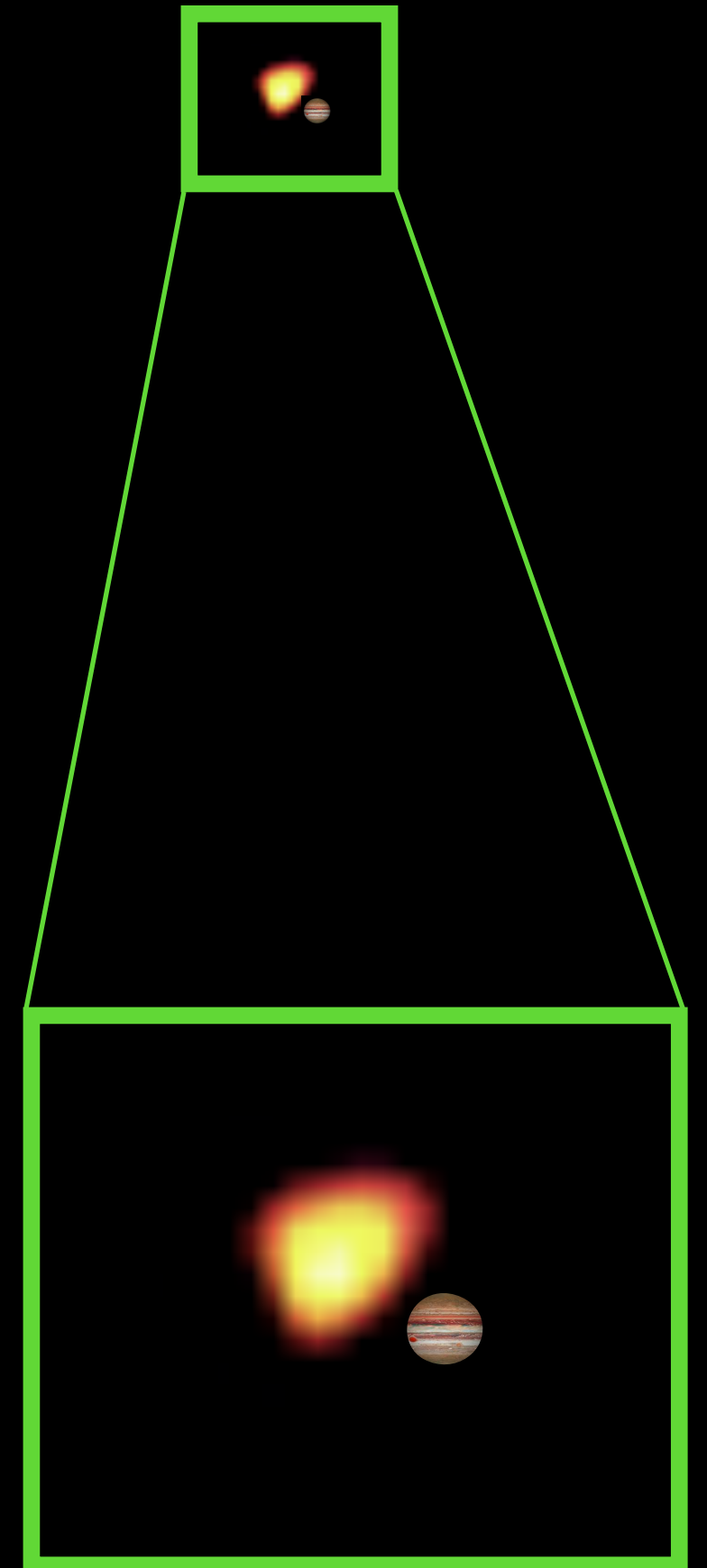
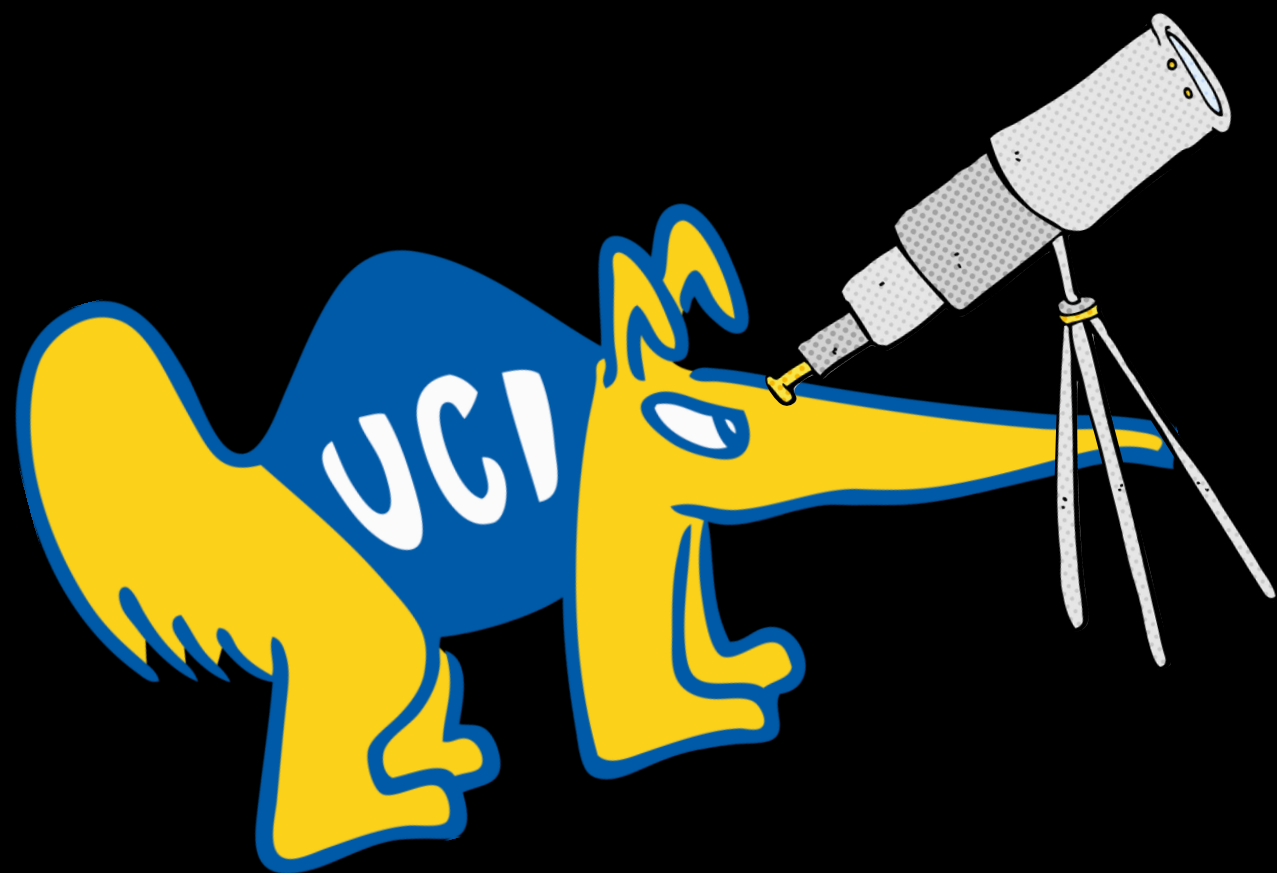
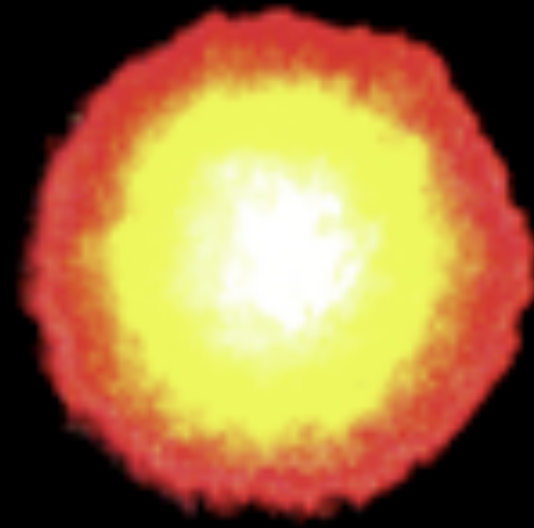


Image credit: Keck Observatory

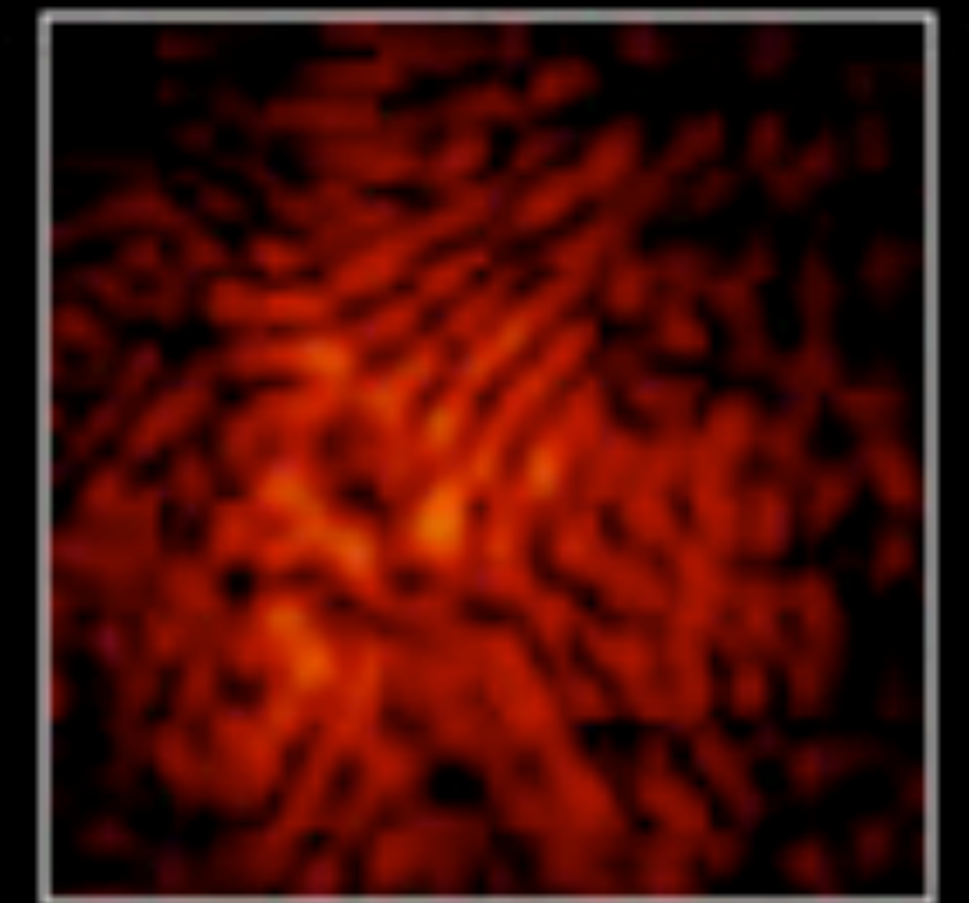
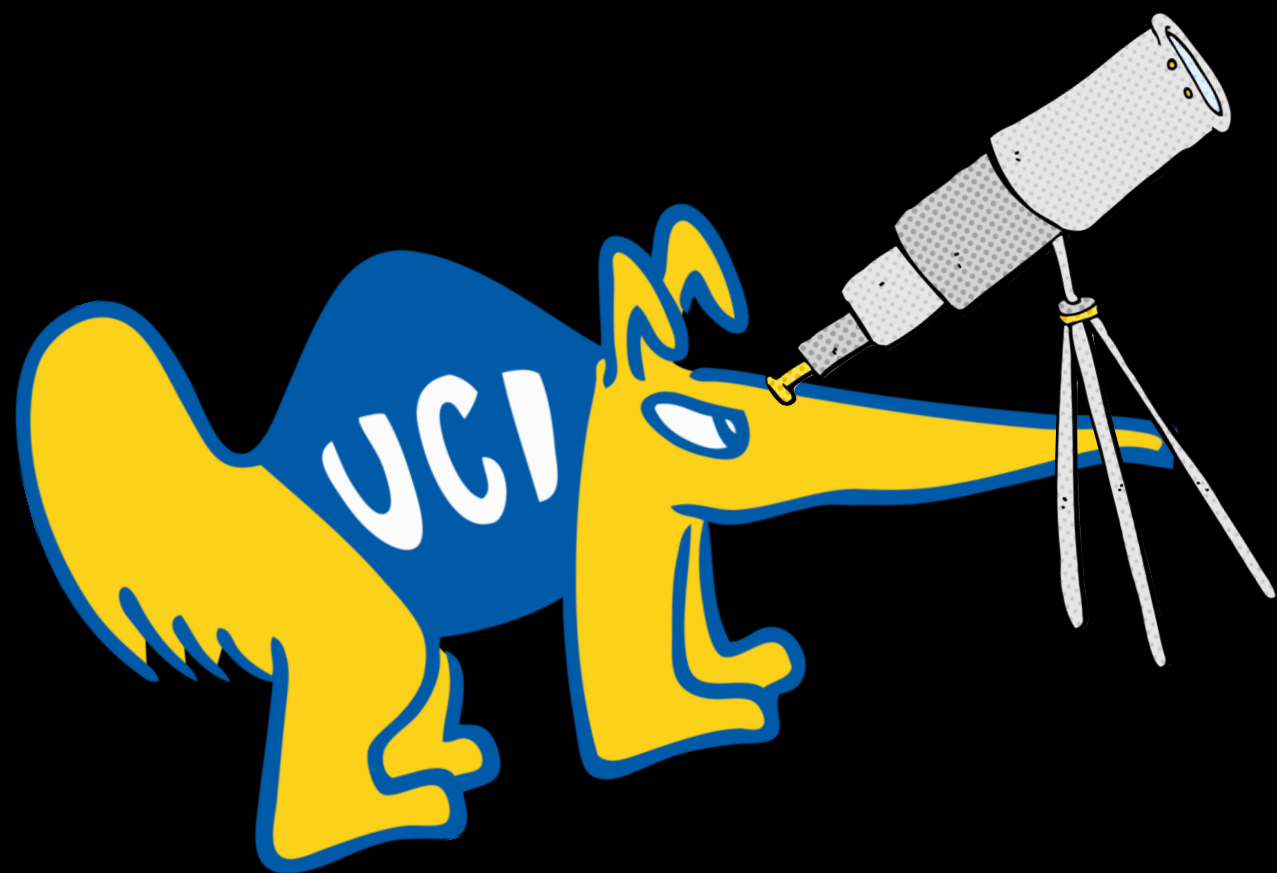
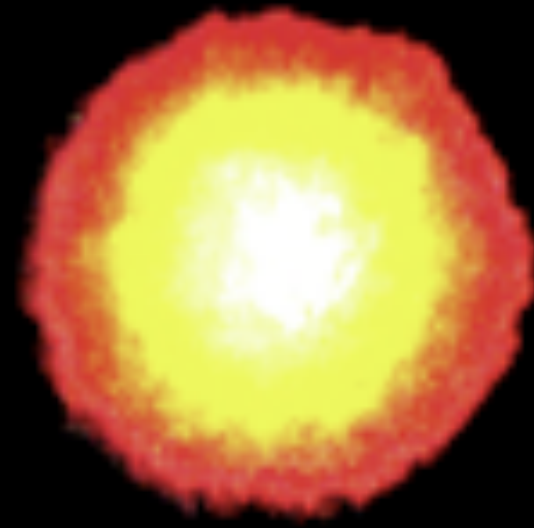
Image credit: Leibniz Institute for Astrophysics



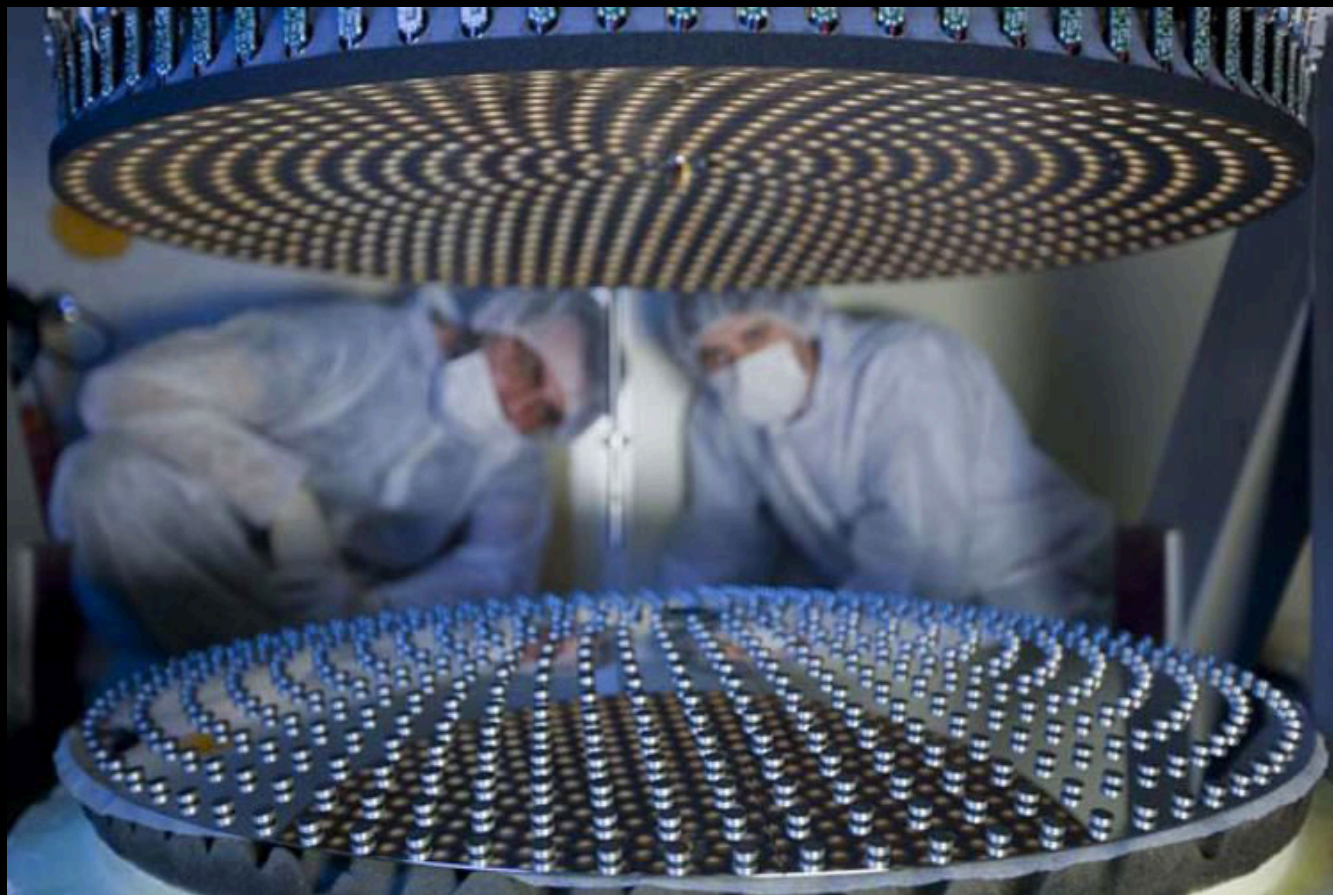
A bigger telescope means a better view...



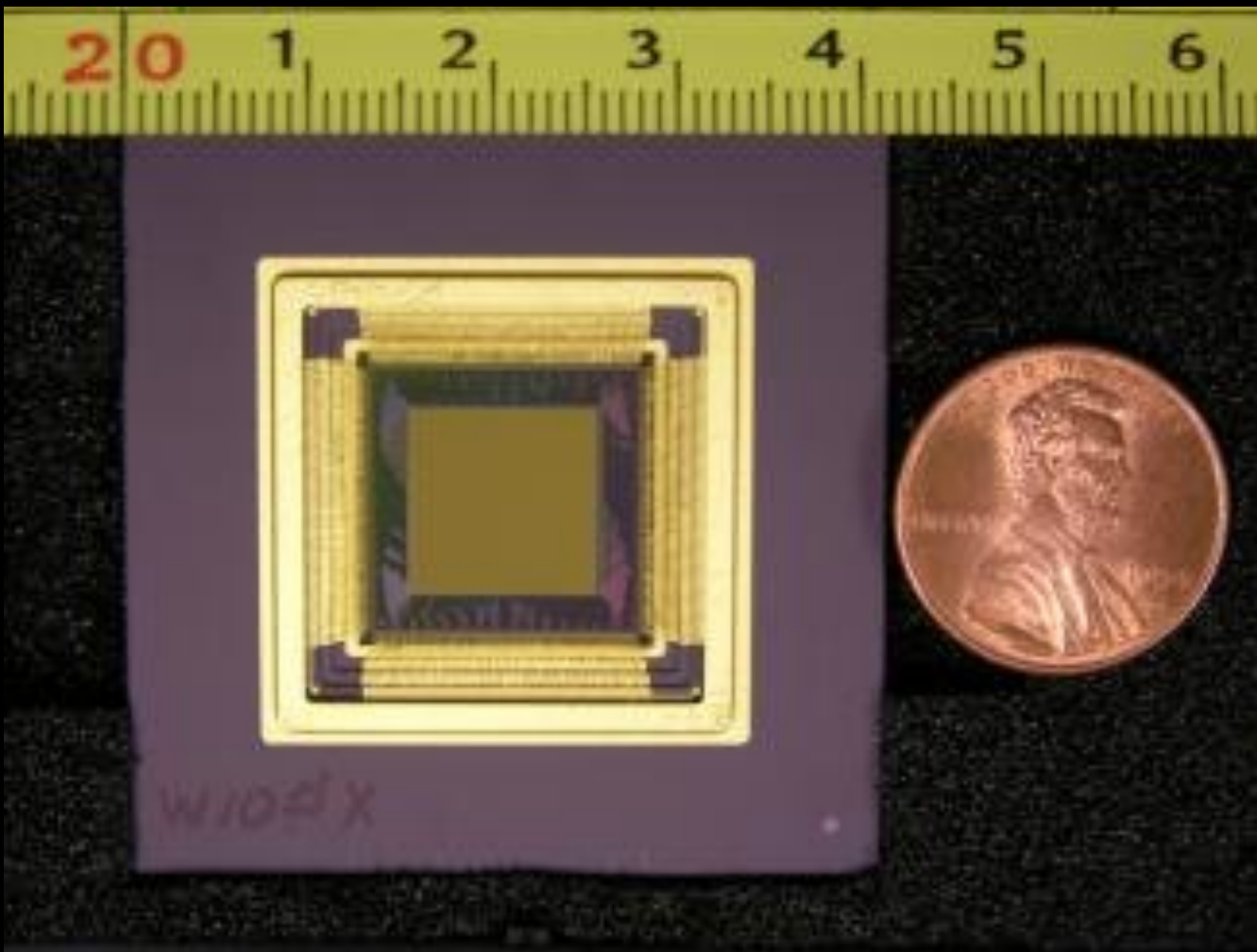
...but stars twinkle.



Adaptive Optics



Microgate



Morzinski+2008

Deformable mirror

Wavefront corrector

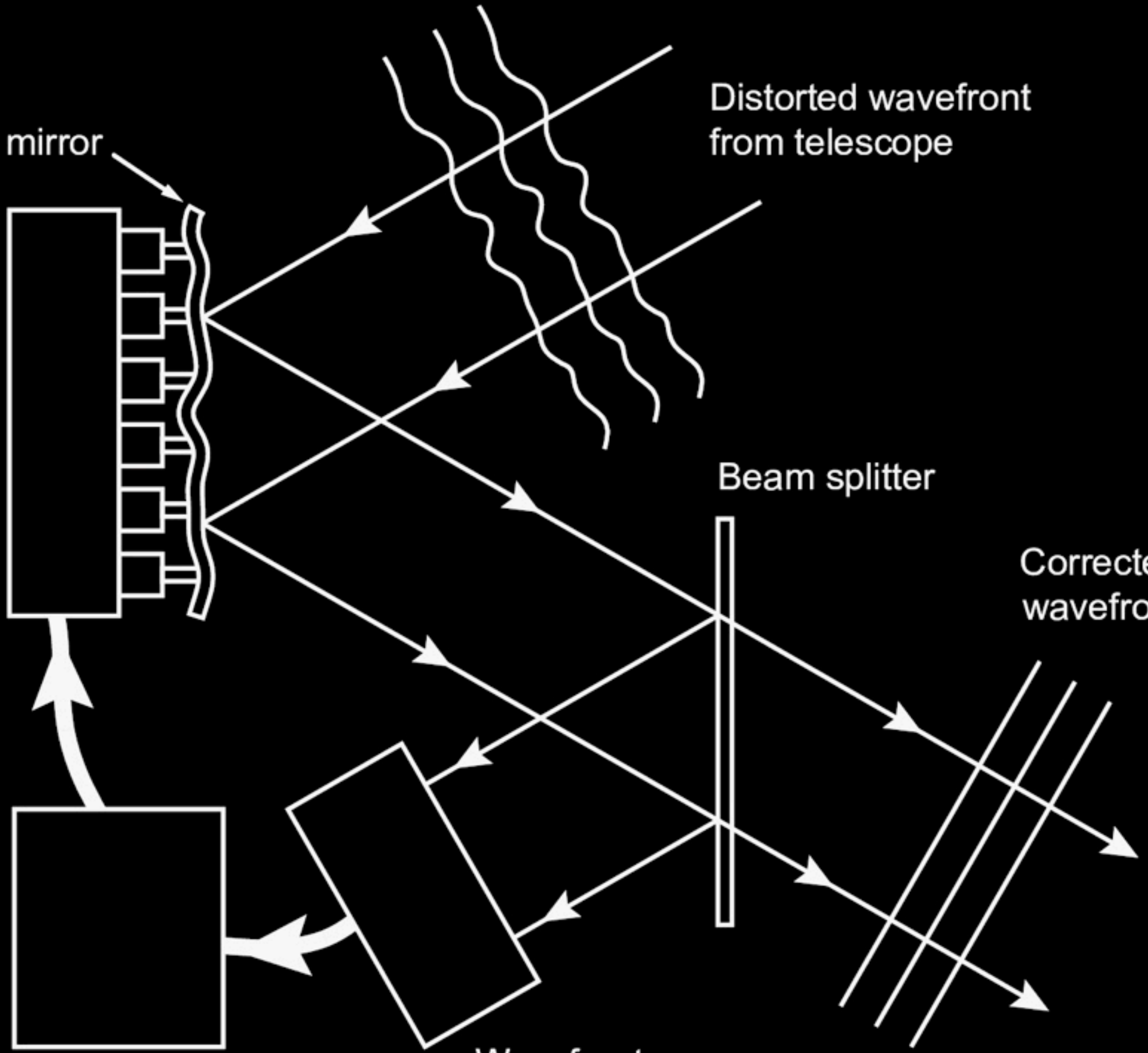
Controller

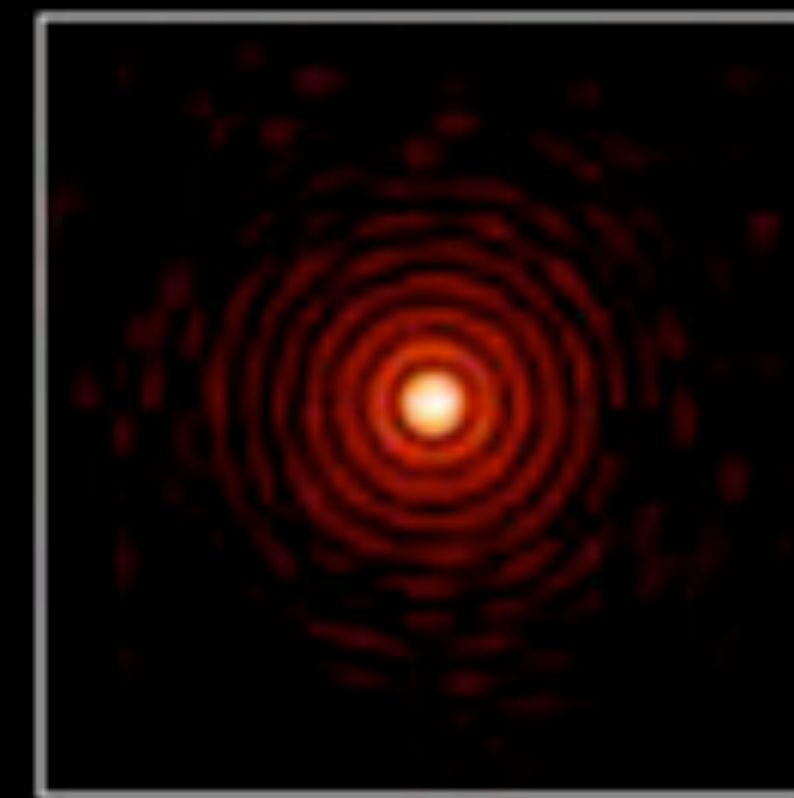
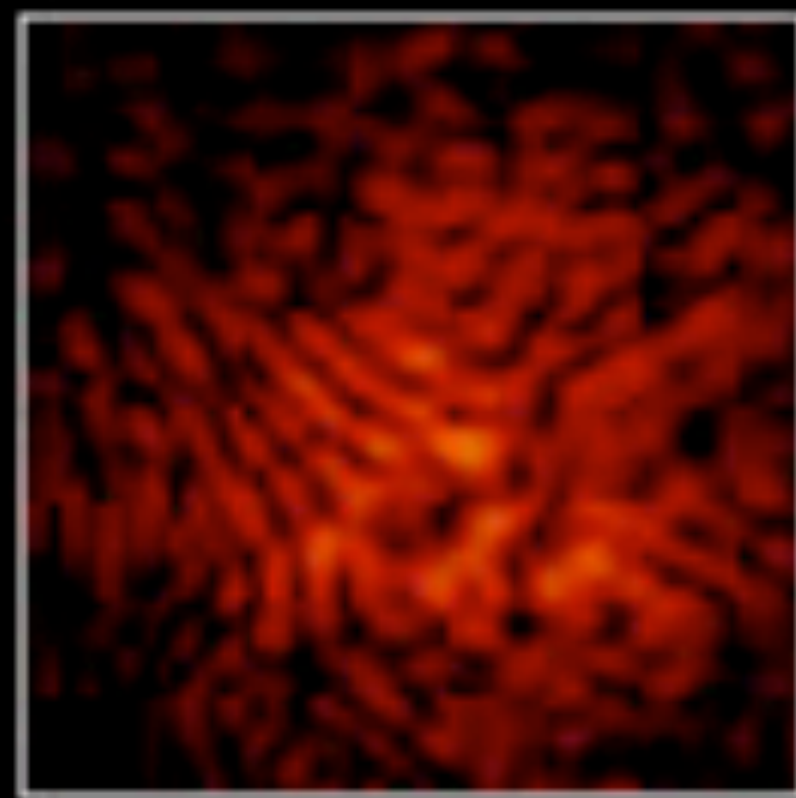
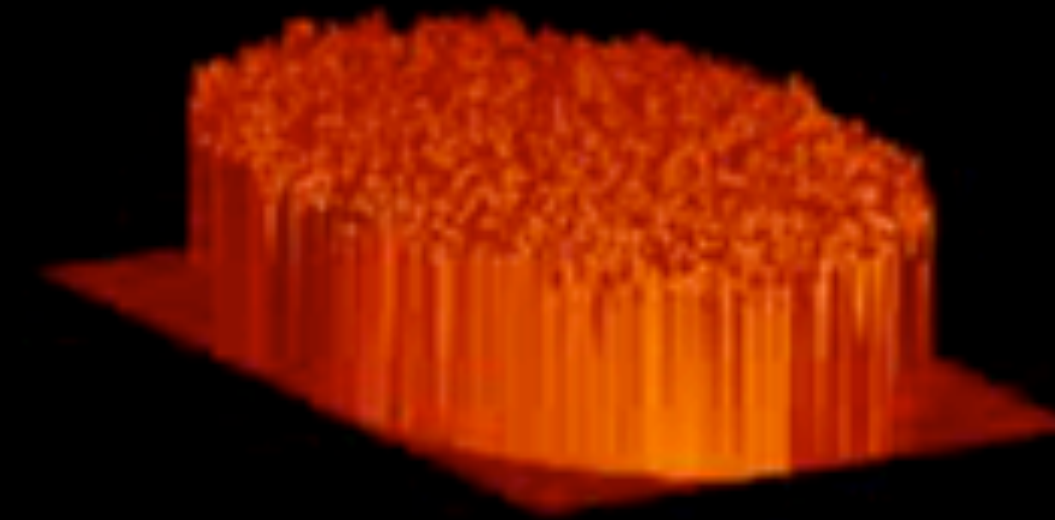
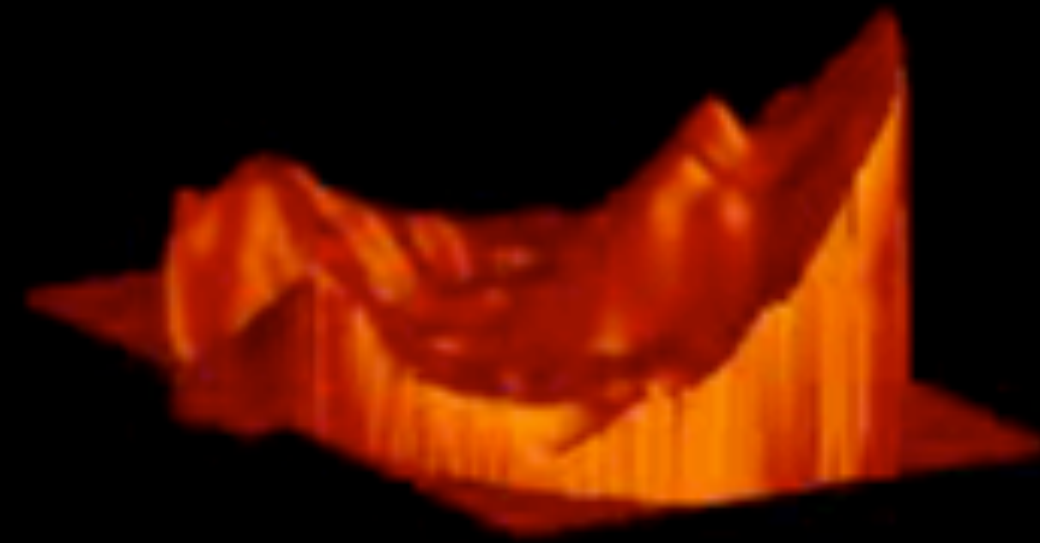
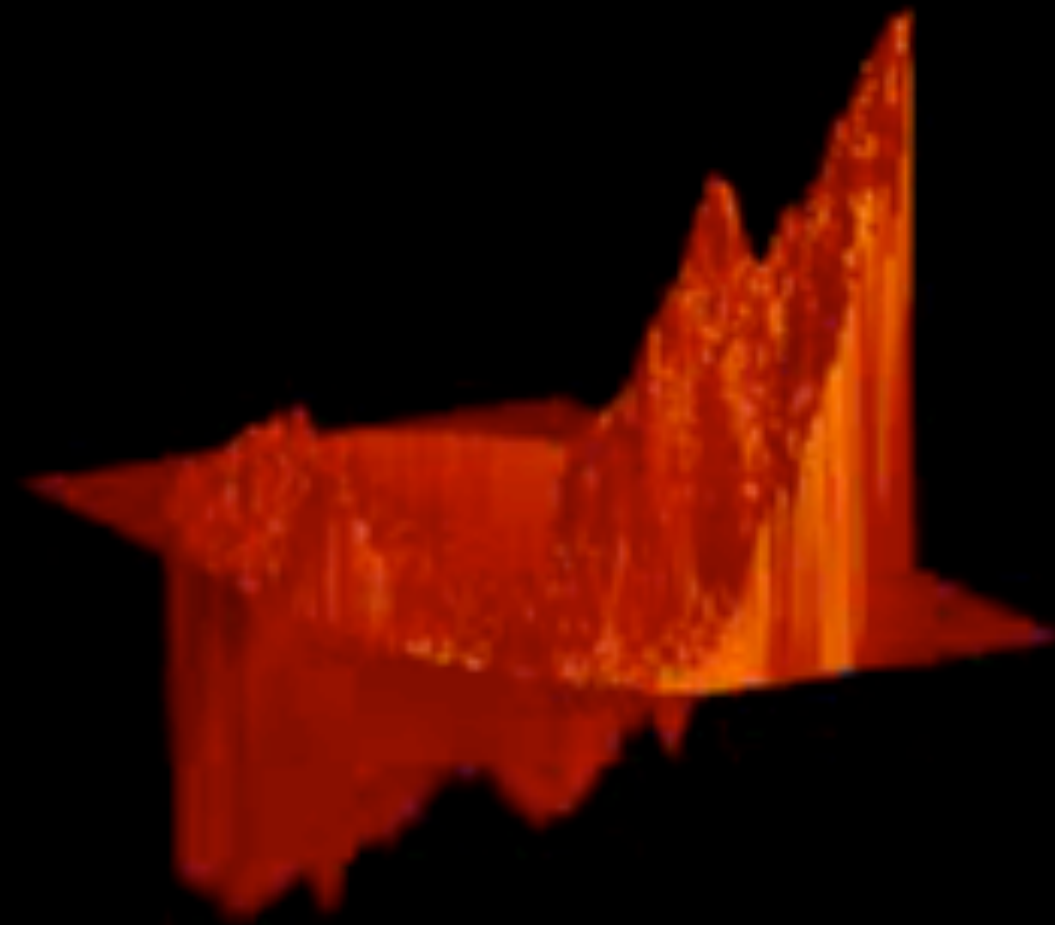
Distorted wavefront from telescope

Beam splitter

Corrected wavefront

Wavefront sensor

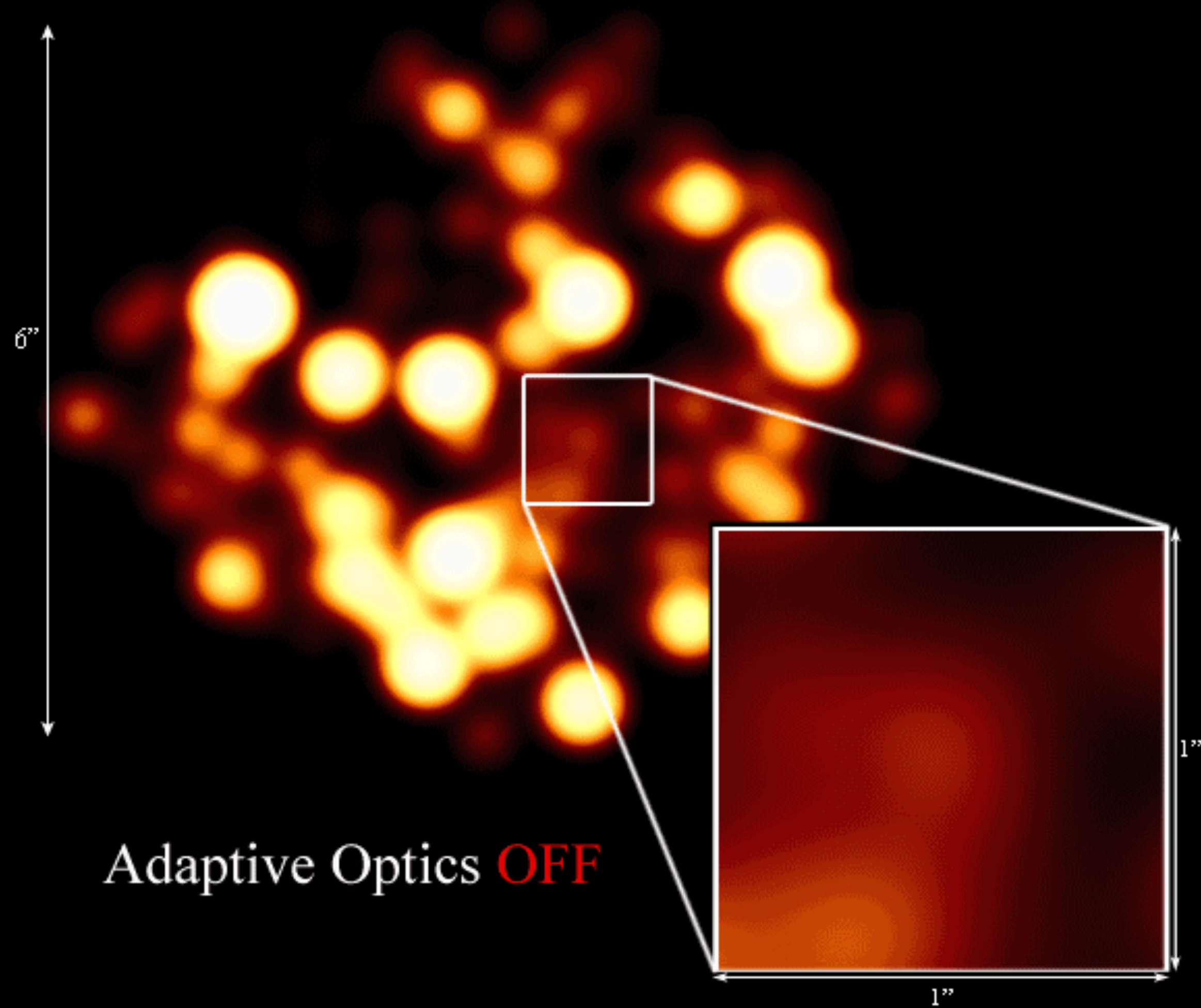




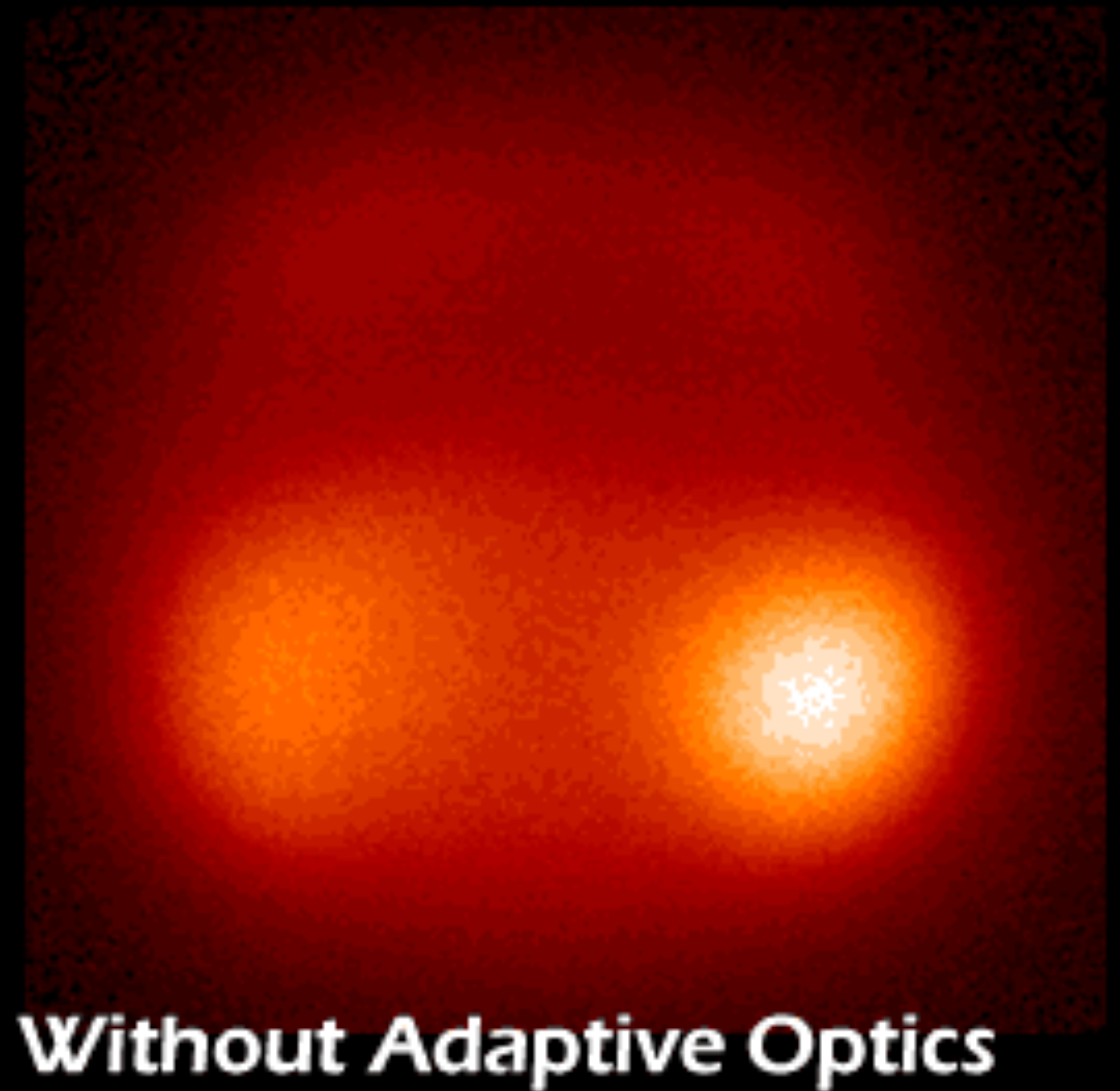
Without Adaptive Optics

With Adaptive Optics

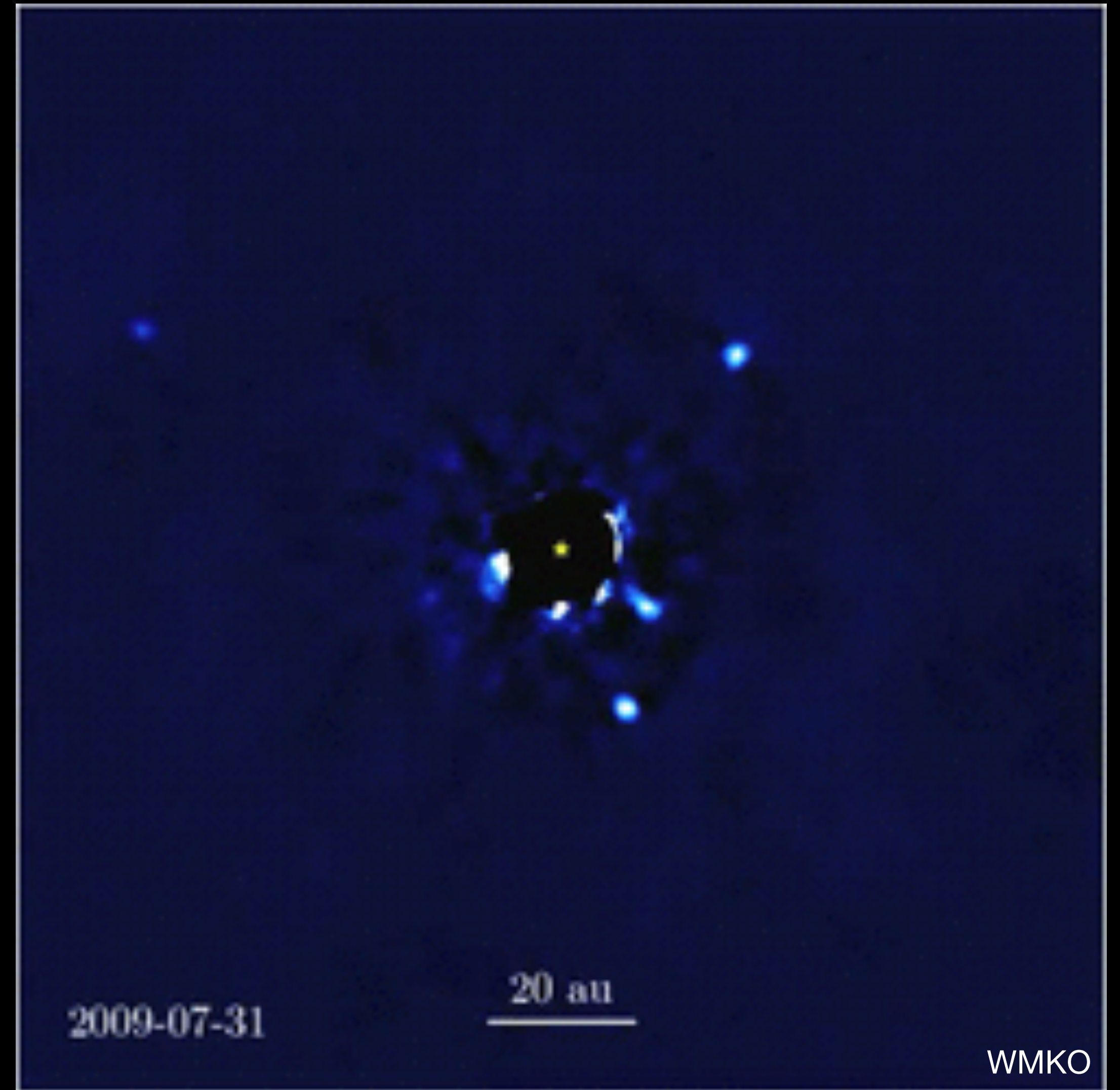
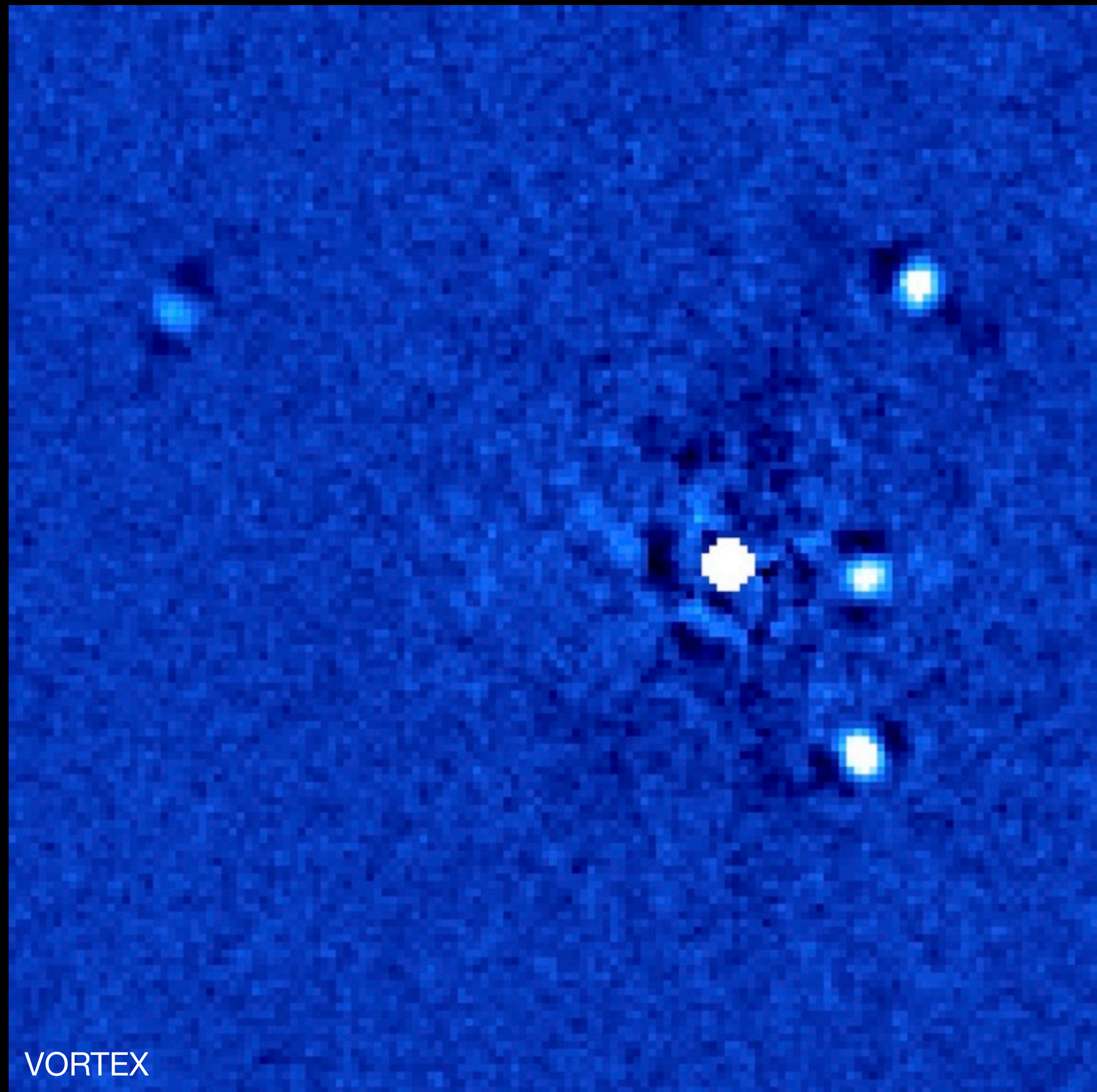
The Galactic Center at 2.2 microns



Neptune at 2.2 microns



Directly Imaged Exoplanets



How faint is a fully-formed exoplanet?

Typical giant exoplanets
are 1 MILLION times
fainter than the stars
they orbit!

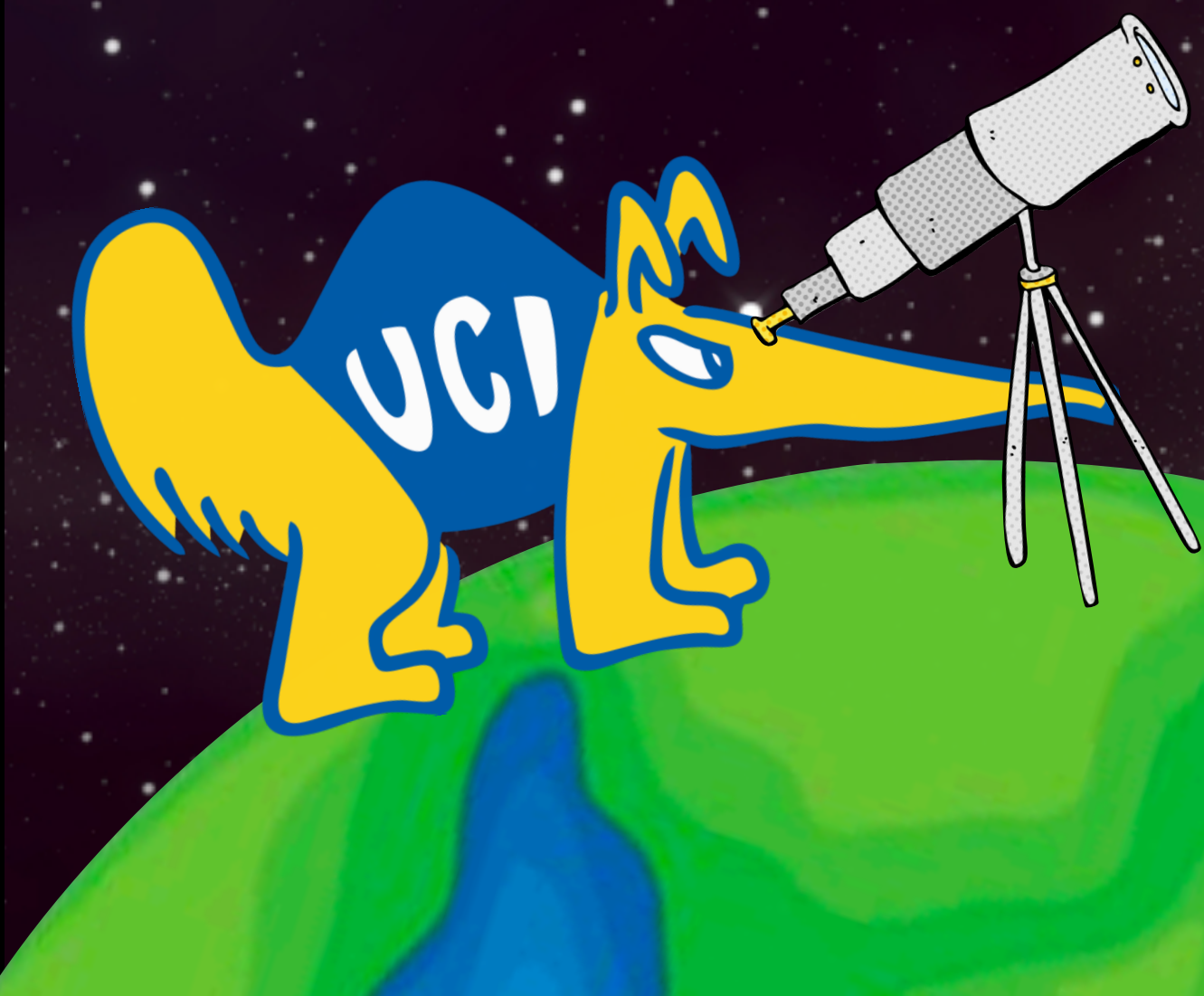


How faint is a forming exoplanet?

Forming giant planets
are 1 HUNDRED to 10
THOUSAND times
fainter than the stars
they orbit!



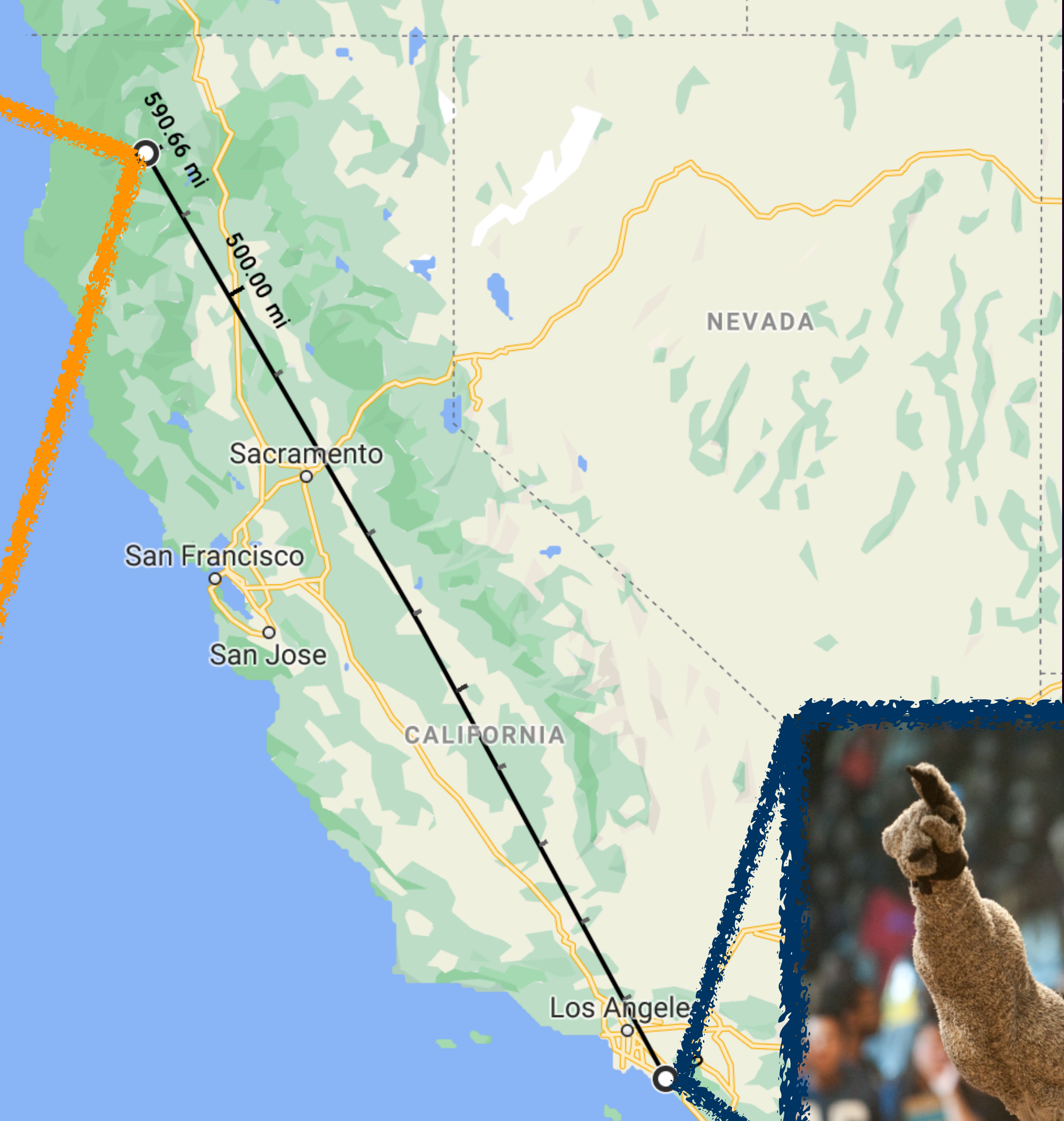
Young stars are really far away!



**~500 light years
(3 quadrillion miles)**

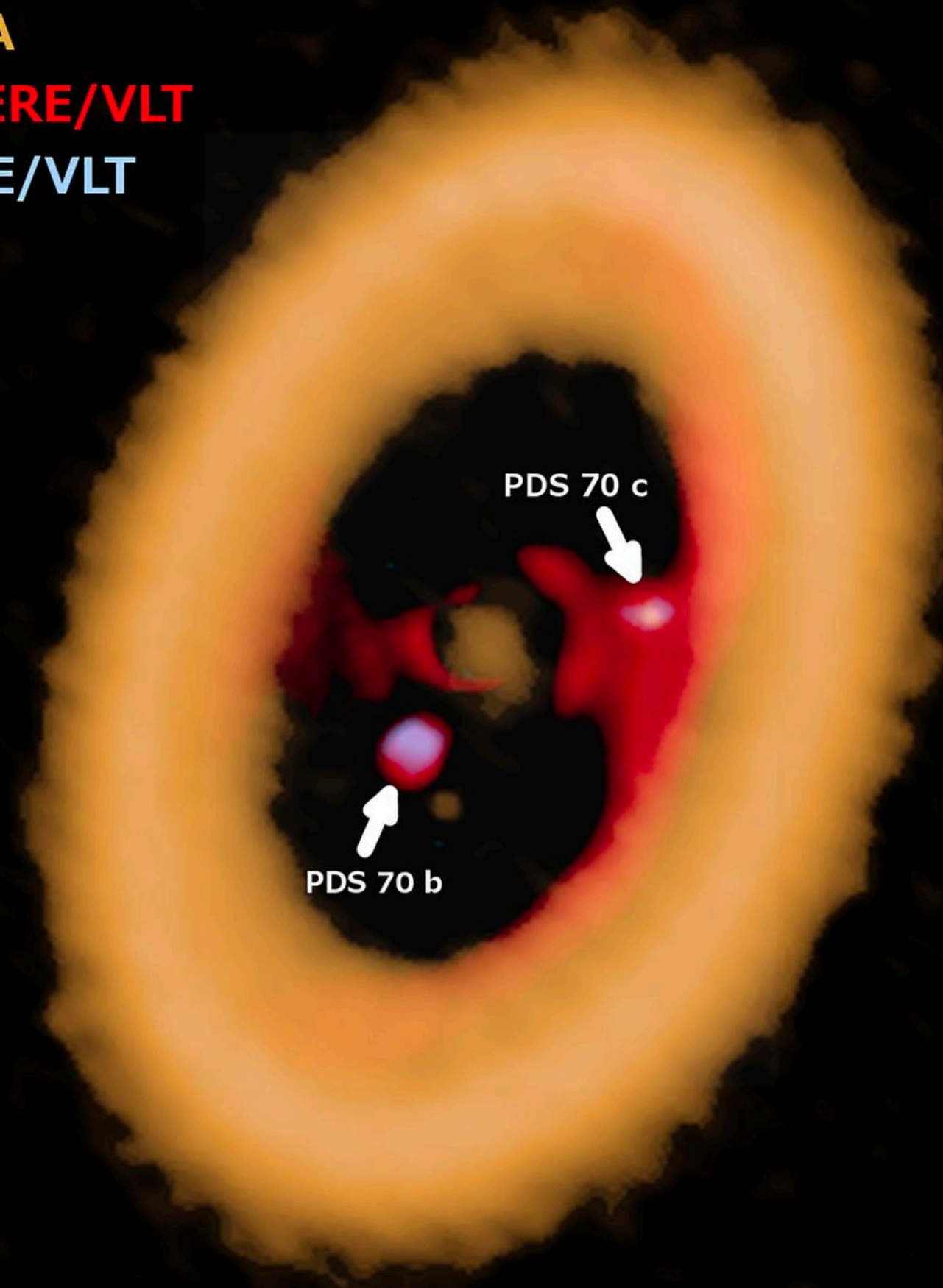


What does that mean?



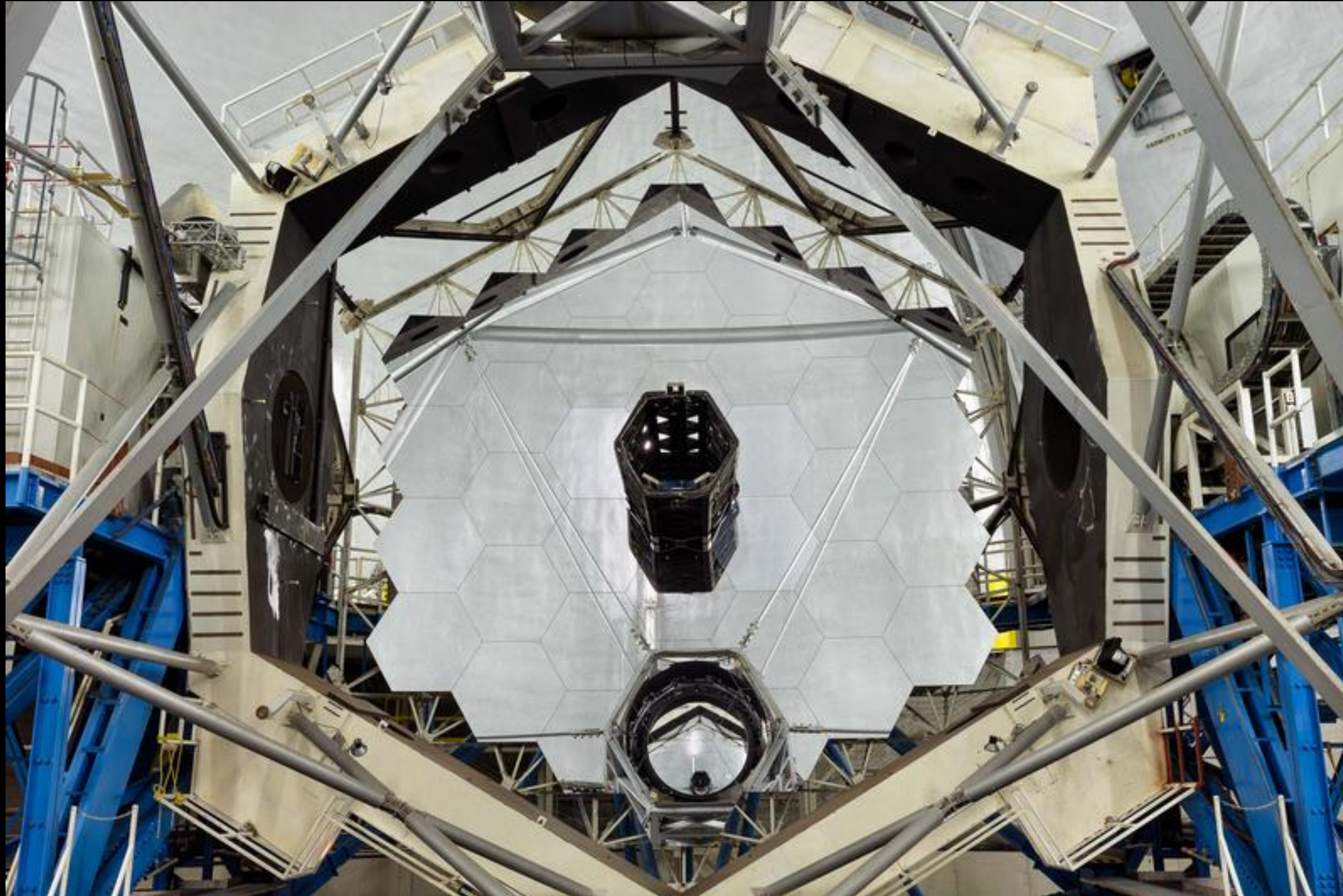
PDS 70: State of the Art with “Traditional” Methods

ALMA
SPHERE/VLT
MUSE/VLT

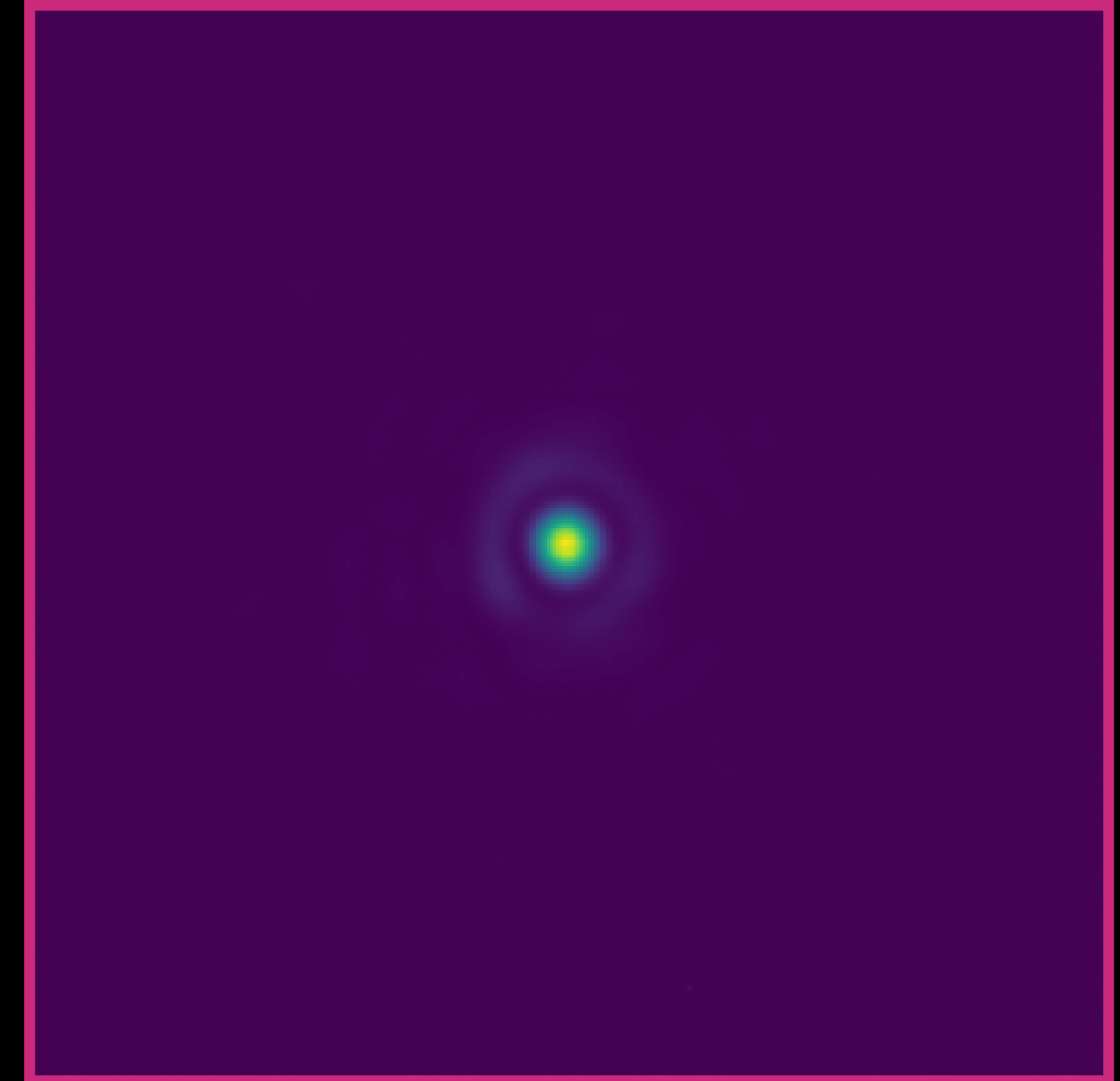


- We are now directly imaging protoplanets!
- Wide angular separations (> 20 AU)
- High masses (~ 5 - 10 Jupiter masses)
- Goals for the protoplanet census:
 - Tighter orbital separations
 - Lower masses
 - Detect and characterize at longer wavelengths ($3 - 5 \mu\text{m}$)

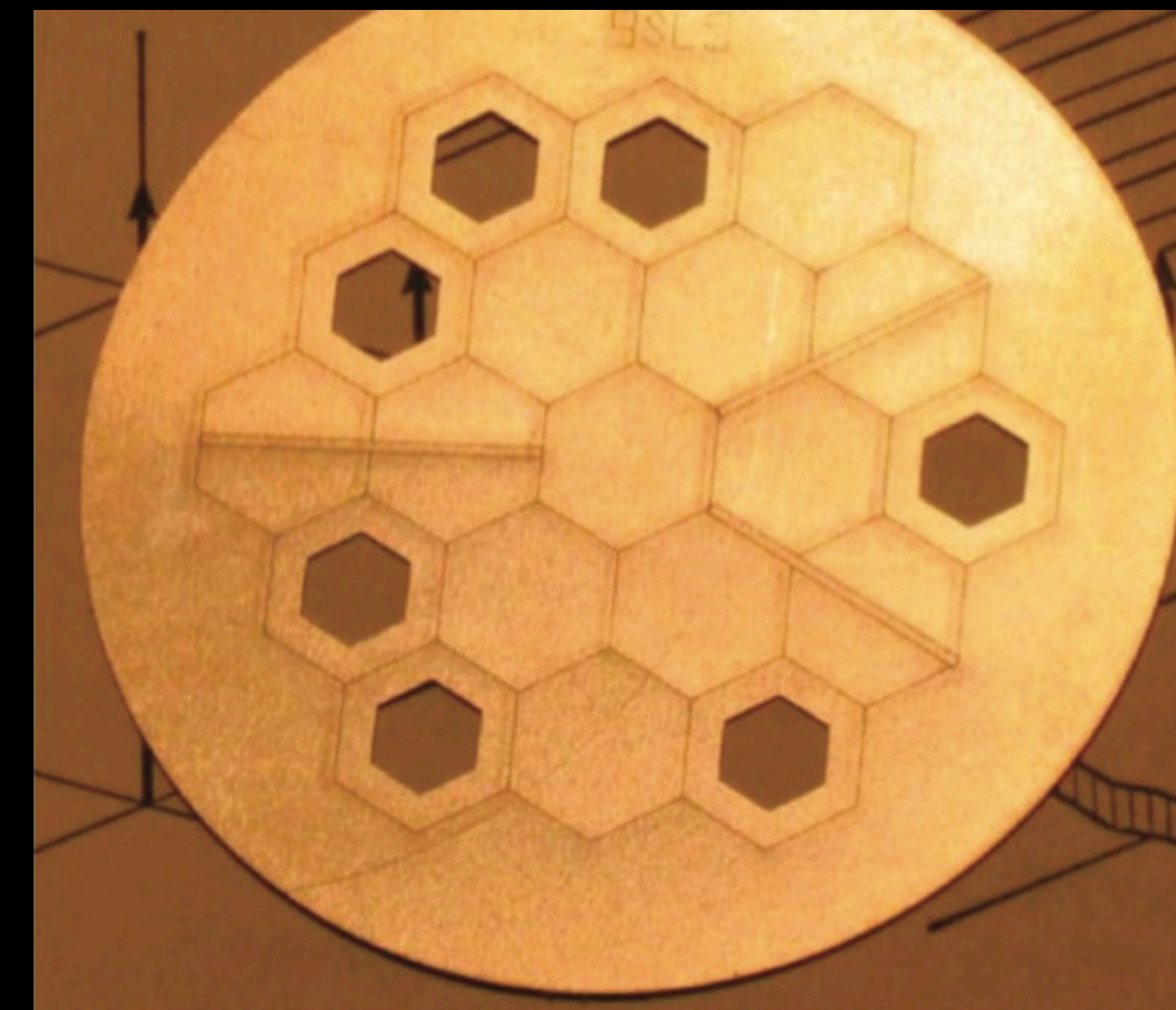
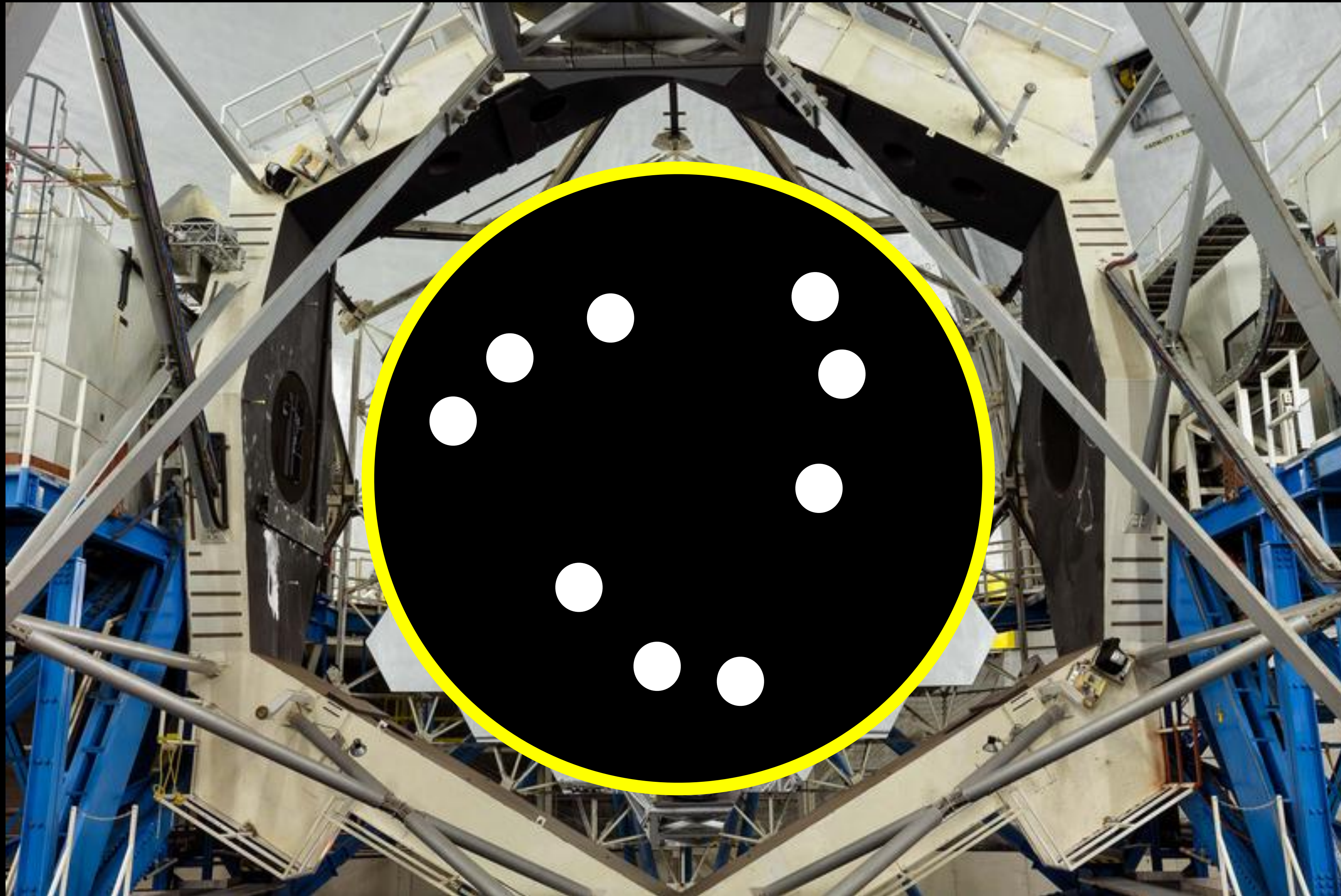
Aperture Masking: Super Resolution



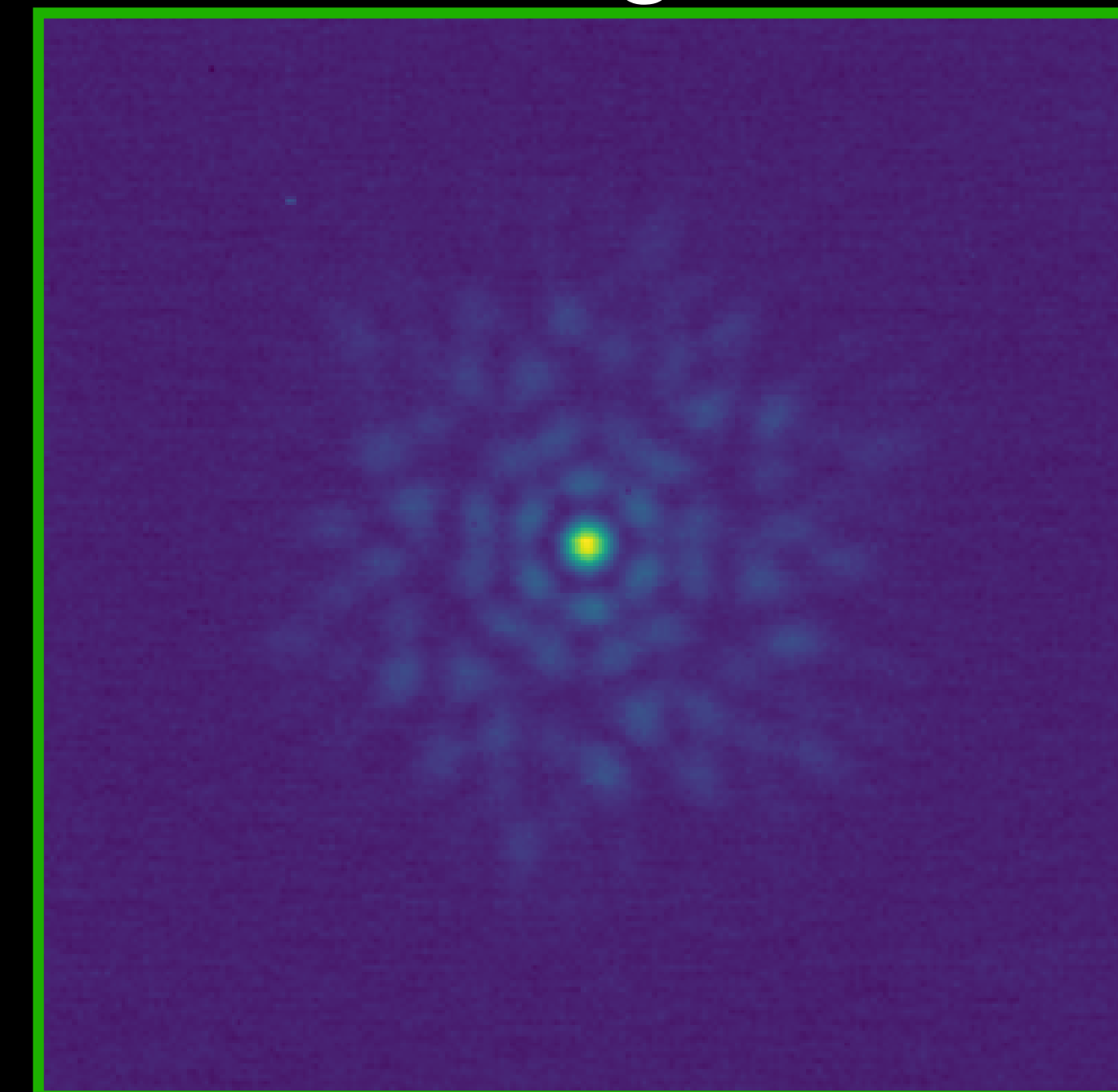
Image



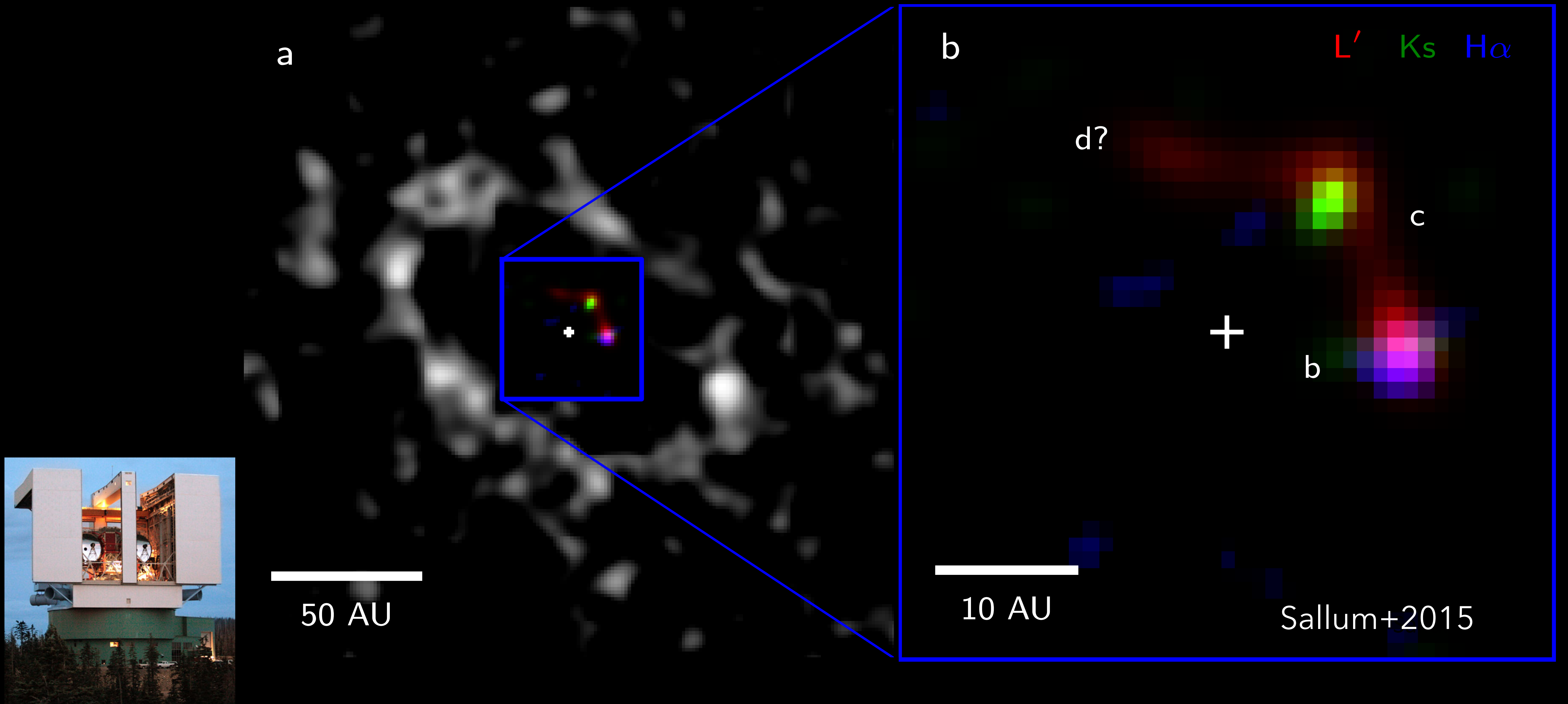
Aperture Masking: Super Resolution



Image

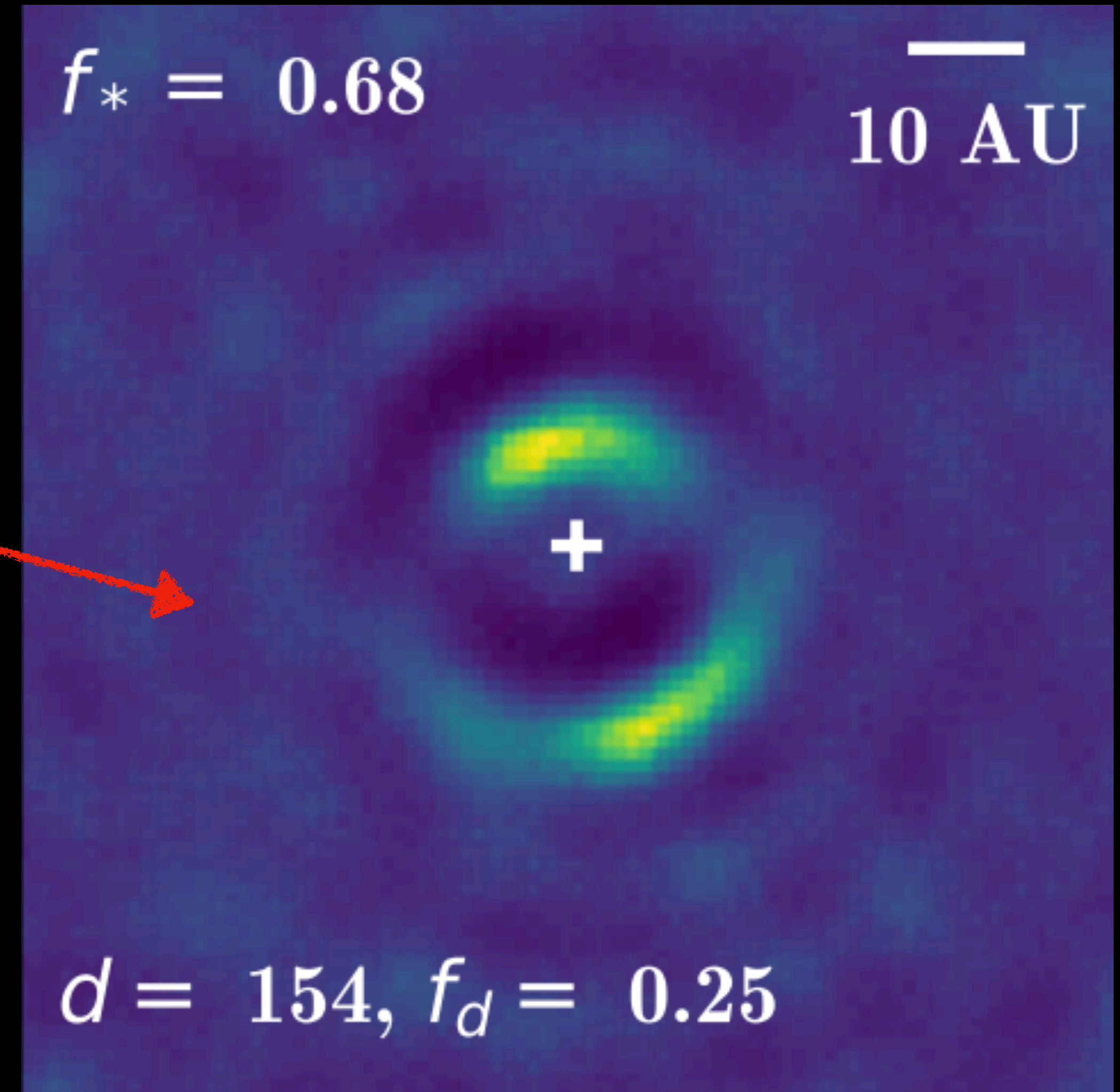


Highlights: planet candidates in protoplanetary disks



Highlights: characterizing the innermost regions of protoplanetary disks

Spirals within a protoplanetary disk millimeter clearing - possible disk-planet interaction!



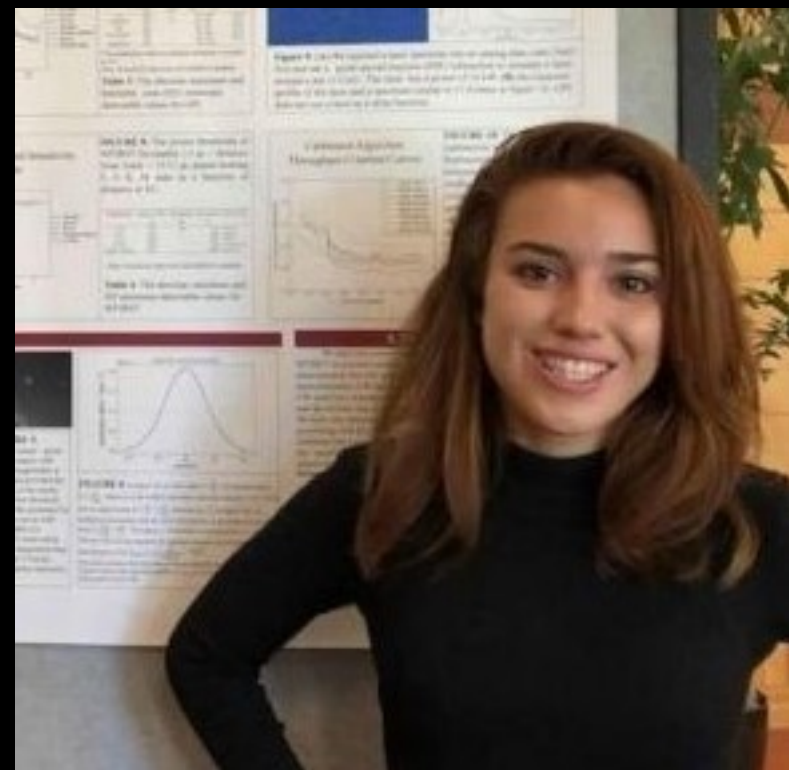
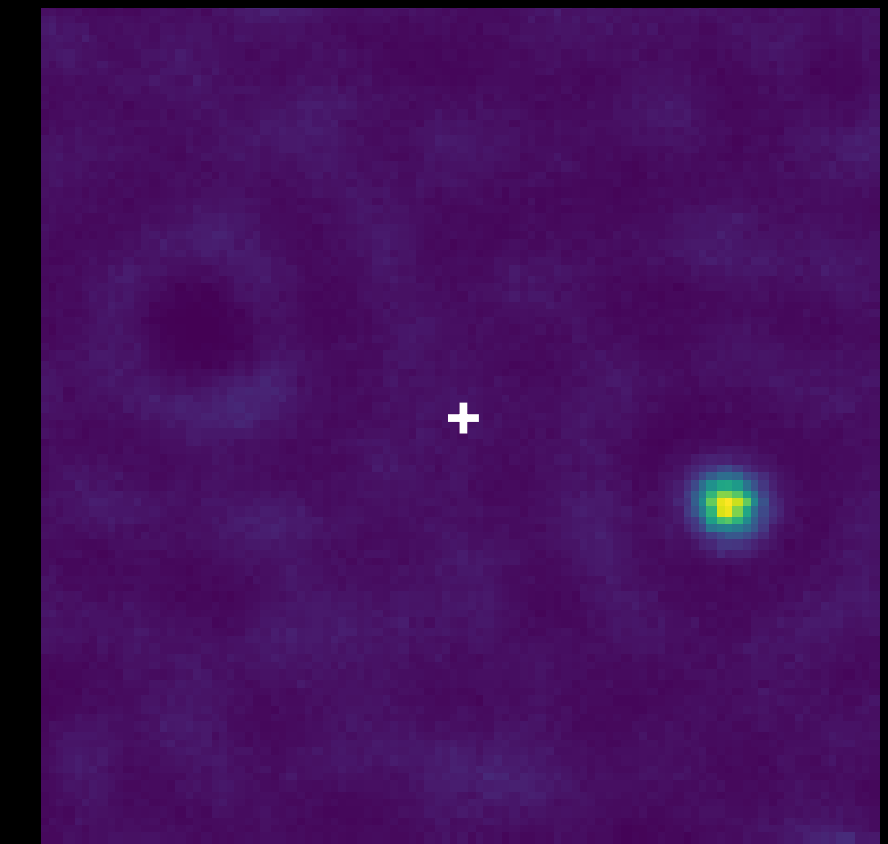
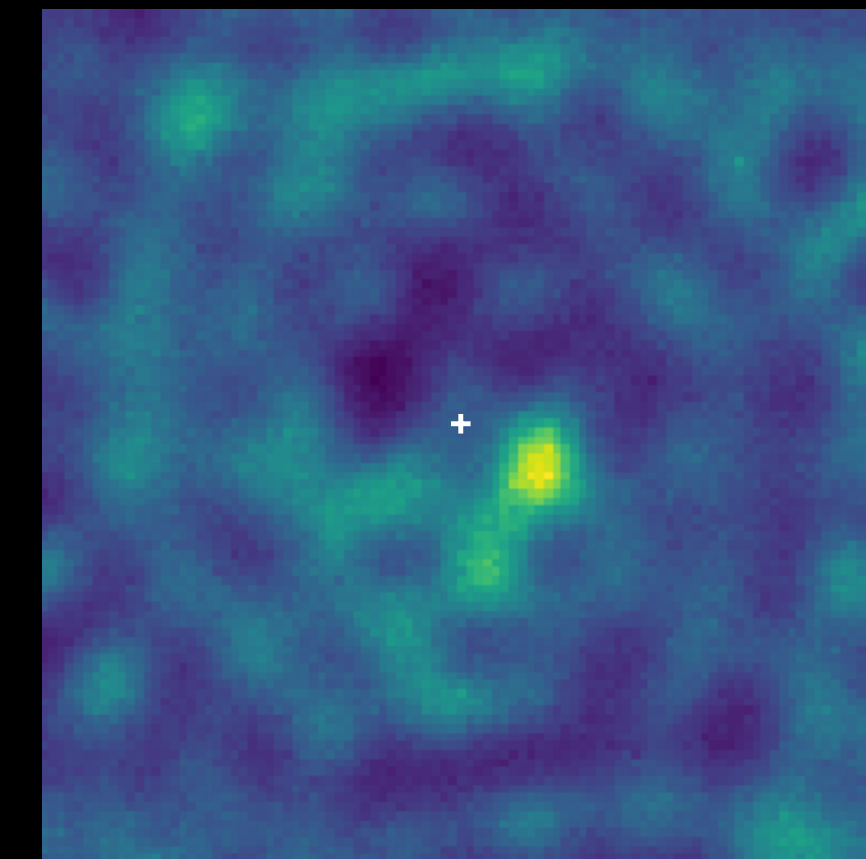
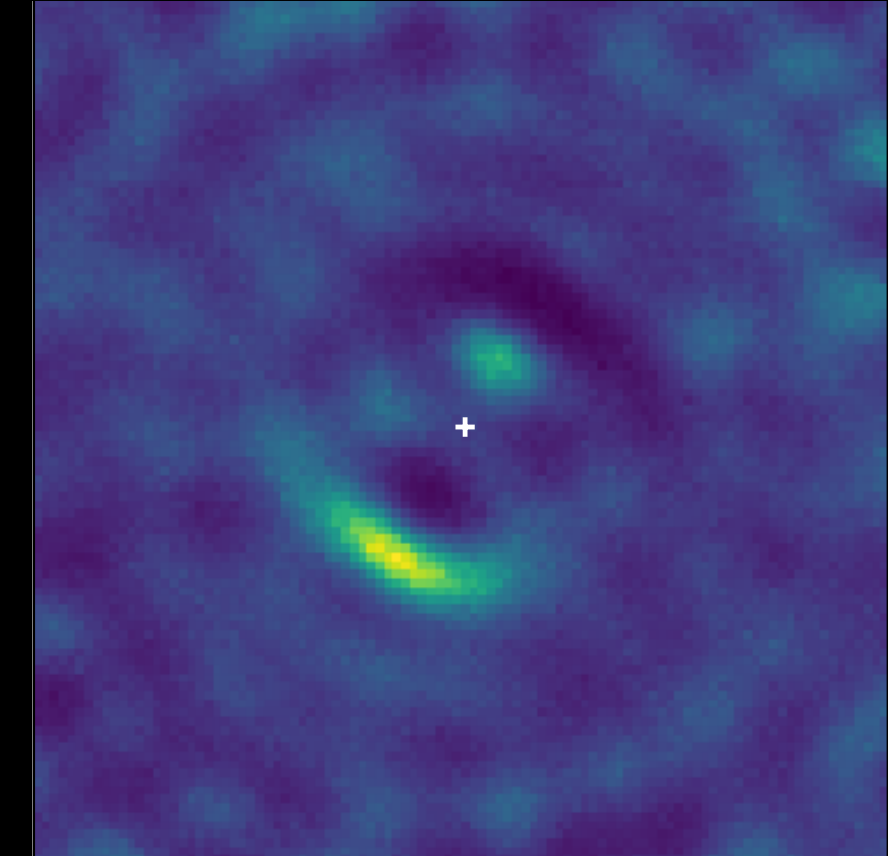
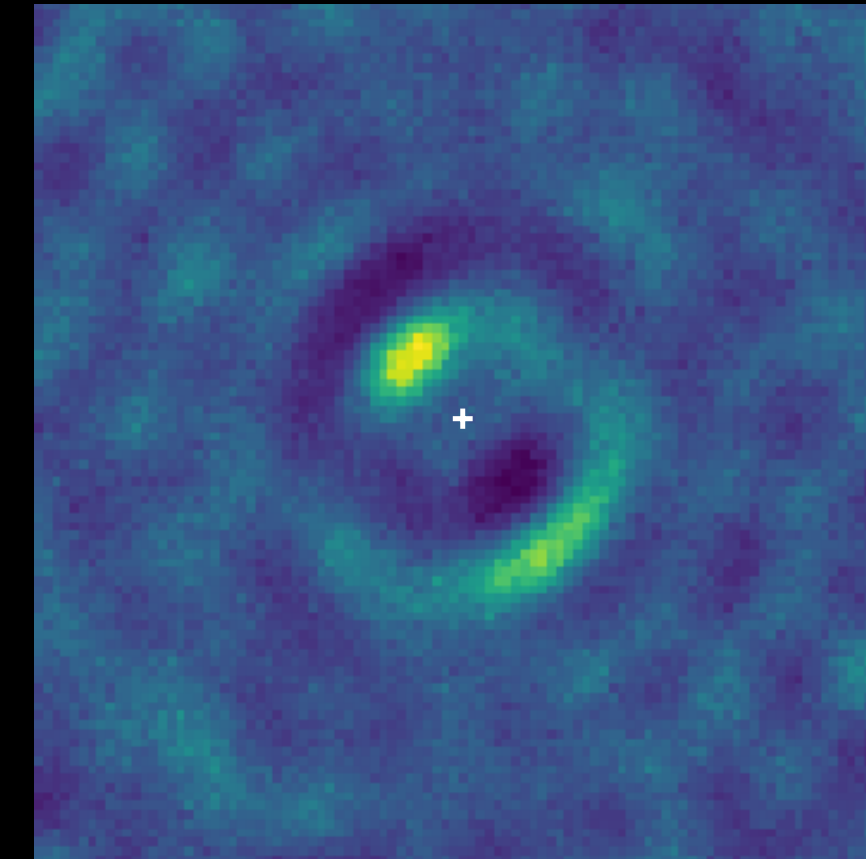
A Keck Masking Survey of Gapped Protoplanetary Disks

A few big questions:

How common are giant forming planets in gapped protoplanetary disks?

How quickly do forming planets sweep up material?

What are the orbital architectures of the youngest planetary systems, and how do they compare to mature ones?



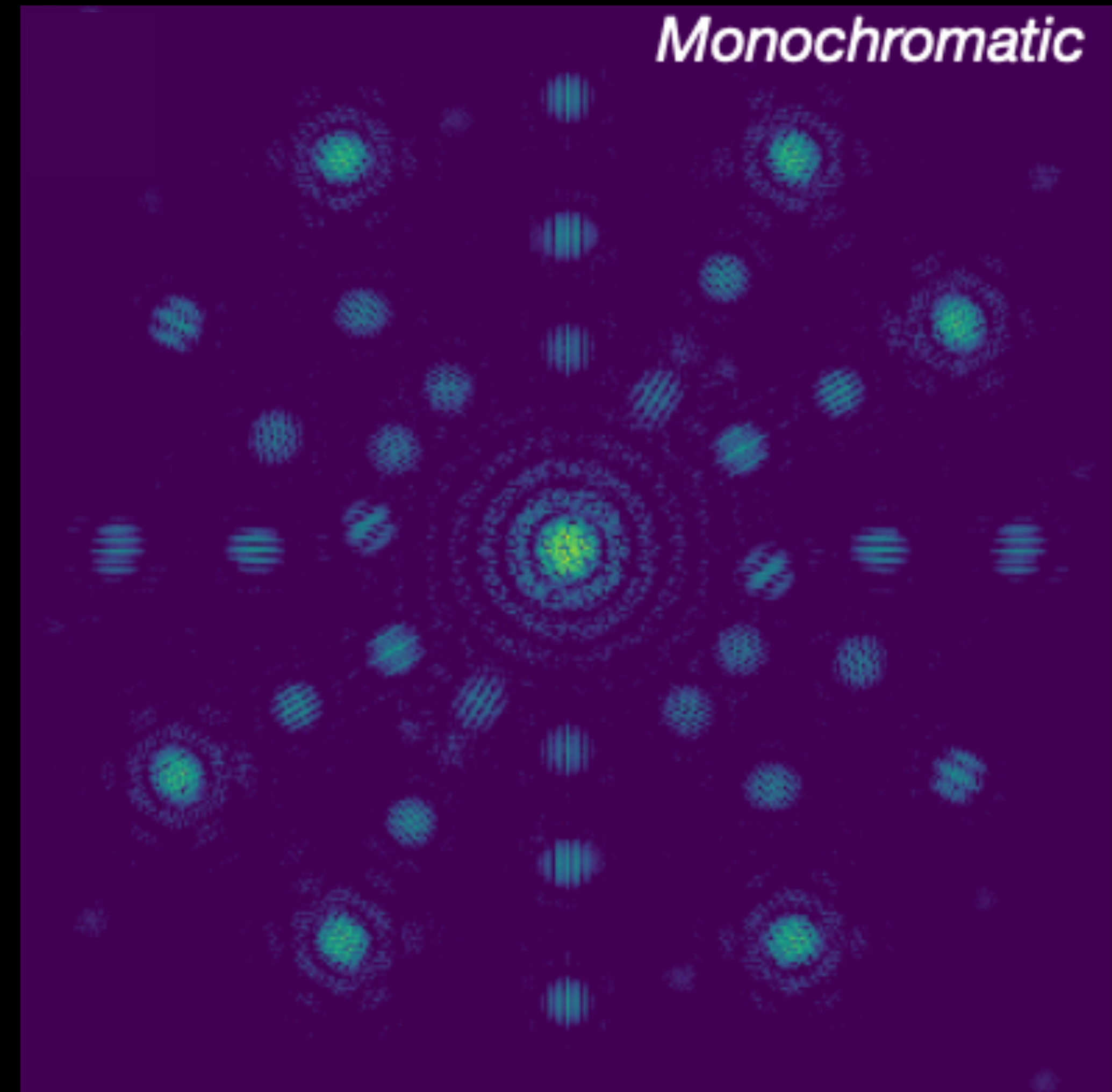
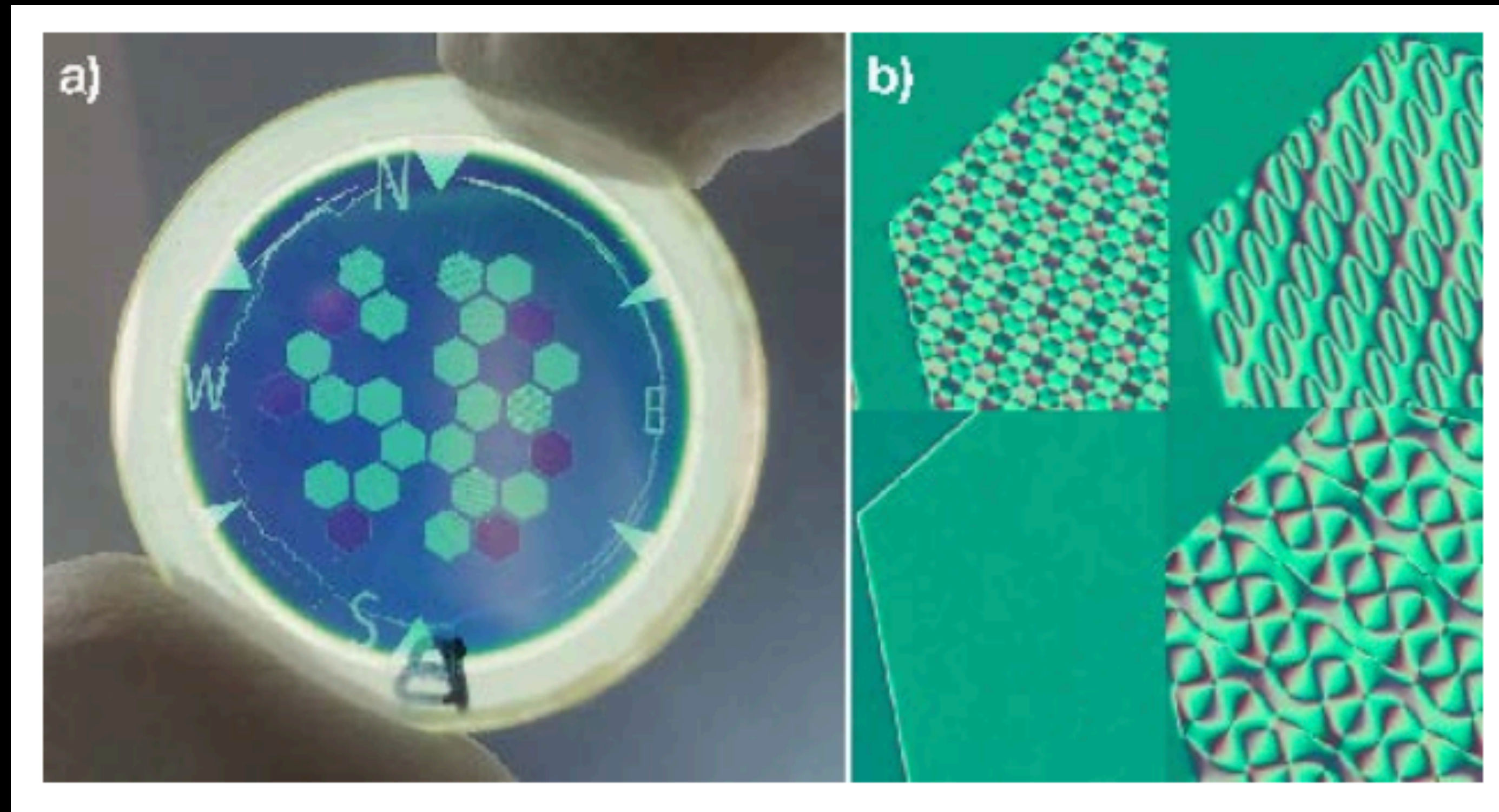
Christina Vides



Upcoming Instruments

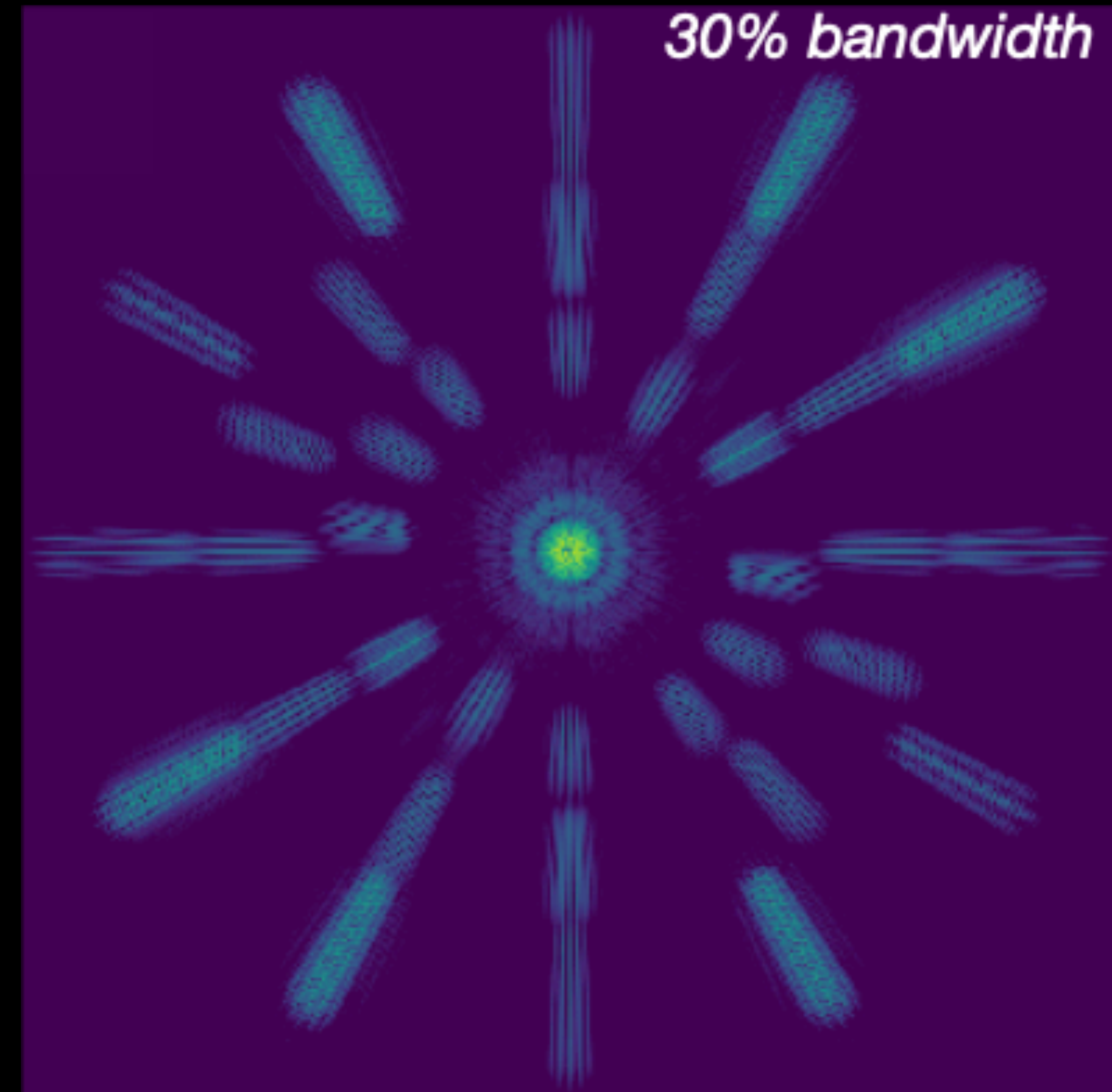
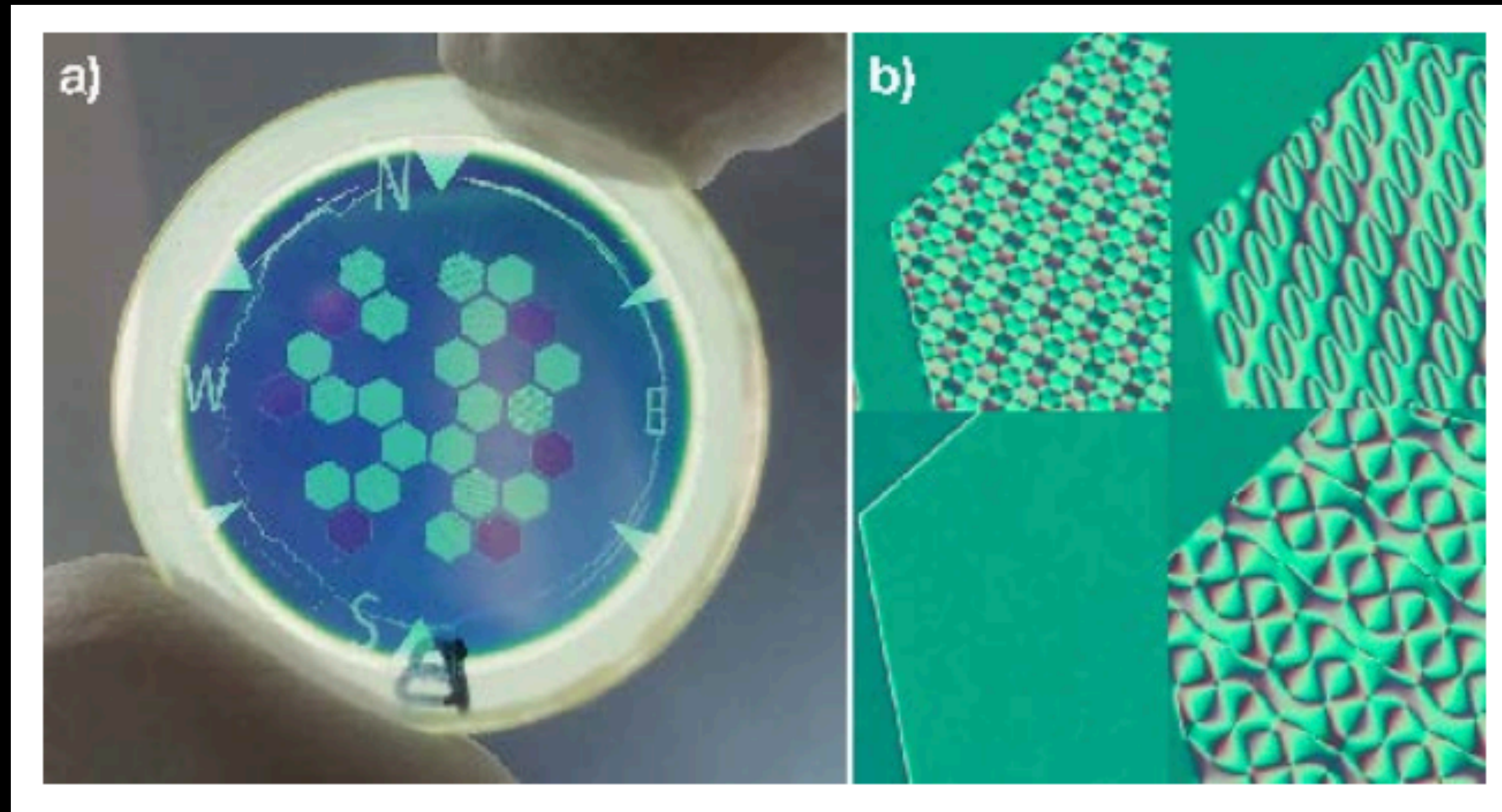
Holographic Aperture Masking (HAM)

Adaptive Optics + Interferometry + Spectral Resolution

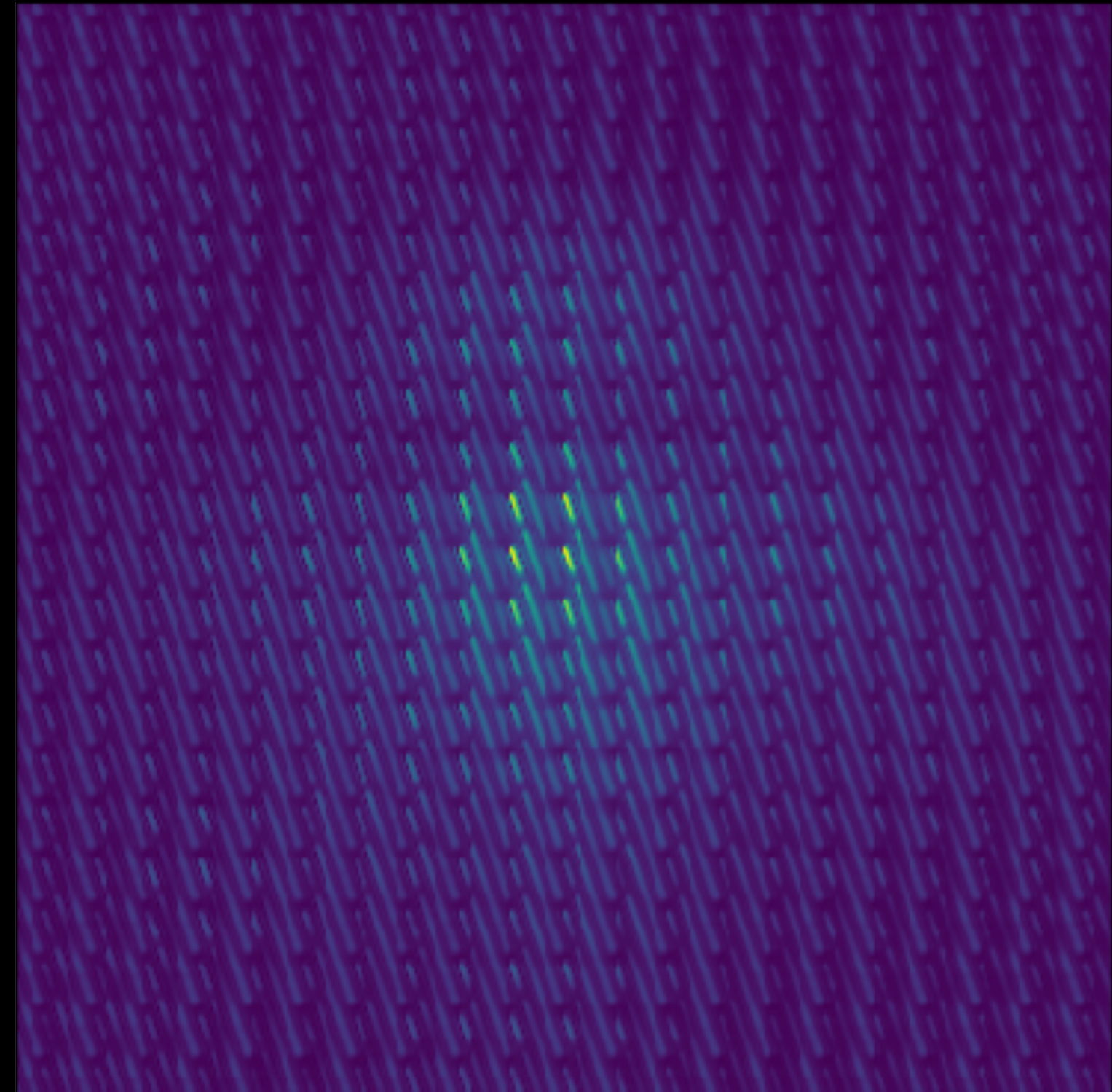
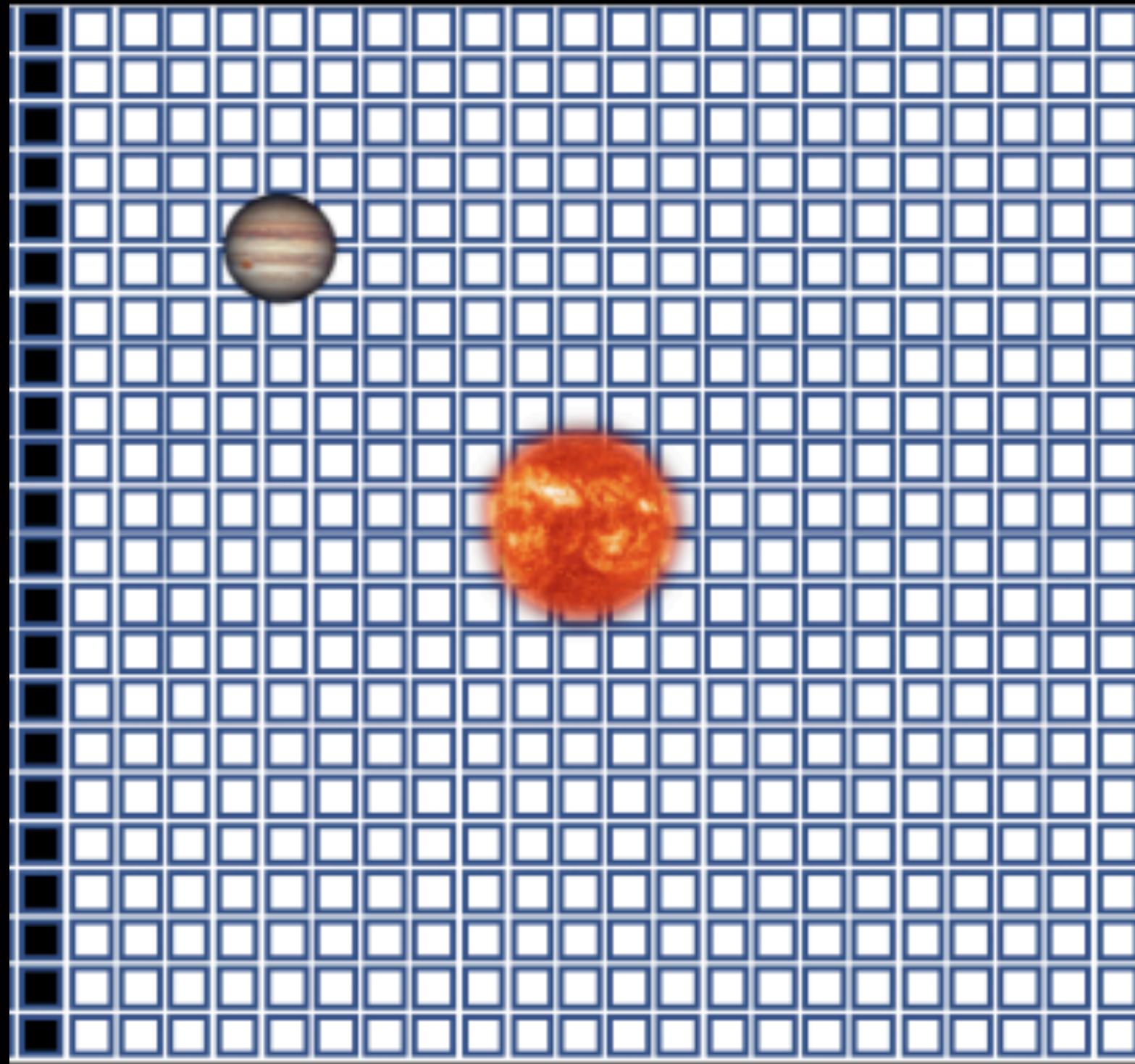


Holographic Aperture Masking (HAM)

Adaptive Optics + Interferometry + Spectral Resolution



Santa Cruz Array of Lenslets for Exoplanet Spectroscopy



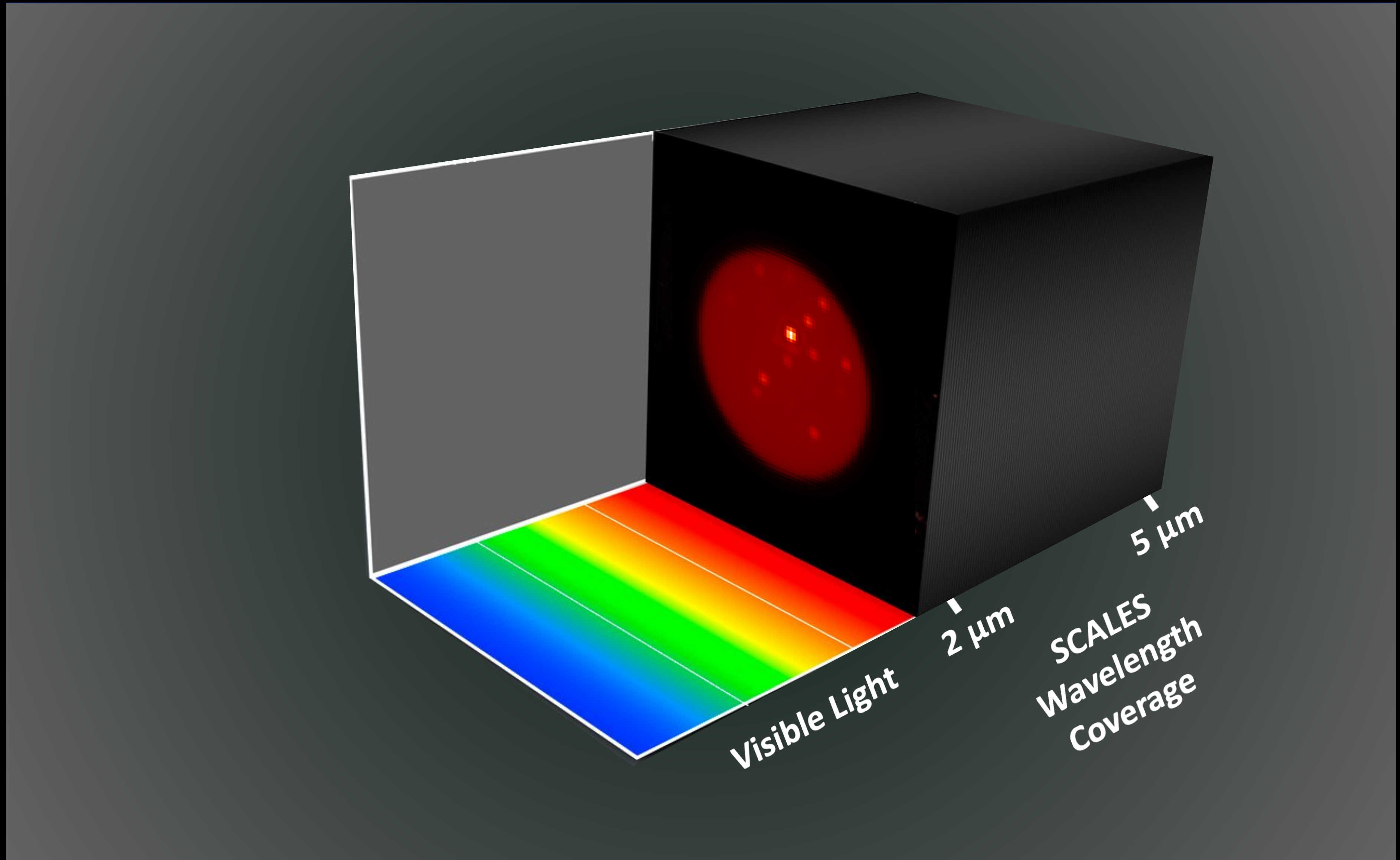
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Instrument Team: Andy Skemer (PI), Deno Stelter (IS), Nick MacDonald (PM), Dimitri Mawet, Mike Fitzgerald, Steph Sallum, Marc Kassis, Phil Hinz, Reni Kupke, Chris Ratliffe, Becky Jensen-Clem, Tim Brandt, Olivier Absil, Itsuki Sakon, Zack Briesemeister, Emily Martin, Brittany Miles, Evan Morris

Science Team: Steph Sallum (PS), Andy Skemer, Natalie Batalha, Natasha Batalha, Geoff Blake, Tim Brandt, Zack Briesemeister, Josh Eisner, Wen-fai Fong, Thomas Greathouse, Tom Greene, Mitsuhiro Honda, Charles Kilpatrick, Katherine de Kleer, Mike Liu, Dimitri Mawet, Brittany Miles, Caroline Morley, Imke de Pater, Diana Powell, Justin Spilker, Kevin Wagner, Yifan Zhou

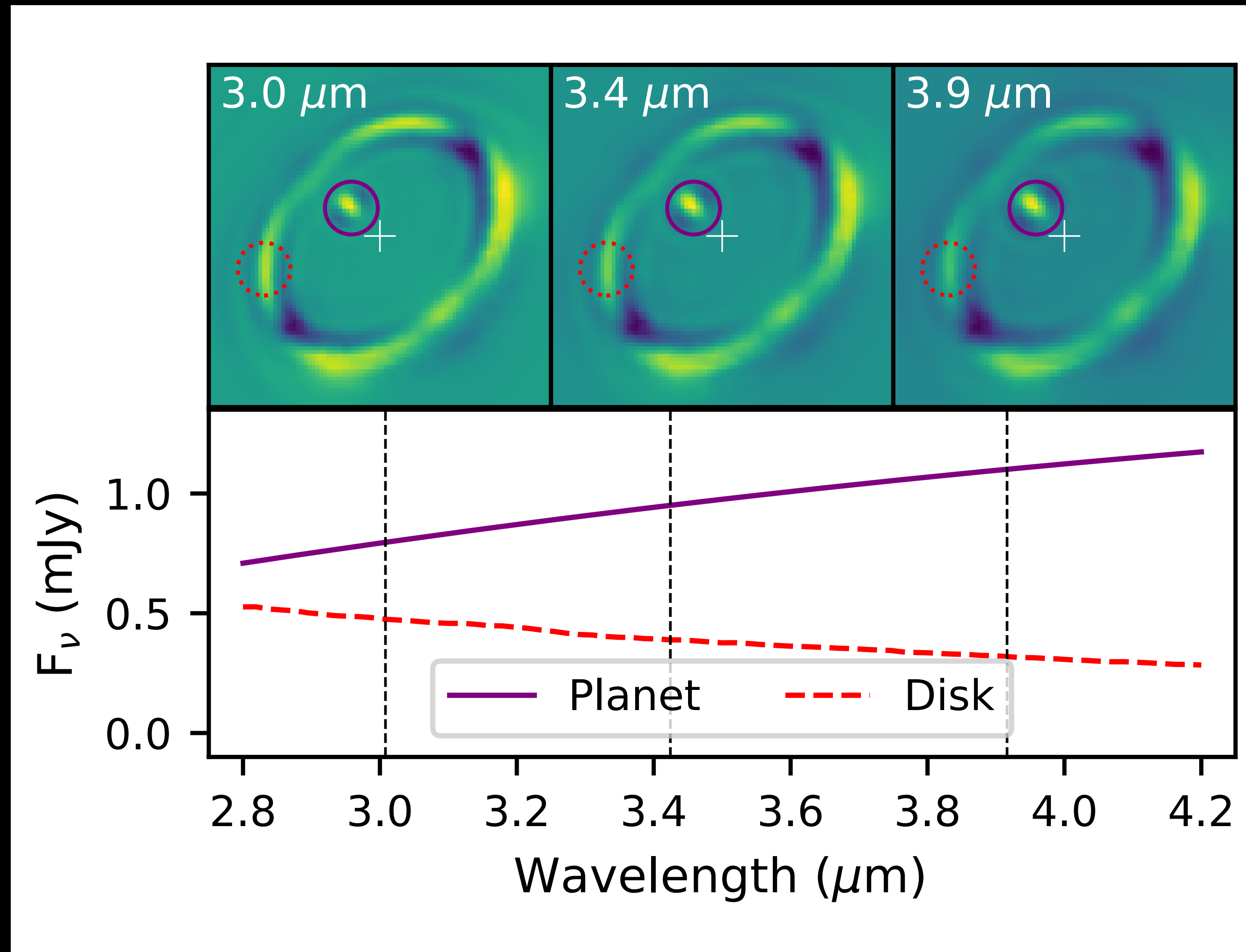
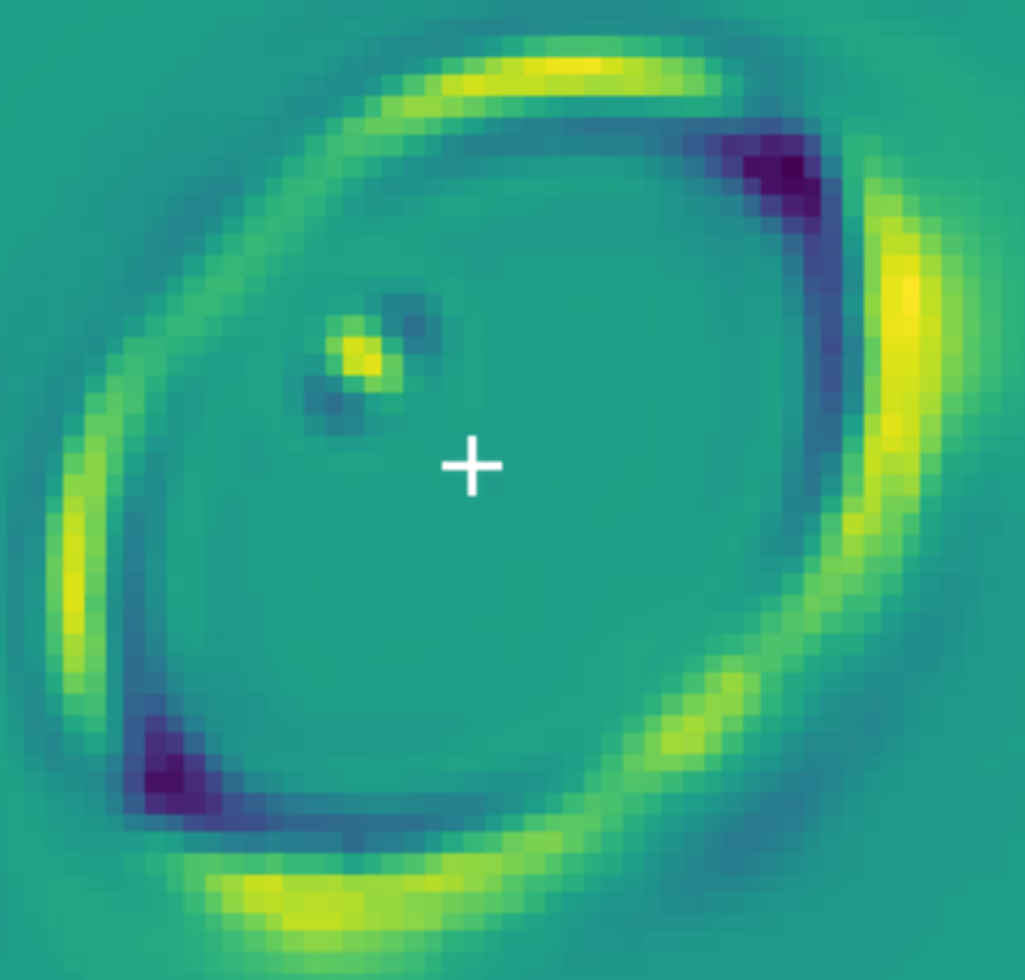
Santa Cruz Array of Lenslets for Exoplanet Spectroscopy

SCALES



SCALES: Planet Formation Science

$\lambda = 2.8$



Looking Ahead

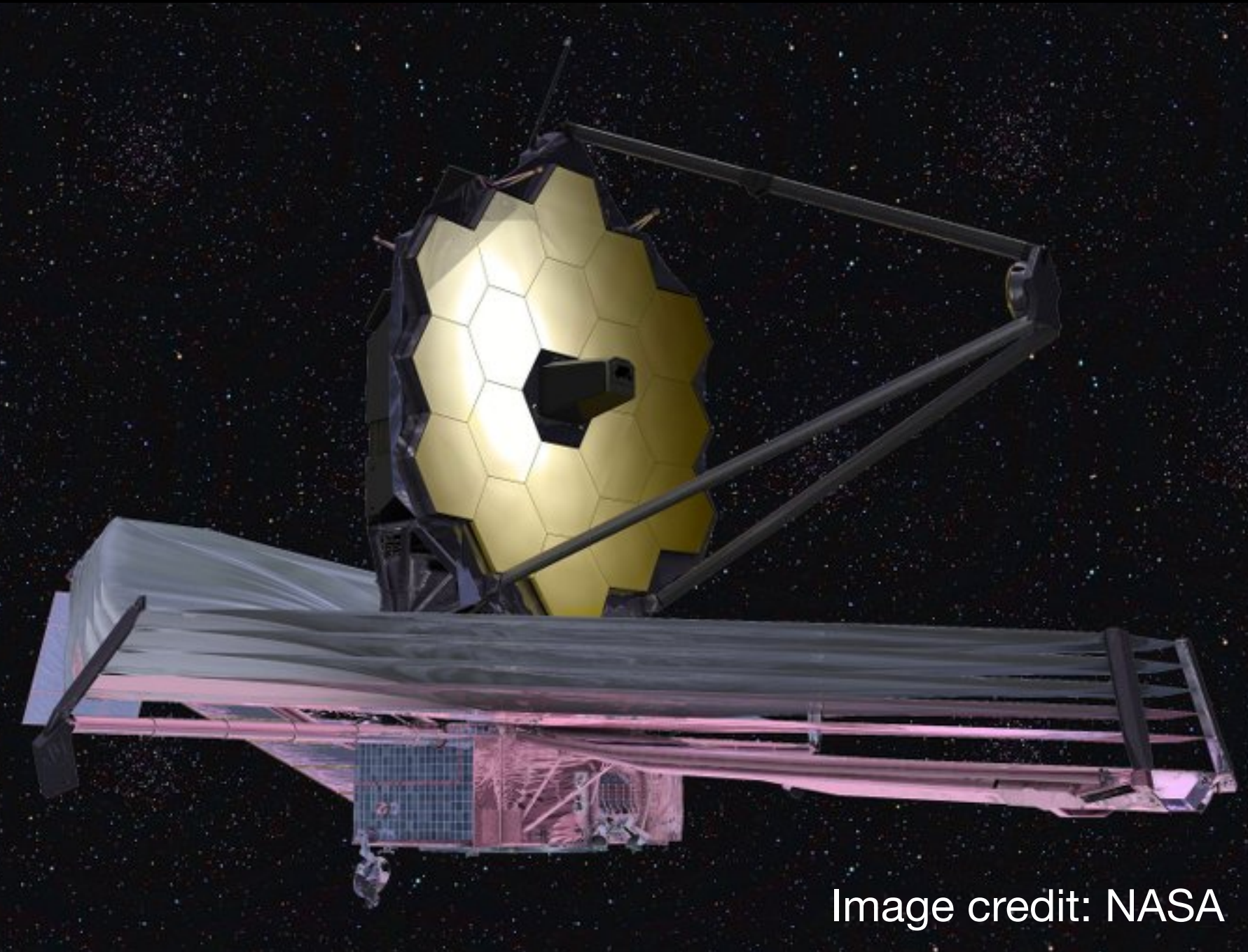


Image credit: NASA

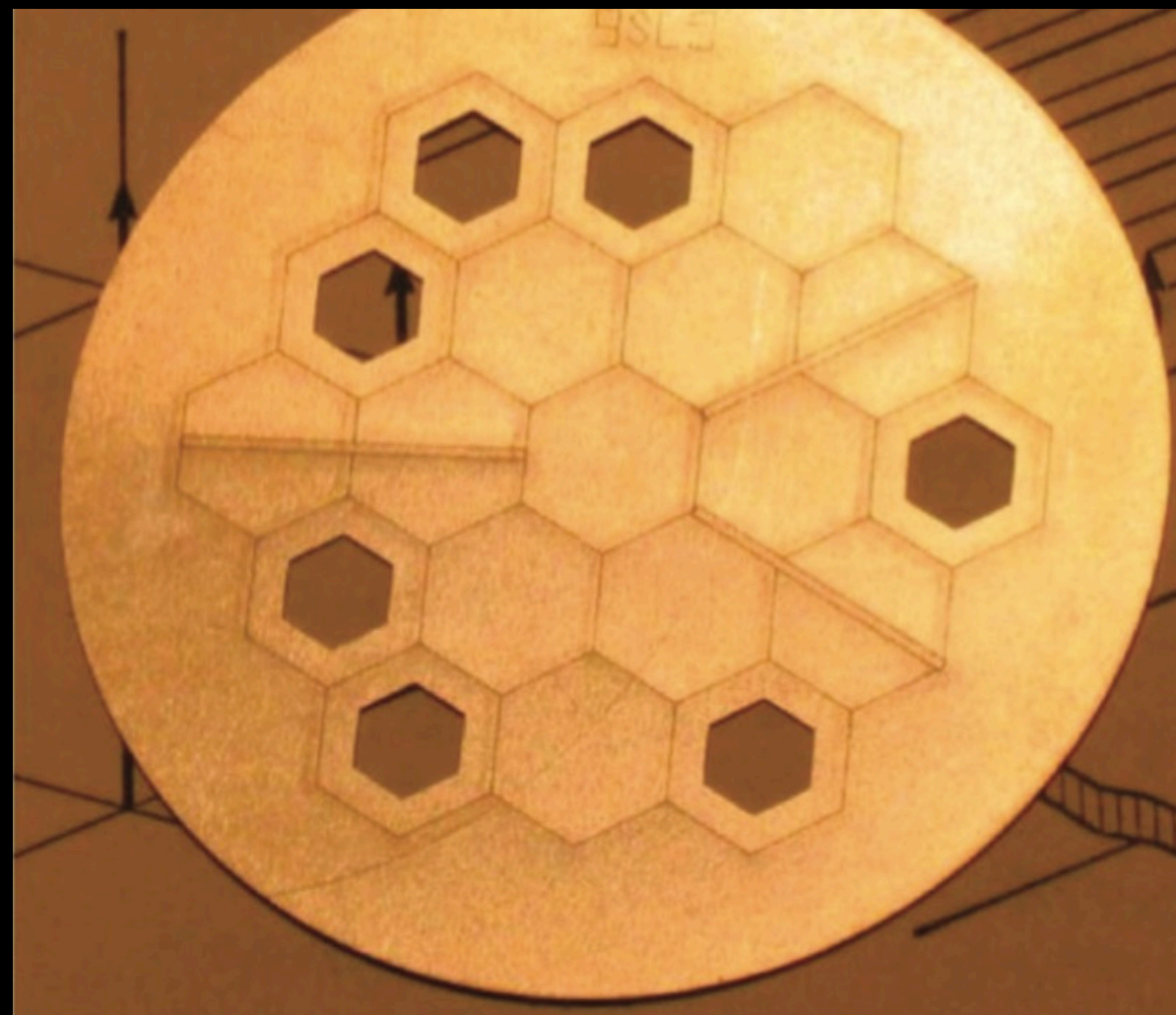


Image credit: TMTIO

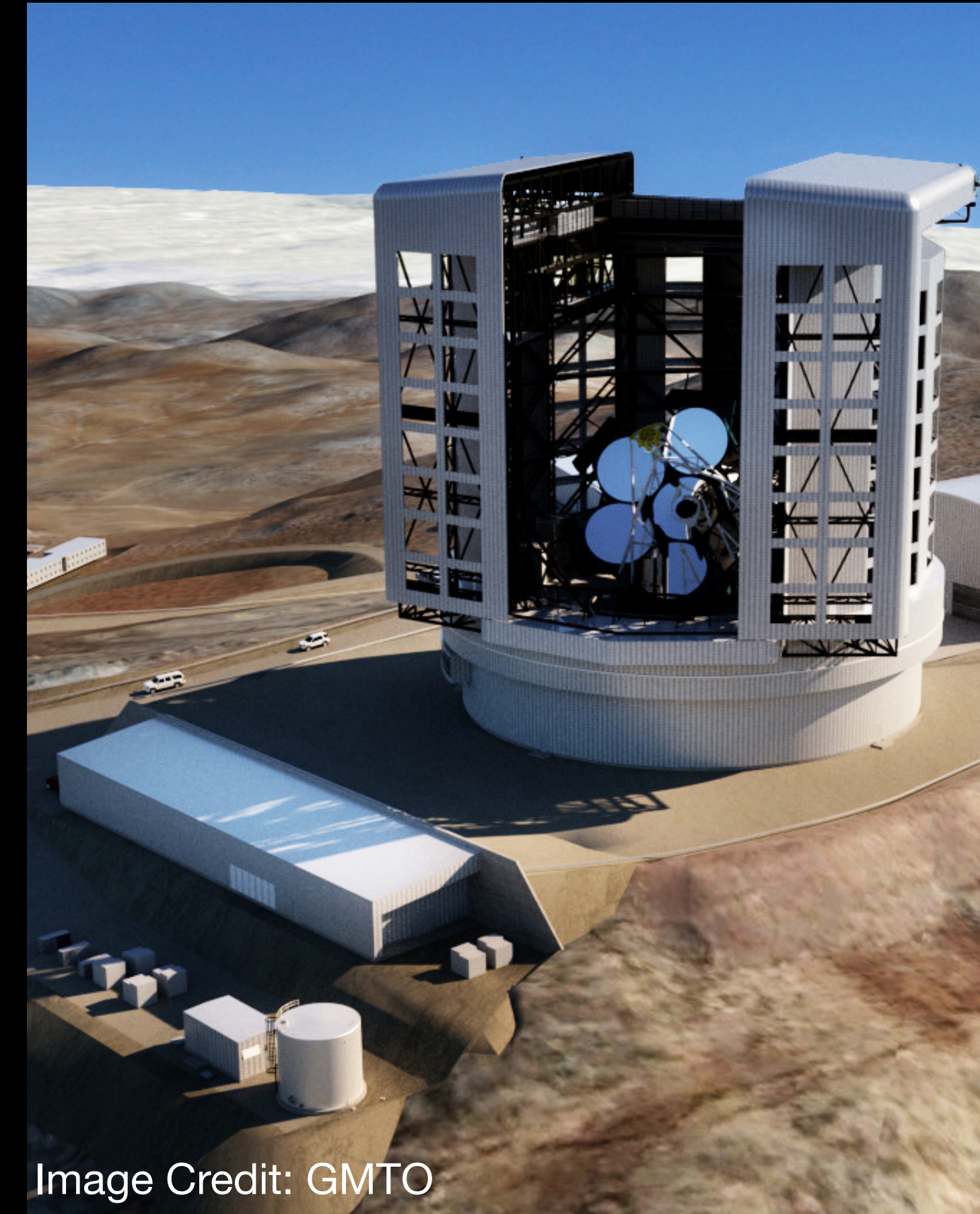


Image Credit: GMTO

UCI School of Physical Sciences

Imaging Planet Formation with Professor Steph Sallum

Welcome, we will begin shortly

*For questions, please utilize the Q&A feature at the bottom of
your screen*

Text PSBLS to 41444 to give!